

The Mu2e Tracker and Calorimeter Systems

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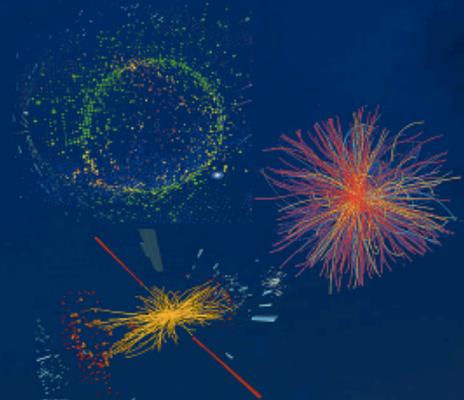
on behalf of the Mu2e Collaboration



EUROPEAN PHYSICAL SOCIETY
CONFERENCE ON HIGH ENERGY PHYSICS

5-12 July 2017 – Lido di Venezia, Italy

- ★ Astroparticle Physics and Cosmology
- ★ Neutrinos and Dark Matter
- ★ Flavour and CP Violation
- ★ Standard Model and Beyond
- ★ Electroweak Symmetry Breaking
- ★ Quantum Field and String Theory
- ★ QCD and Heavy Ions
- ★ Accelerators and Detectors
- ★ Outreach, Education, and Diversity



Charge Lepton Flavour Violation

CLFV strongly suppressed in Standard Model: $BR \leq 10^{-50} \Rightarrow$ **its observation indicates New Physics**

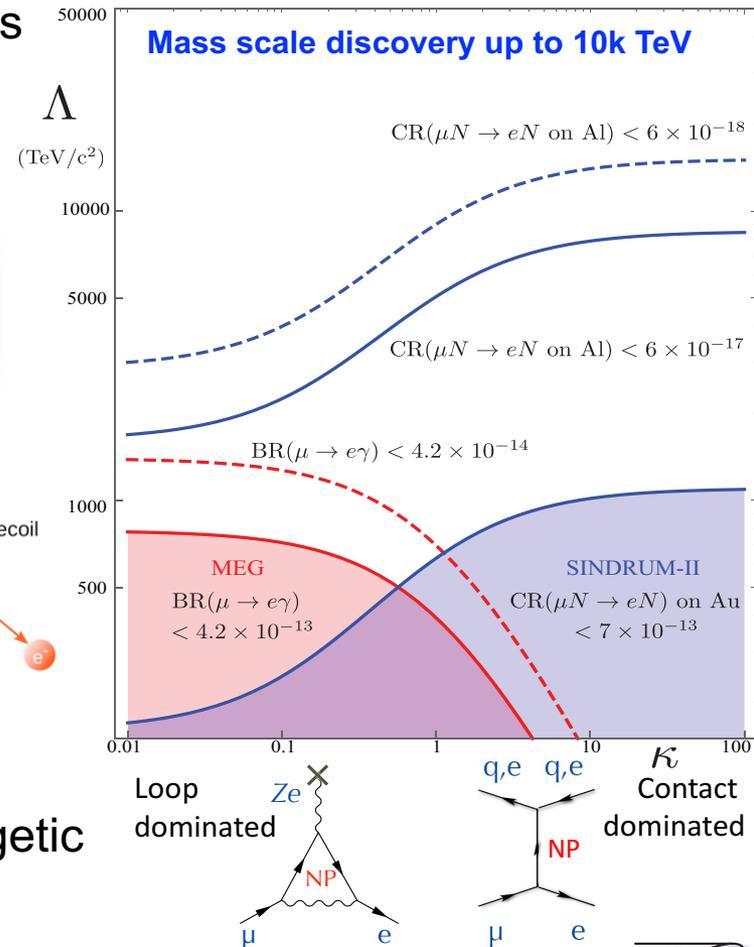
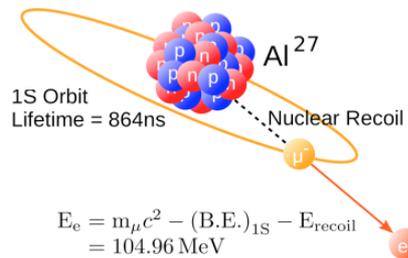
CLFV@Mu2e: coherent neutrinoless conversion of a muon to an electron in the field of a nucleus \Rightarrow discovery sensitivity on many NP models

Goal: **10^4 improvement w.r.t. previous conversion experiment** (SINDRUM II)

$$R_{\mu e} = \frac{\Gamma(\mu^- + N(A, Z) \rightarrow e^- + N(A, Z))}{\Gamma(\mu^- + N(A, Z) \rightarrow \text{all muon capture})} \leq 6 \times 10^{-17} \text{ (@90\%CL)}$$

Experimental technique:

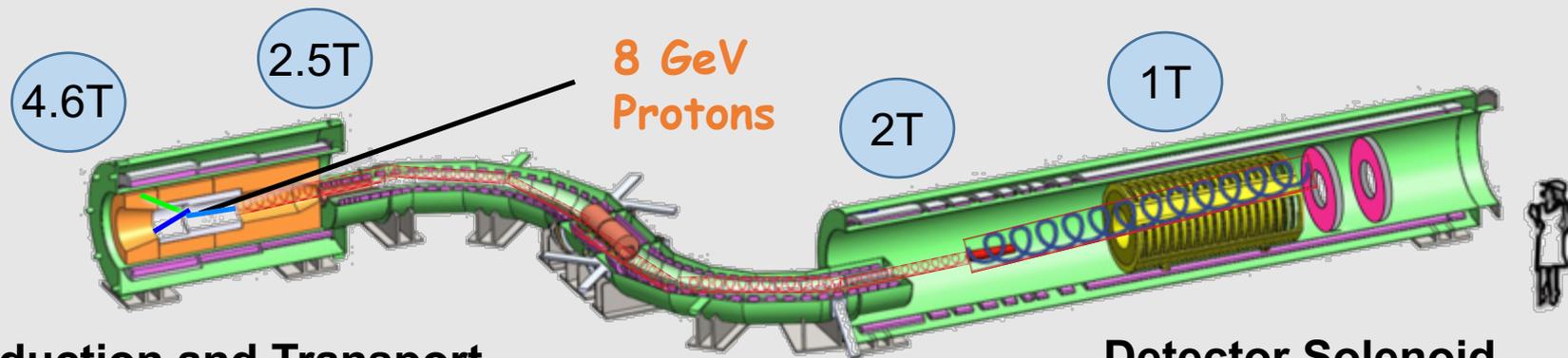
- ✗ Beam of low momentum muons
- ✗ Muons stopped in Al target
- ✗ Muons trapped in orbit around the nucleus
- ✗ Look for $\mu^- N(A, Z) \rightarrow e^- N(A, Z)$ events: mono-energetic e^- with $E \sim M_\mu$, produced with $\tau_\mu^{\text{Al}} = 864 \text{ ns}$



The Mu2e Experiment

SES @ 2.4×10^{-17} requires demanding detector technologies:

- 10^{18} μ stopped
 - 10^{20} p on target
 - $N_{\text{bckg}} < 0.5$
- in 3 years running



Production and Transport Solenoids

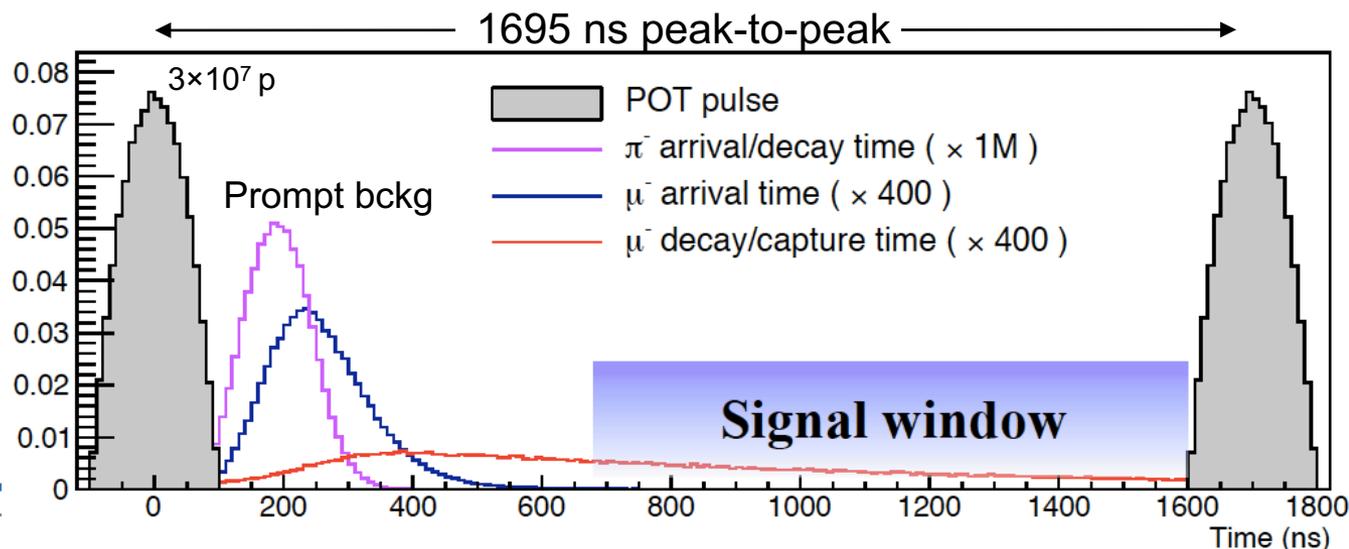
Production, selection and transport of low momentum muon beam

Detector Solenoid

- Muon capture on Al target
- Tracker, EM Calorimeter
- Outside: Cosmic Ray Veto

Bunch structure:

- Pulsed proton beam and a delayed live gate to suppress prompt backgrounds
- Narrow proton pulses
- Out-of-time protons suppressed by $O(10^{10})$

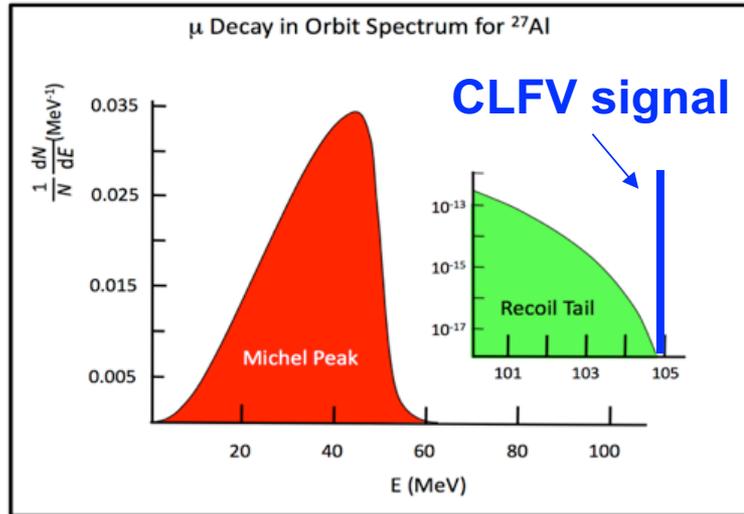


The Tracker

One of two main bckg:

Decay In Orbit

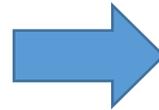
$\mu \rightarrow e \nu_{\mu} \nu_e$
(~39% on Al)



- ✗ Nuclear modifications push DIO spectrum near conversion electron
- ✗ DIO and CLFV signal, **Conversion Electron (CE)**, overlap after energy loss and detector resolution

Detector requirements:

1. Small amount of X_0
2. $\sigma_p < 180 \text{ keV @ } 105 \text{ MeV}$
3. Good rate capability:
 - 20 kHz/cm² in live window
 - Beam flash of 3 MHz/cm²
4. dE/dx capability to distinguish e^-/p
5. Operate in $B = 1 \text{ T}$, 10^{-4} Torr vacuum
6. Maximize/minimize acceptance for CE/DIO



Low mass straw drift tubes design:

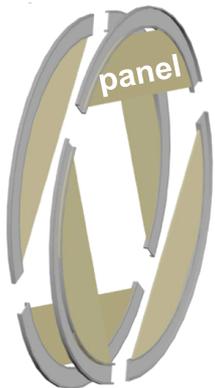
- 5 mm diameter, 33 – 117 cm length
- 15 μm Mylar wall, 25 μm Au-plated W wire
- 80:20 Ar:CO₂ @ 1 atm
- Dual-ended readout



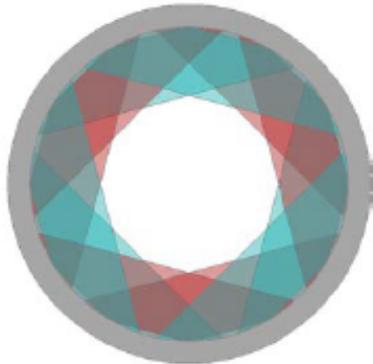
Tracker Design

- ✗ > 20,000 tubes arranged in planes on stations
- ✗ Self-supporting panel consists of 2×48 straws, two staggered layers
- ✗ 6 panels assembled into a plane, 2 planes assembled into a station, 18 stations
- ✗ Rotation of panels and planes for stereo reconstruction

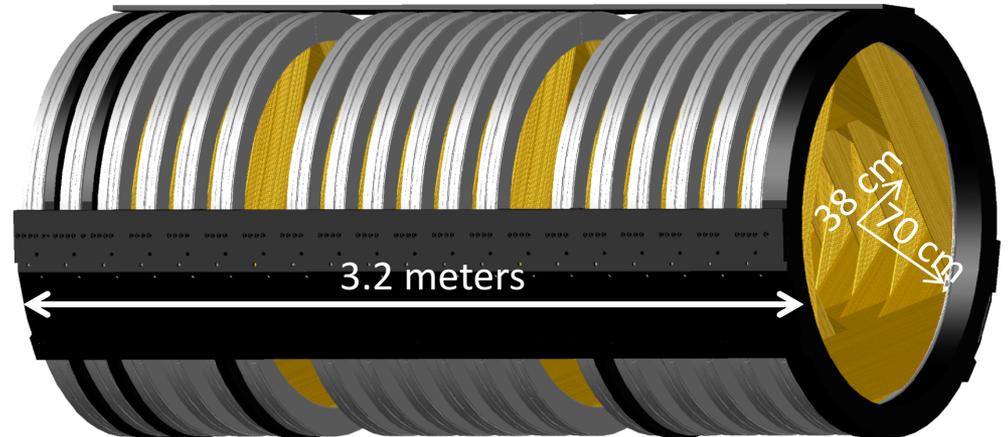
Tracker Plane



Tracker Station:
2 rotated planes



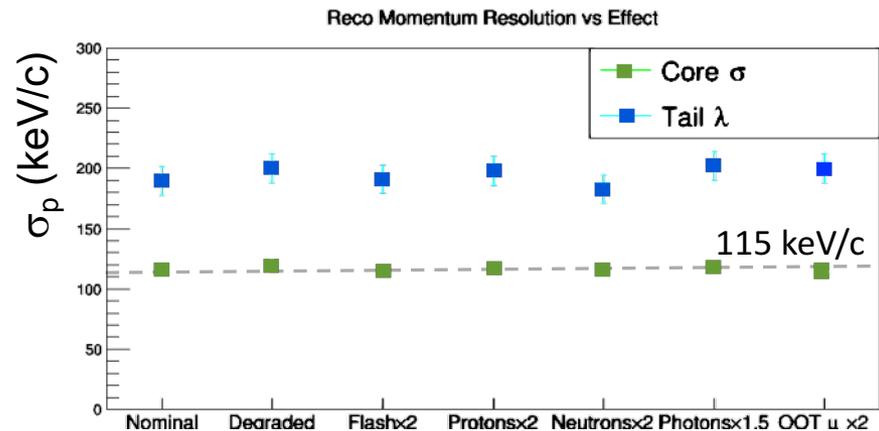
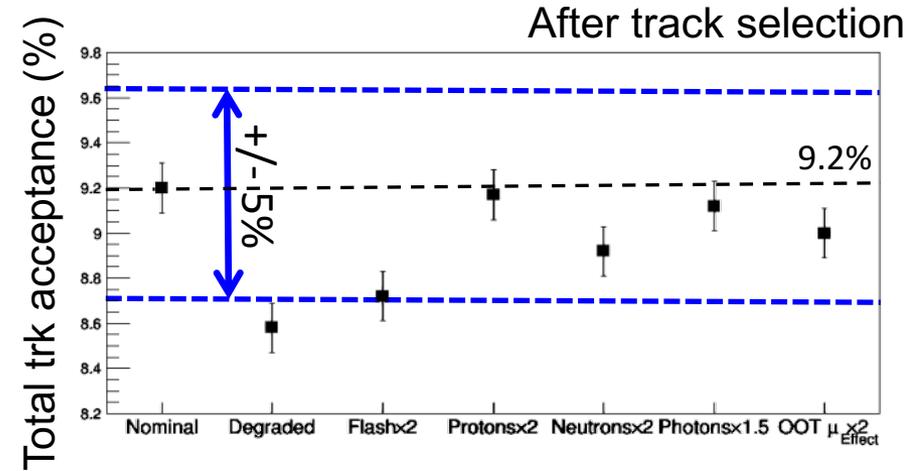
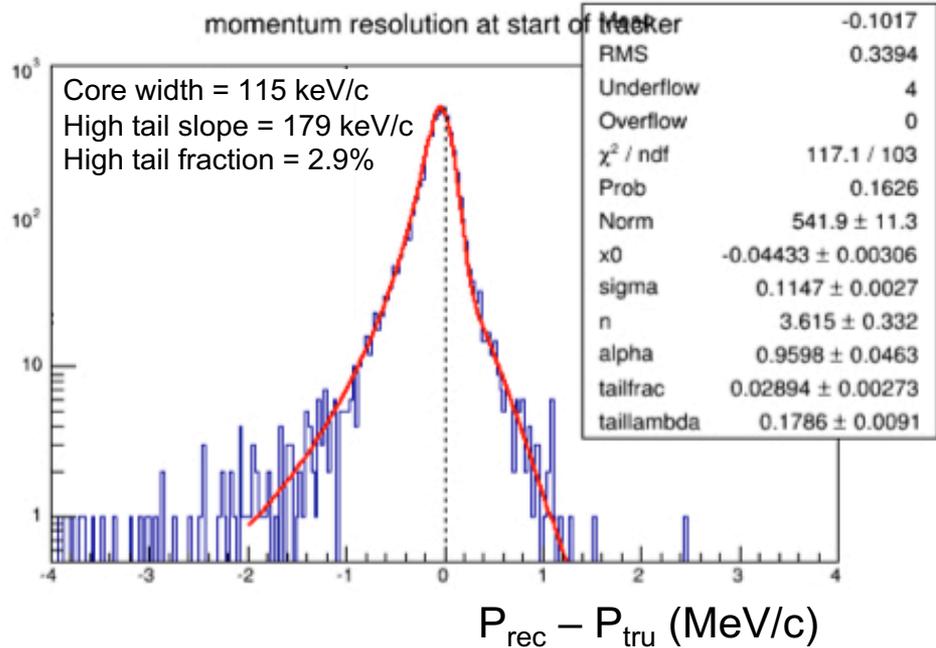
Tracker: 18 stations



- ✗ Inner 38 cm is purposely un-instrumented
 - Blind to beam flash (low momentum particles)
 - Blind to >99% of DIO spectrum

Tracker Performances

Expected tracker performances from full simulation

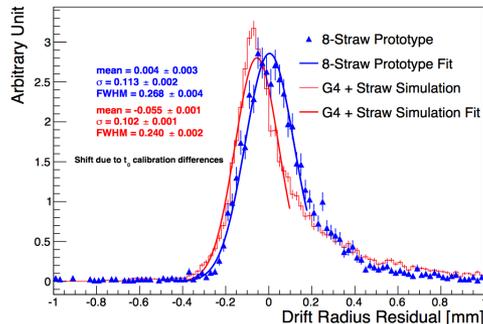
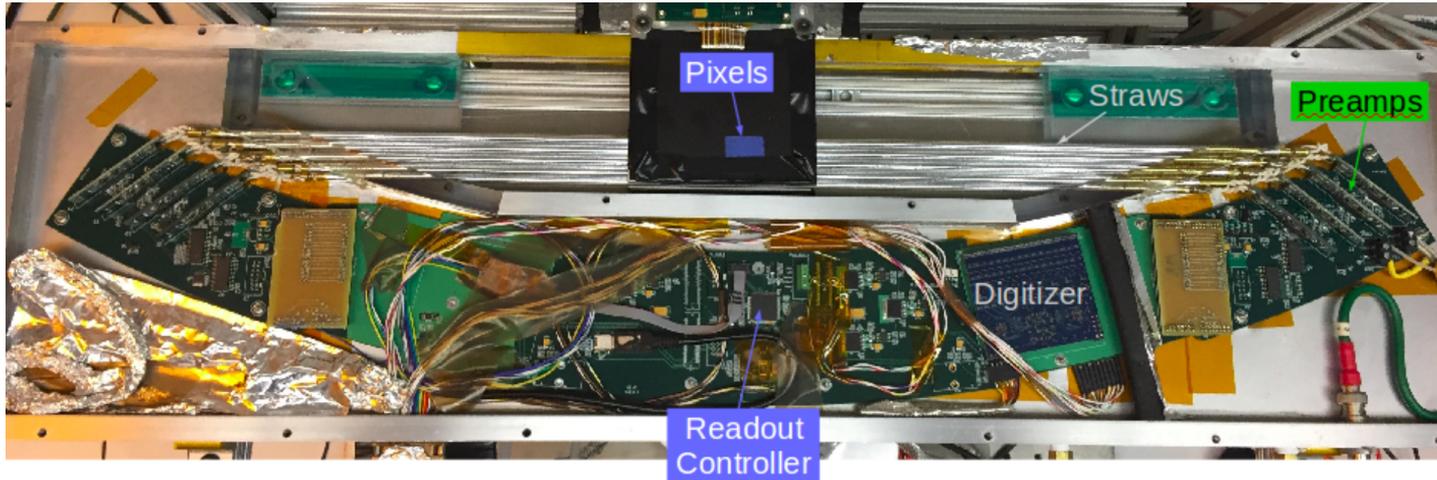


Variations in accidental hit rate

- ✗ Well within physics requirements
- ✗ Robust against increases in rate
- ✗ Inefficiency dominated by geometric acceptance

Mu2e Tracker: 8 channel prototype

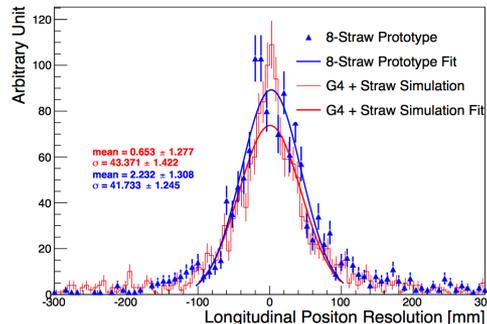
Test with cosmics to measure gain, resolution...



Transverse Resolution
(Data vs MC)

$$\sigma_{data} = 0.113 \pm 0.002 \text{ mm}$$

$$\sigma_{MC} = 0.102 \pm 0.001 \text{ mm}$$



Longitudinal Resolution
(Data vs MC)

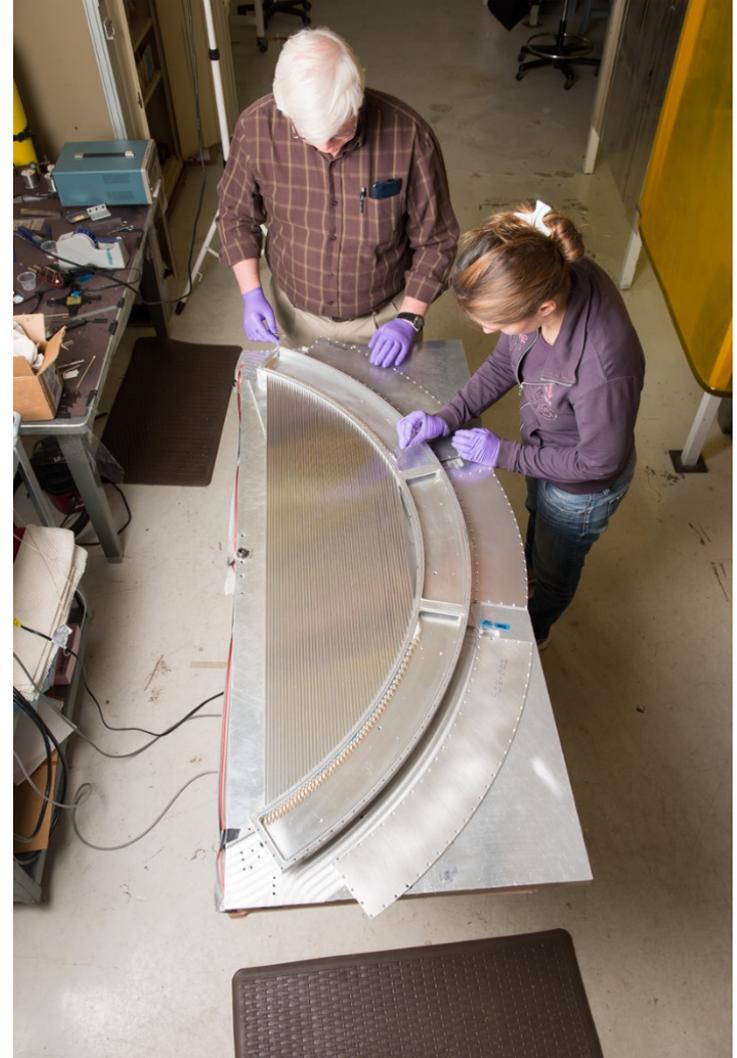
$$\sigma_{data} = 42 \pm 1 \text{ mm}$$

$$\sigma_{MC} = 43 \pm 1 \text{ mm}$$

Parameter	Value	Reference
N electrons per ionization	$\langle N \rangle = 2$	NIMA 301, 202(1991)
Energy per ionization electron	39 eV	NIST (27-100 eV) and G4
Avg. Straw Gain	70k	Prototype (PAM, ^{55}Fe)
Threshold Value	12 mV	Prototype (DVM, ^{55}Fe)
Threshold Noise	3 mV	Spice Sim. (V. Rusu)
Shaping Time	22 ns	Prototype (^{55}Fe)

Panel Prototype

- ✗ First pre-production prototype, with final design, recently built and being tested
- ✗ Orders placed for final production
- ✗ FEE prototypes tested successfully
- ✗ Vertical slice test to be performed on fully instrumented panels with entire FEE chain



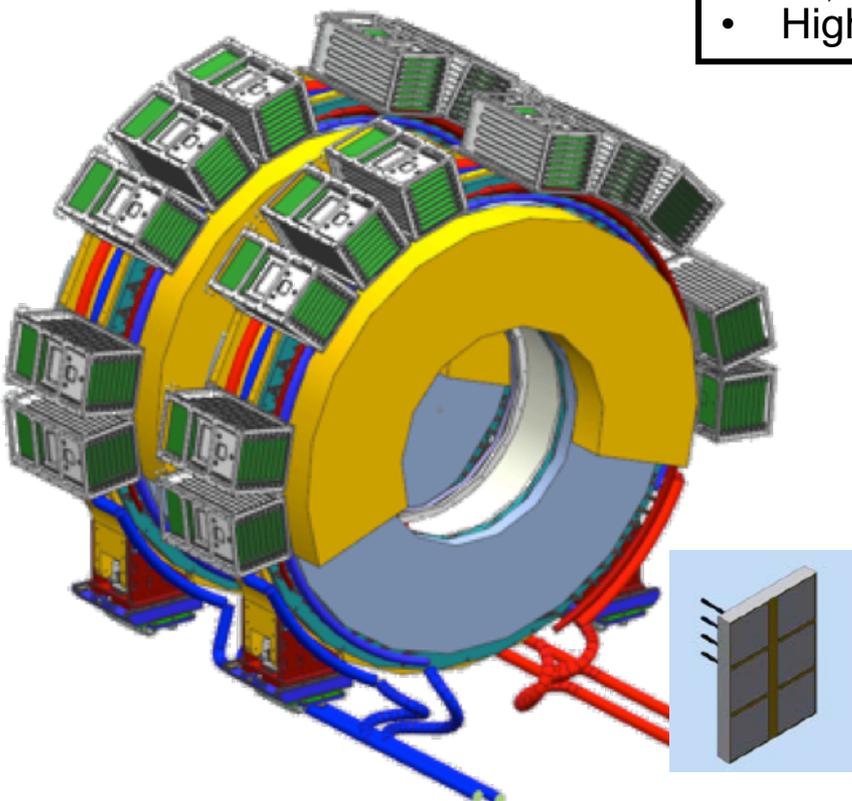
The Electromagnetic Calorimeter

Calorimeter provides confirmation for CE and other crucial functions:

- ✗ PID: e/μ separation
- ✗ EMC seeded track finder
- ✗ Standalone trigger

Requirements:

- $\sigma_E/E = \mathcal{O}(5\%)$ for CE
- $\sigma_T < 500$ ps for CE
- $\sigma_{X,Y} \leq 1$ cm
- High acceptance for CE
- Fast ($\tau < 40$ ns)
- Operate in 1T and 10^{-4} Torr
- Redundancy in readout
- Radiation hard: 90 krad photons and 3×10^{12} n/cm²



EMC Design:

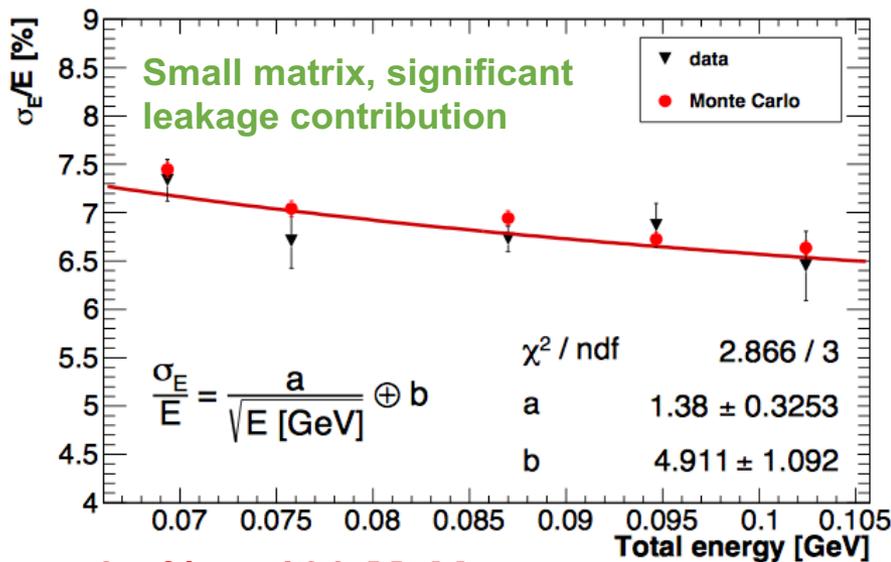
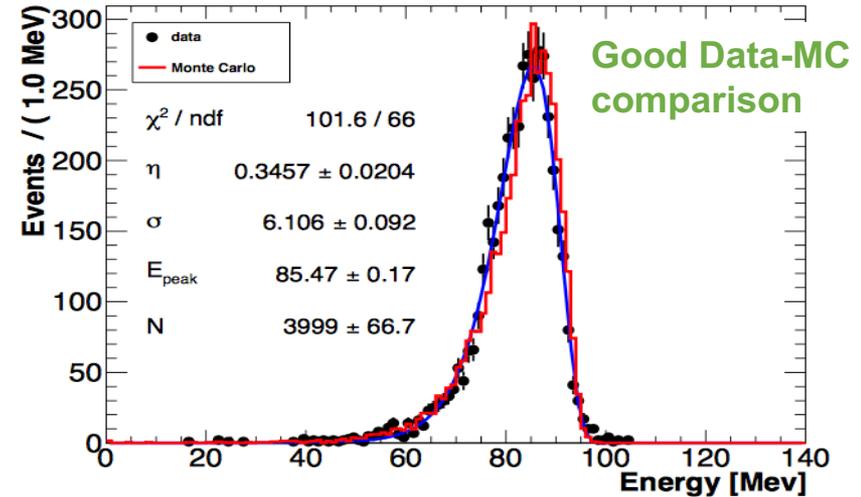
- ✗ Two disks, $R_{in}=374$ mm, $R_{out}=660$ mm, $10X_0$ length, ~ 75 cm separation
- ✗ 674+674 square x-sec **pure CsI crystals**, $(34 \times 34 \times 200)$ mm³
- ✗ For each crystal, two custom array (2×3 of 6×6 mm²) **large area UV-extended SiPMs**
- ✗ Analog FEE directly mounted on SiPM
- ✗ Calibration/Monitoring with 6 MeV radioactive source and a laser system

Disks spaced by $\frac{1}{2} \lambda$ of the helix (min-max distance from axis) for CE tracks

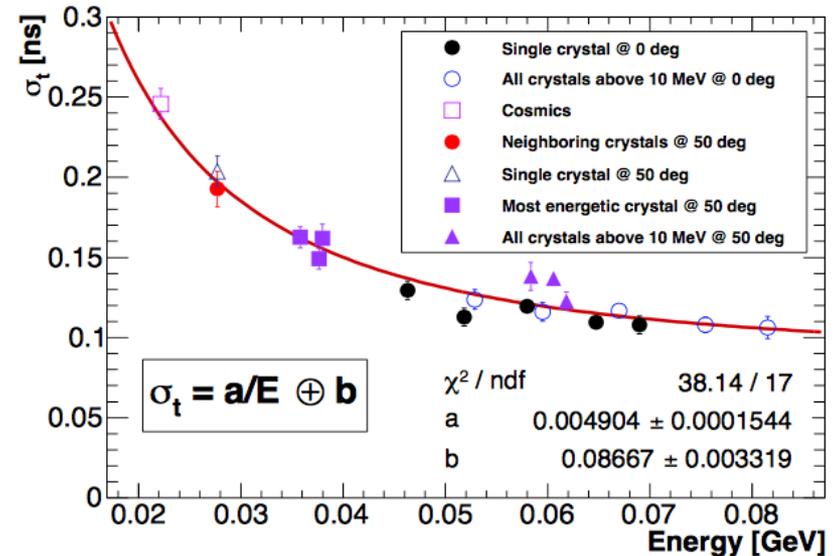
Calorimeter Performances

JINST 12 (2017) P05007

- ✗ Small prototype tested @ BTF (Frascati) in April 2015, 80–120 MeV e^-
- ✗ 3×3 array of (30×30×200) mm² undoped CsI crystals coupled to Hamamatsu MPPC
- ✗ DAQ readout: 250 Msps CAEN V1720 Wave Form Digitizer



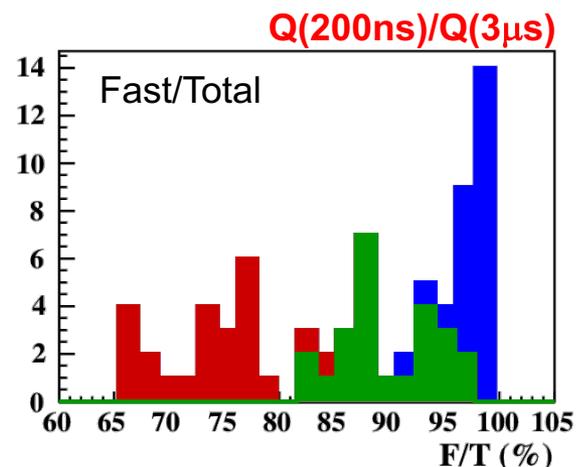
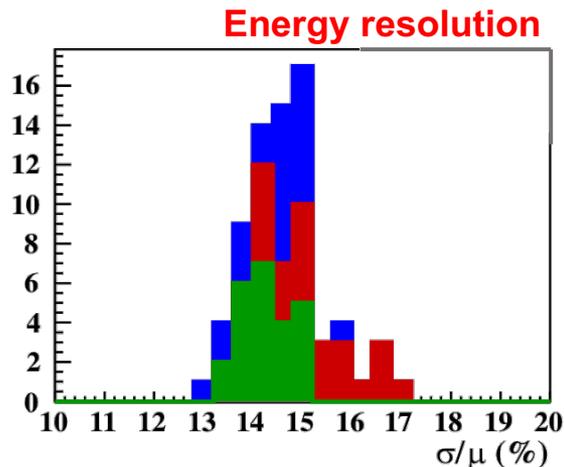
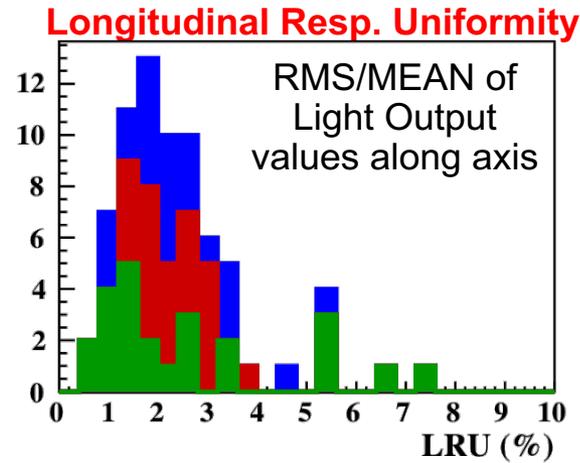
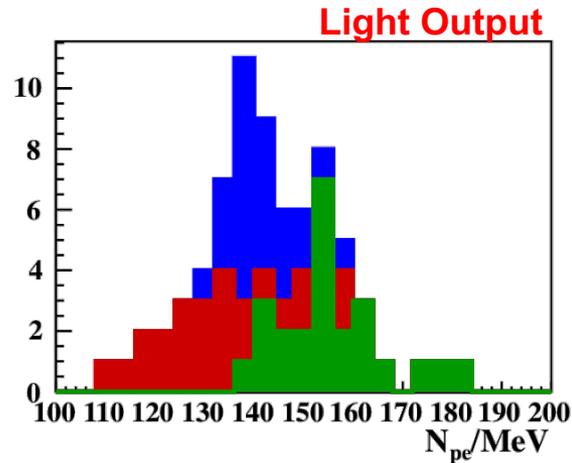
$\sigma_E \sim 6.5\%$ at 100 MeV



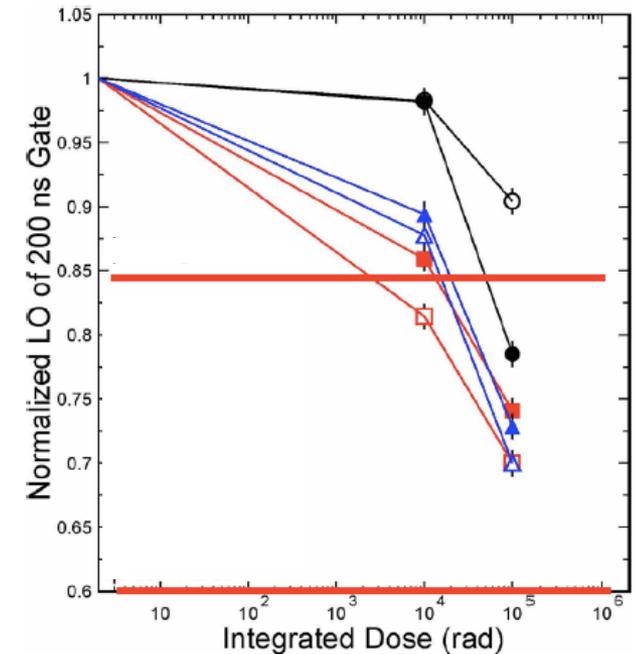
$\sigma_T \sim 110$ ps at 100 MeV

Test of pre-production crystals

- ✗ 3×24 pre-production crystals from three different vendors
- ✗ Optical properties tested with 511 keV γ 's along the crystal axis
- ✗ Crystals are wrapped with 150 μm of Tyvek and coupled to an UV-extended PMT



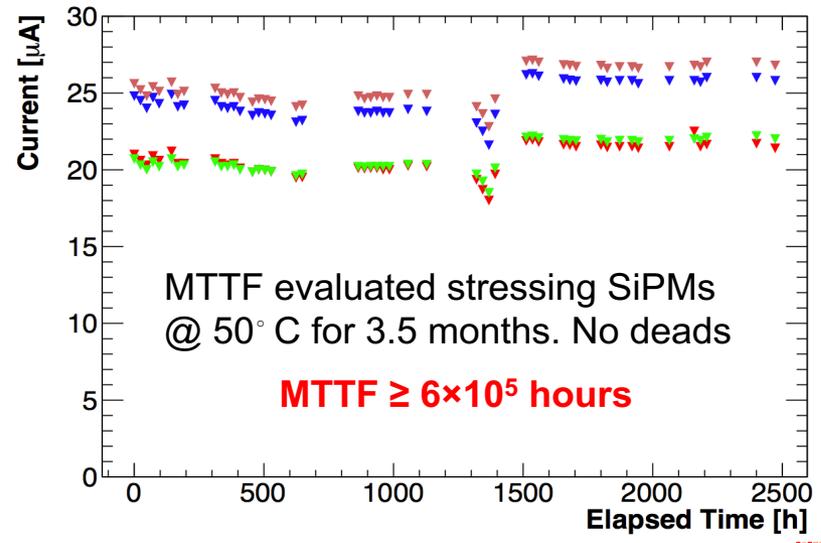
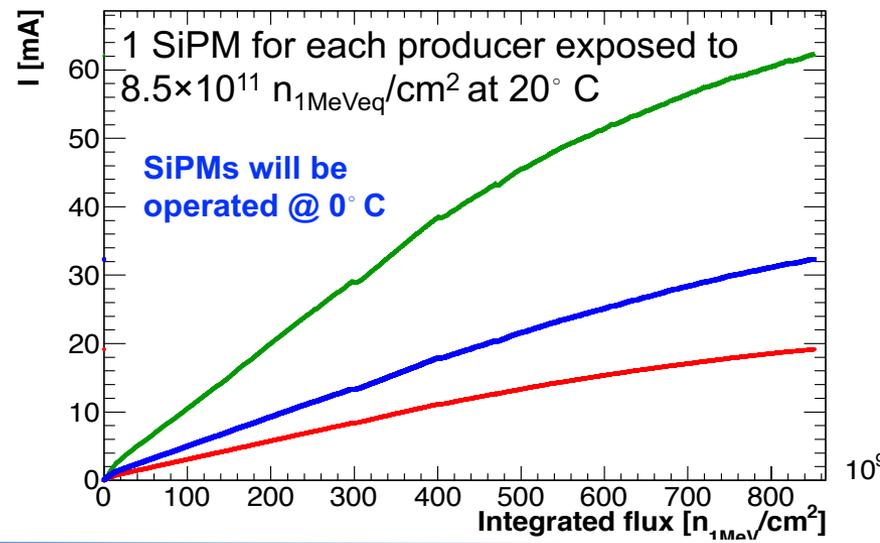
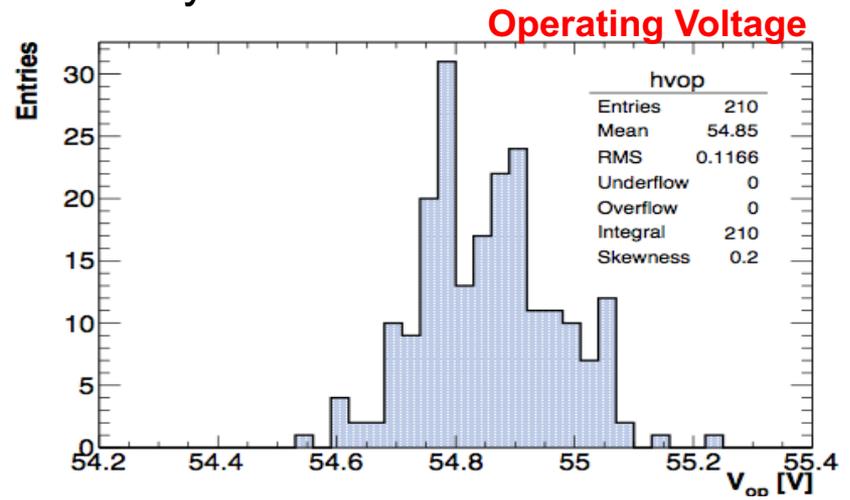
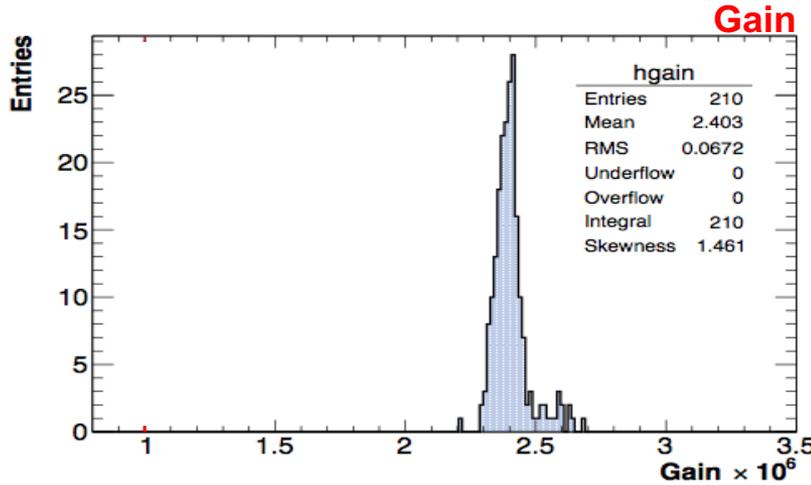
Irradiation test up to 100 krad



All satisfy Mu2e
100 krad requirement
(40% max. loss)

Test of pre-production SiPMs

- ✗ 3×50 Mu2e pre-production SiPMs from three different vendors
- ✗ 3×35 were characterized, all six cells in the array



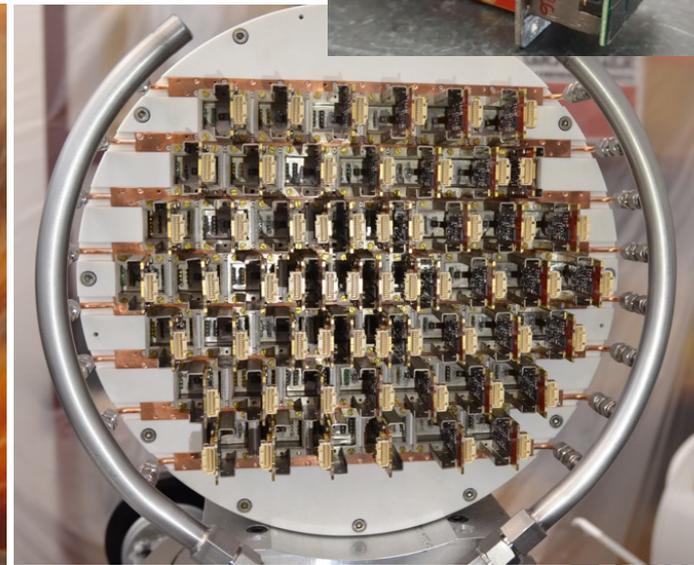
Module 0

Large EMC prototype: 51 crystals, 102 SiPMs, 102 FEE boards

Mechanics and cooling system similar to the final ones

Goals:

- ✗ Integration and assembly procedures
- ✗ Work under vacuum, low temperature, irradiation env.
- ✗ Test beam with 60–120 MeV e^- done, analysis in progress



Conclusions

- ✗ The Mu2e experiment will exploit the world's highest intensity muon beams of the Fermilab Muon Campus to search for CLFV, improving current sensitivity by a factor 10^4
- ✗ A low mass straw tube tracker and a pure CsI crystal calorimeter with SiPM readout have been selected to satisfy the demanding requirements
- ✗ Both systems are concluding the prototyping phase
- ✗ Production phase is starting, moving to full regime for end 2017
- ✗ Detector installation in 2020, followed by Mu2e commissioning and data

The Muon Campus at Fermilab



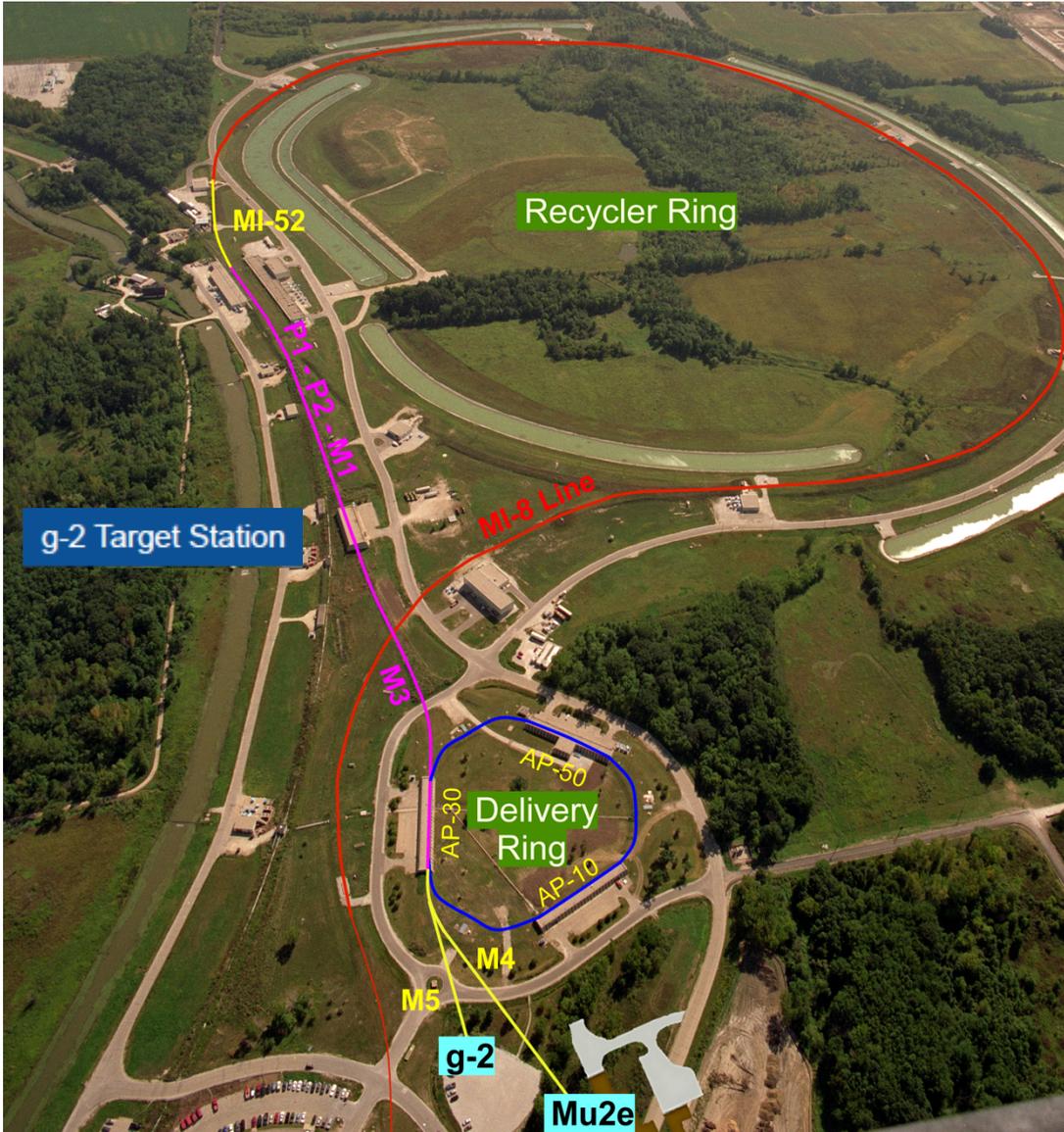
Mu2e

Muon (g-2)



**The Mu2e Collaboration
Over 200 scientists from 37 institutions**

Beam for Muon Campus



Recycler: fixed 8 GeV proton ring

Beams both to Muon Campus and neutrino experiments

Separate runs for g-2 and Mu2e

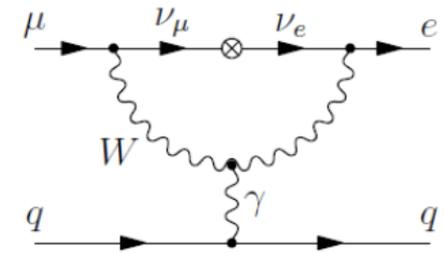
g-2: target before the delivery ring, 3.1 GeV π^+ selected, clean, polarized μ^+ beam

Mu2e: 8 GeV protons to Mu2e hall

Accelerator Readiness Review took place in March 2017

Beam commissioning for g-2 started beginning of April
First $p/\pi/\mu$ beam in the ring in June (no target and del. ring)

Standard Model
 $\mu \rightarrow e$ conversion



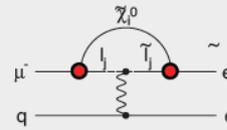
W. Altmannshofer, *et al*, arXiv:0909.1333 [hep-ph]

	AC	RVV2	AKM	δ LL	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	★★★	★	★	★	★	★★★	?
ϵ_K	★	★★★	★★★	★	★	★★	★★★
$S_{\psi\phi}$	★★★	★★★	★★★	★	★	★★★	★★★
$S_{\phi K_S}$	★★★	★★	★	★★★	★★★	★	?
$A_{CP}(B \rightarrow X_s \gamma)$	★	★	★	★★★	★★★	★	?
$A_{7,s}(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★★★	★★★	★★	?
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★	★	★	?
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★	★	★	★	★	★	★
$B_s \rightarrow \mu^+ \mu^-$	★★★	★★★	★★★	★★★	★★★	★	★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$\mu \rightarrow e \gamma$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
$\tau \rightarrow \mu \gamma$	★★★	★★★	★	★★★	★★★	★★★	★★★
$\mu + N \rightarrow e + N$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
d_n	★★★	★★★	★★★	★★	★★★	★	★★★
d_e	★★★	★★★	★★	★	★★★	★	★★★
$(g-2)_\mu$	★★★	★★★	★★	★★★	★★★	★	?

Table 8: "DNA" of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models ★★★ signals large effects, ★★ visible but small effects and ★ implies that the given model does not predict sizable effects in that observable.

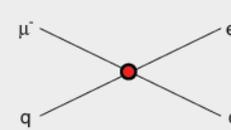
Supersymmetry

rate $\sim 10^{-15}$



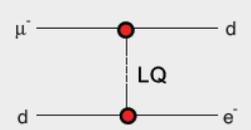
Compositeness

$\Lambda_c \sim 3000$ TeV



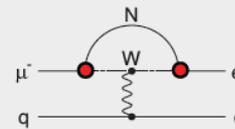
Leptoquark

$M_{LQ} = 3000 (\lambda_{\mu d} \lambda_{ed})^{1/2}$ TeV/c²



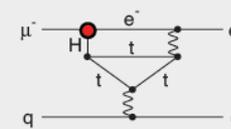
Heavy Neutrinos

$|U_{\mu N} U_{eN}|^2 \sim 8 \times 10^{-13}$



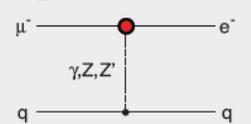
Second Higgs Doublet

$g(H_{\mu e}) \sim 10^{-4} g(H_{\mu\mu})$



Heavy Z'
 Anomal. Z Coupling

$M_{Z'} = 3000$ TeV/c²



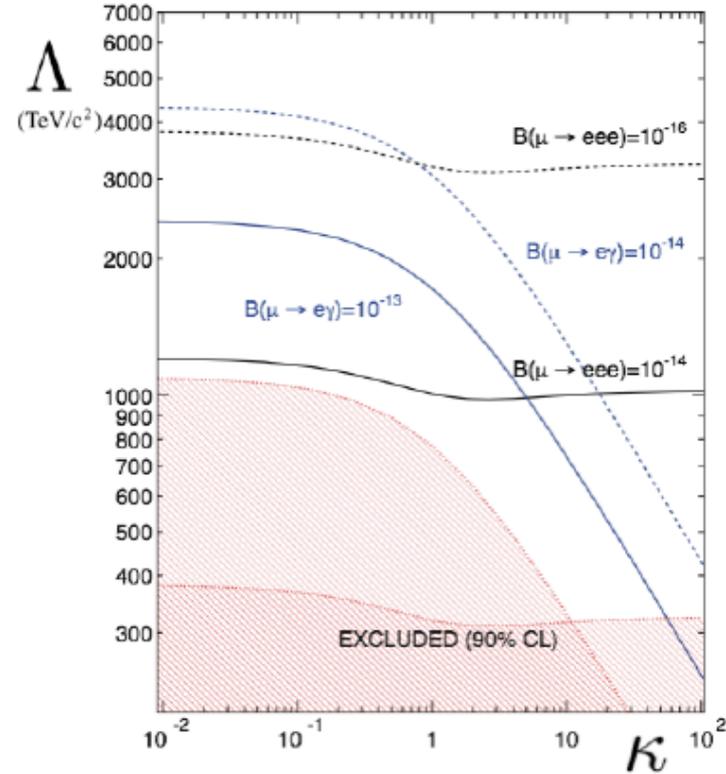
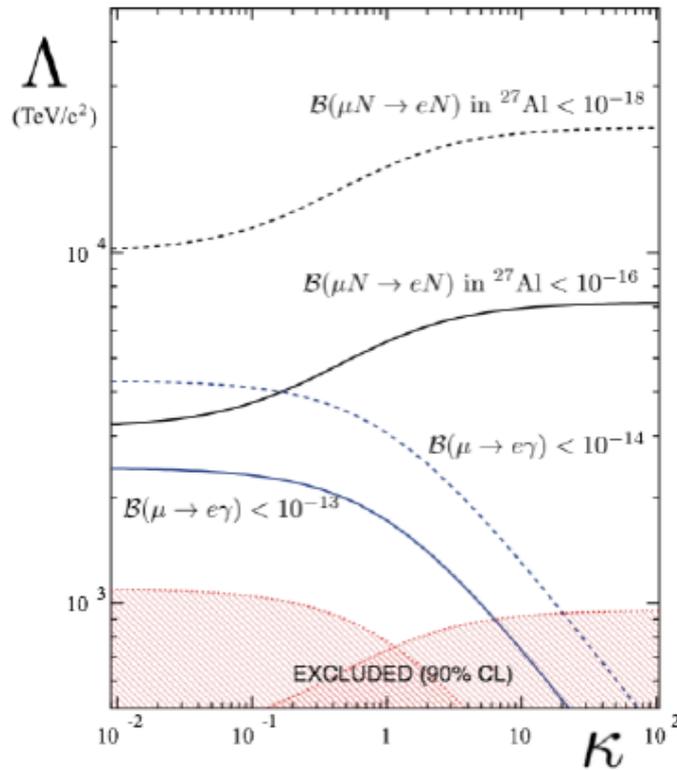
If SUSY seen at LHC \rightarrow rate $\sim 10^{-15}$

Implies **O(40)** reconstructed signal events with **negligible background** in Mu2e for many SUSY models.

Mu2e keeps discovery sensitivity for all SUSY benchmark point for LHC Phase2



Model independent Lagrangian

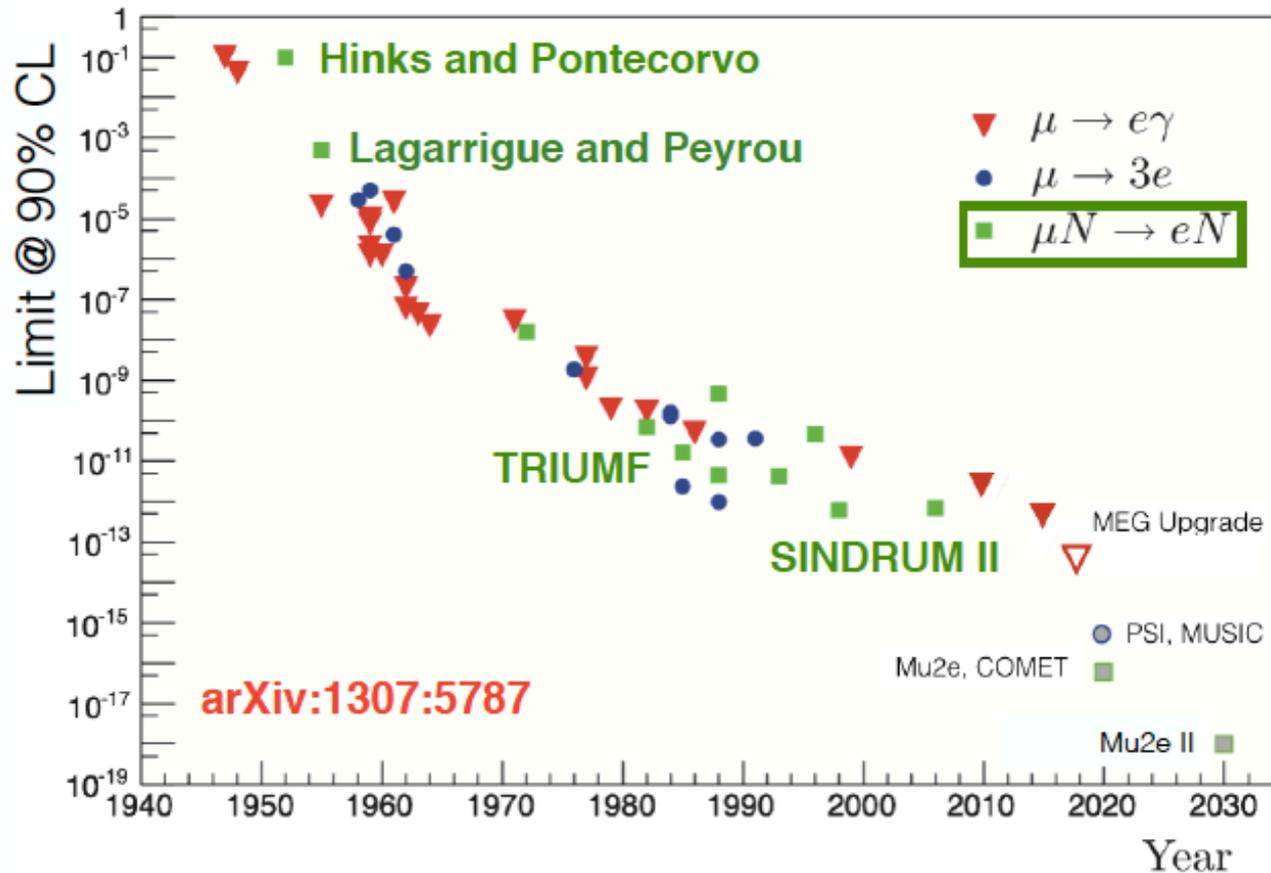


$$L_{\text{CLFV}} = \frac{m_\mu}{(\kappa + 1)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(\kappa + 1)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L (\bar{e} \gamma^\mu e)$$

“dipole term”

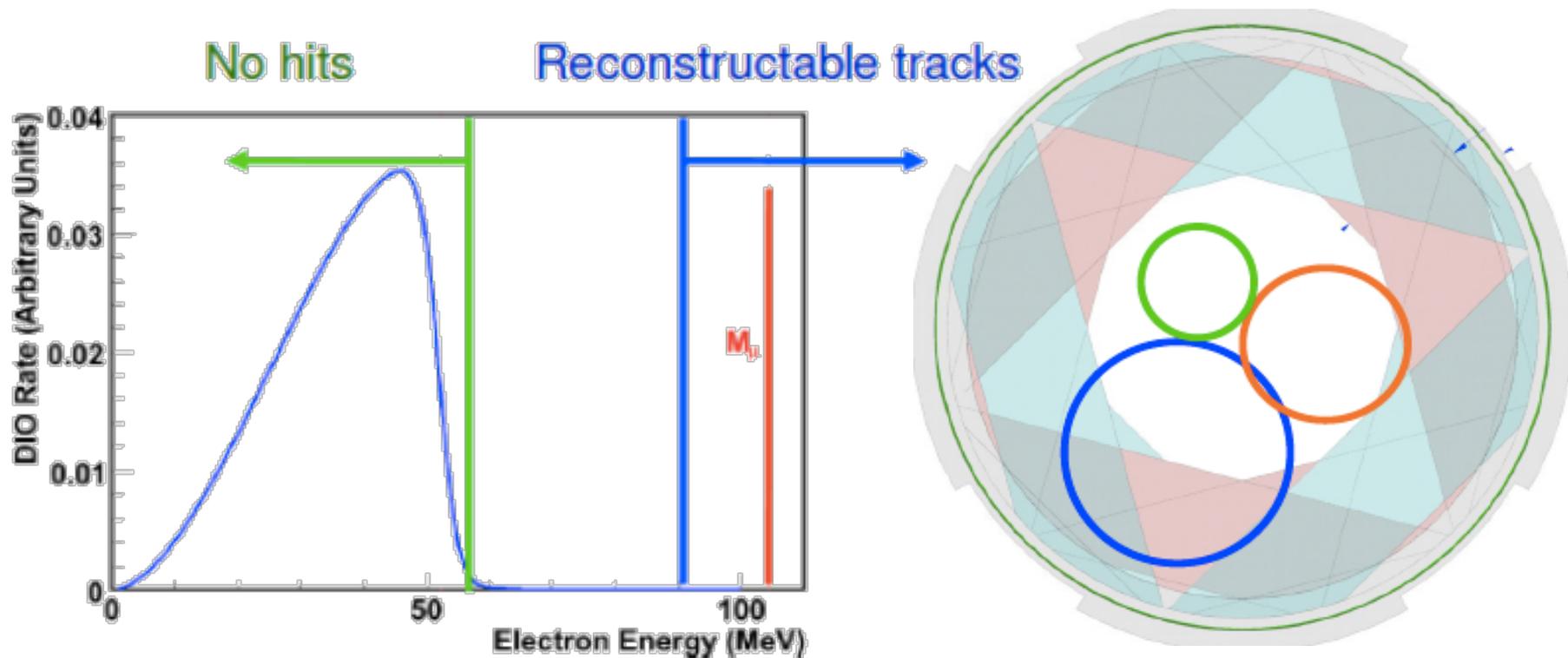
“contact term”

CLFV along the years



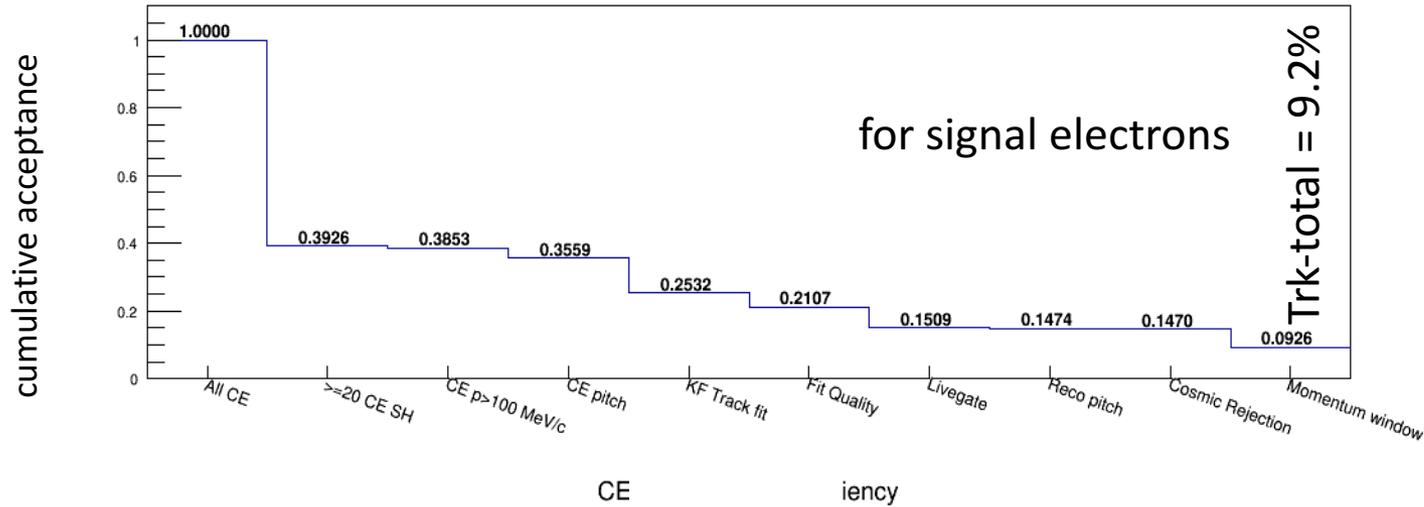
- Mu2e will improve by a factor 10^4 the present best limit!

DIO background

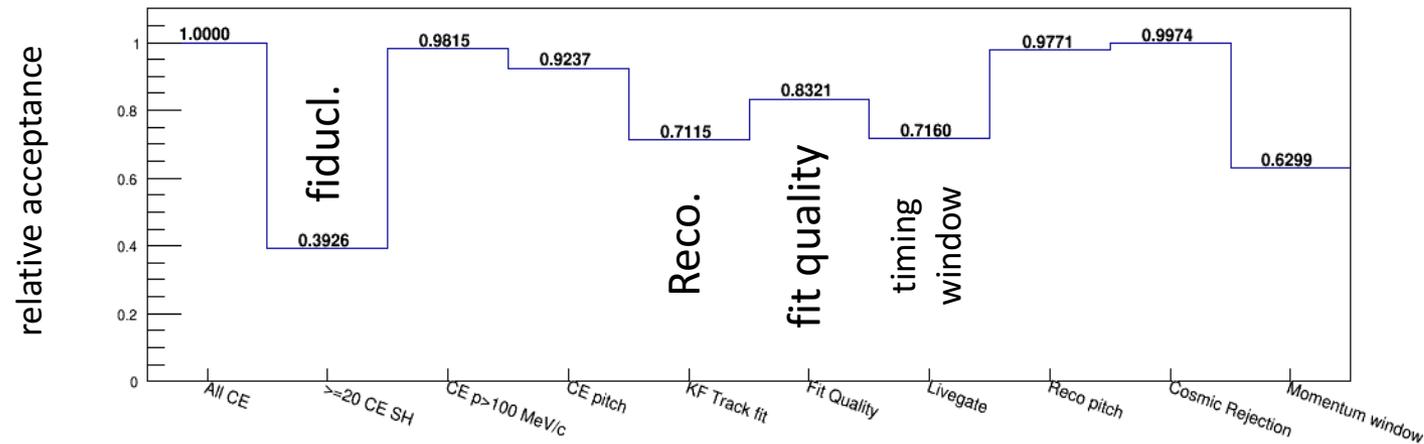


The tracker is blind to most of the DIO background

Track reconstruction and selection



Inefficiency dominated by geometric acceptance

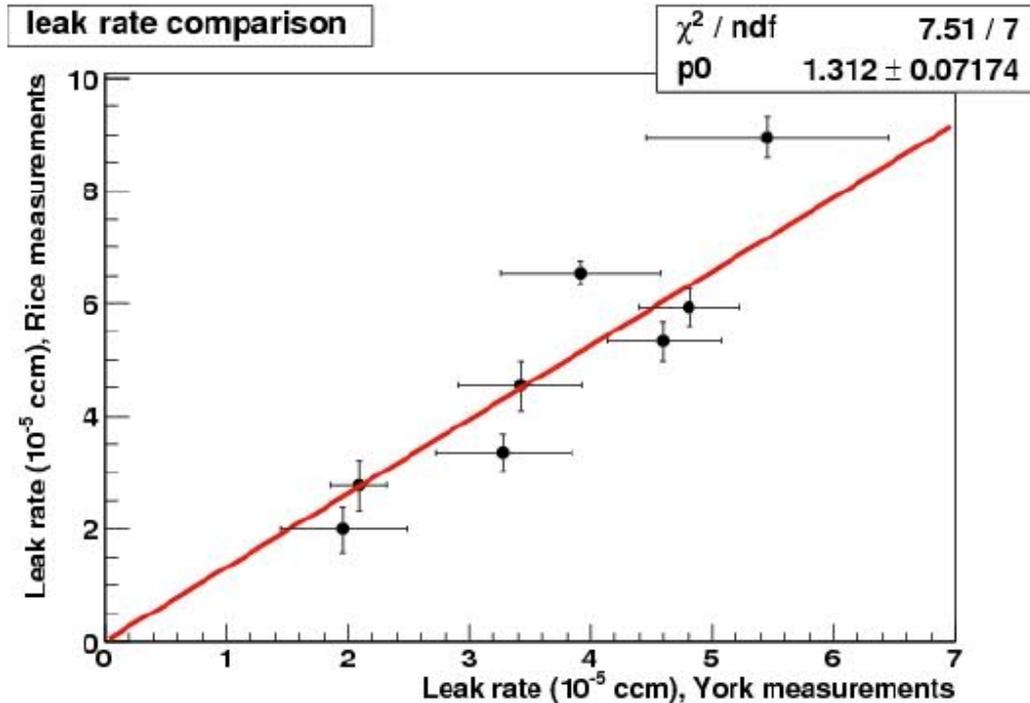


After calorimeter PID and CRV deadtime, Total = 8.5%

Straw leak test

Two methods:

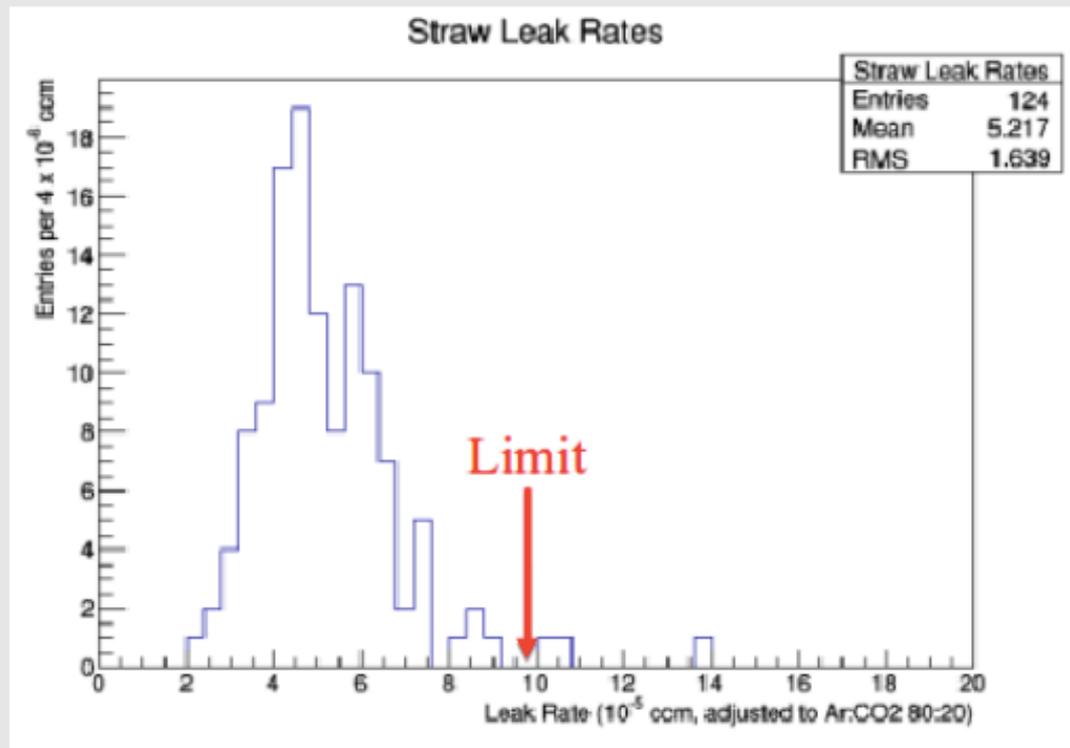
- CO₂ permeation >100 straw/day: do every straw
Needs cross calibration
- Vacuum Absolute measurement
~ 1 per day (sample of straws)



Agrees within uncertainty
in correcting for difference
in diffusion of Argon vs CO₂

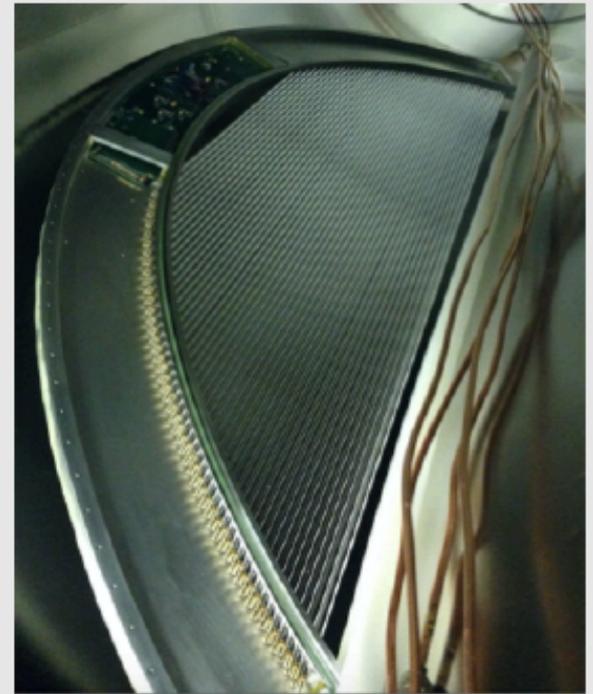
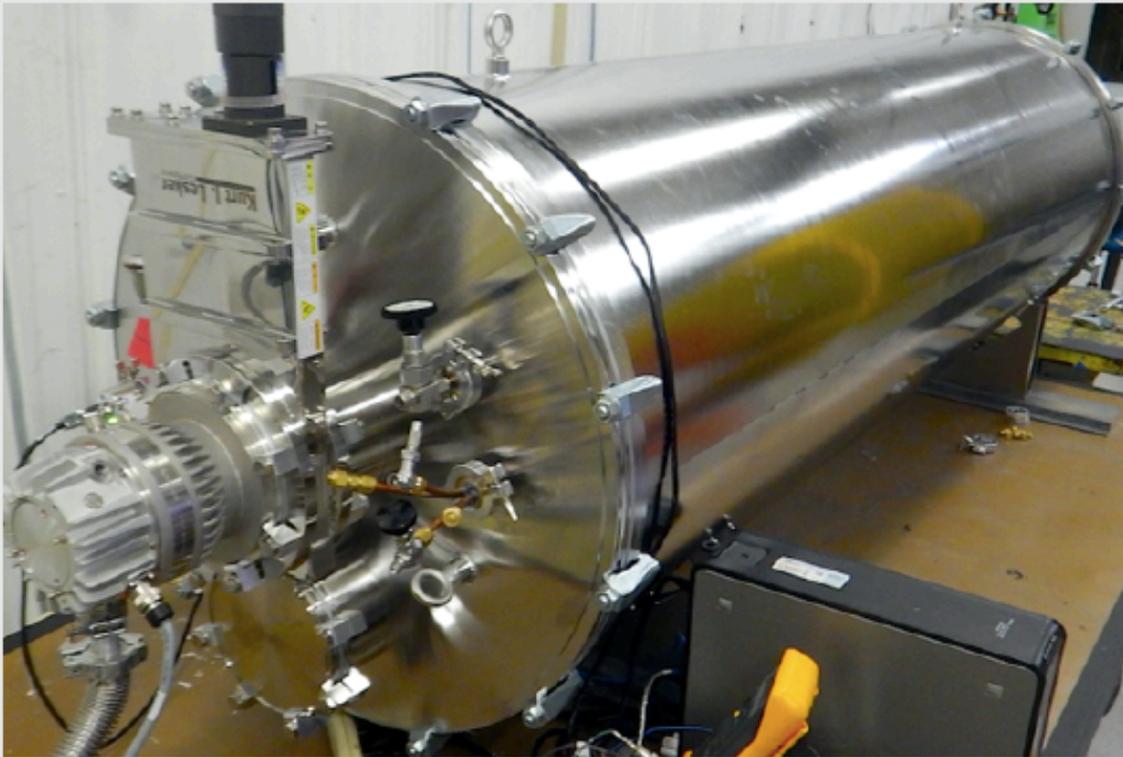
Straw leak test

- The full tracker leak rate limit is $6 \text{ cm}^3/\text{min}$.
- many possible sources
- individual straw leak limit is $9.6 \times 10^{-5} \text{ cm}^3/\text{min}$
- 124 straws tested at FNAL last summer; 121 passed



Panel leak test

- Large vacuum vessel to test 6 panels per day



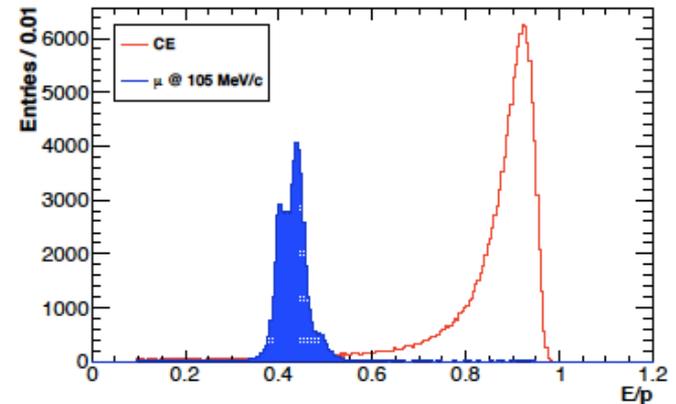
Calorimeter: e/ μ separation

With a CRV inefficiency of 10^{-4} an additional rejection factor of ~ 200 is needed to have < 0.1 fake events from cosmics in the signal window

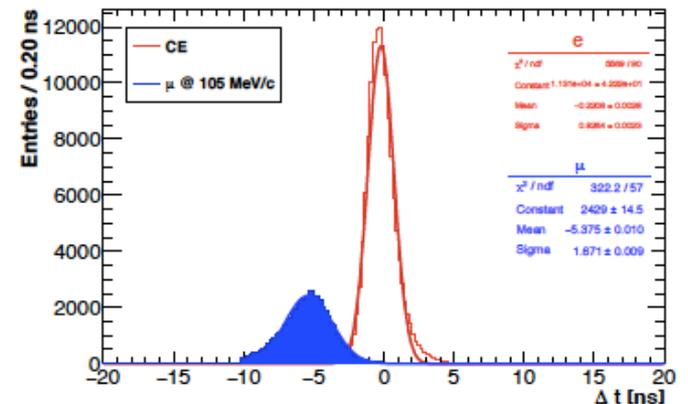
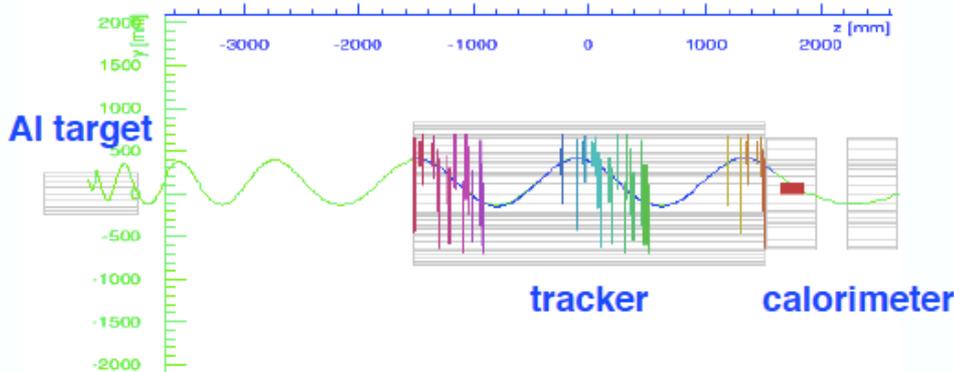
- 105 MeV/c e^- are ultra-relativistic, while 105 MeV/c μ have $\beta \sim 0.7$ and a kinetic energy of ~ 40 MeV
- Likelihood rejection combines

$$\Delta t = t_{\text{track}} - t_{\text{cluster}} \text{ and } E/p:$$

$$\ln L_{e,\mu} = \ln P_{e,\mu}(\Delta t) + \ln P_{e,\mu}(E/p)$$



μ mimicking the CE

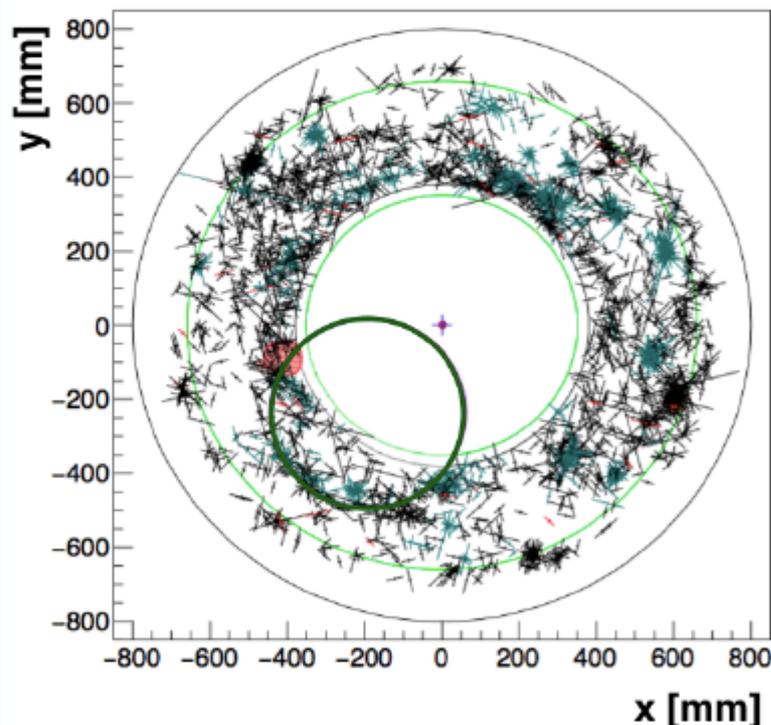


A rejection factor of 200 can be achieved with $\sim 95\%$ efficiency for CE

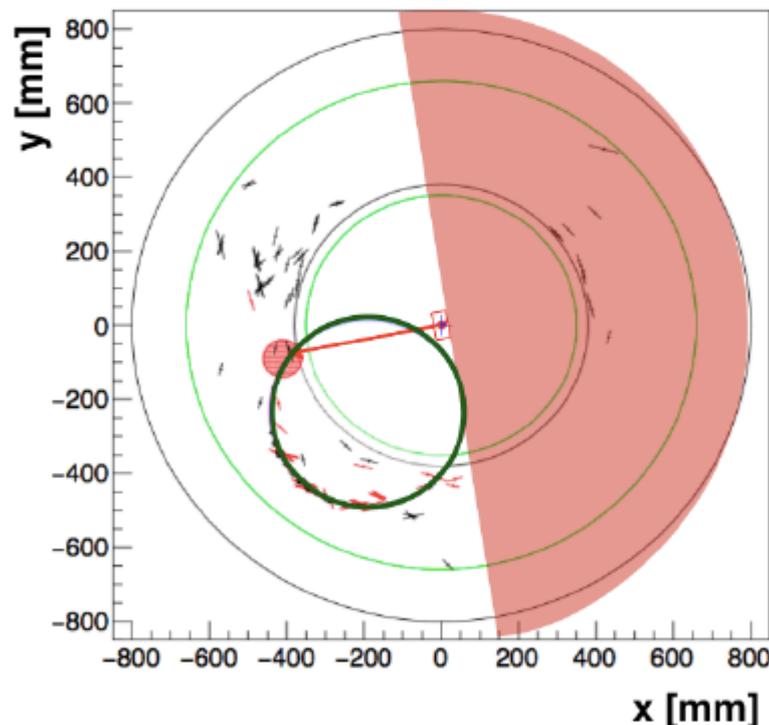
Calorimeter seeded track finder

- Cluster time and position are used for filtering the straw hits:
 - ✓ time window of ~ 80 ns
 - ✓ spatial correlation

no selection



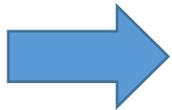
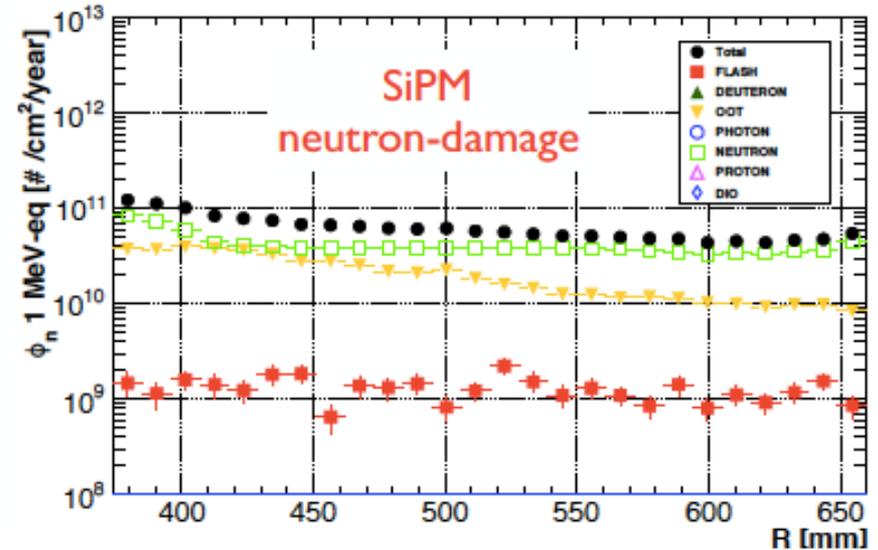
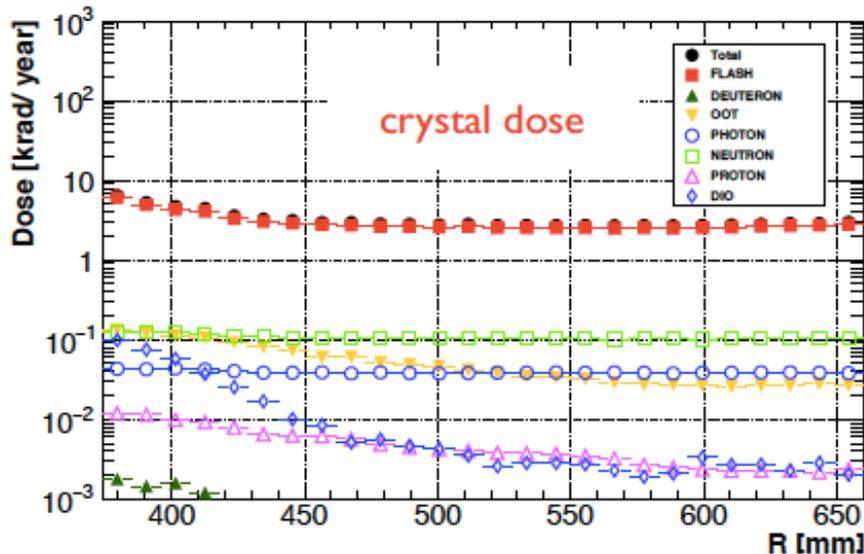
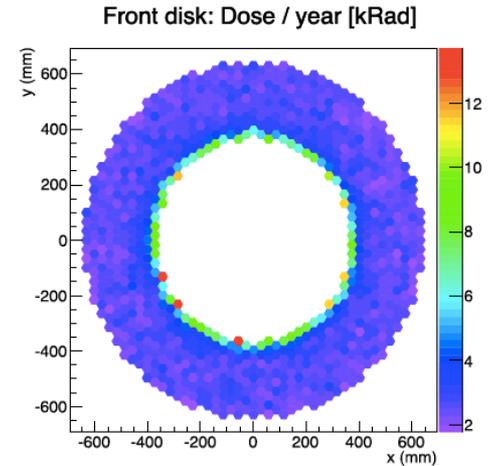
calorimeter selection



- black crosses = straw hits, red circle = calorimeter cluster, green line = CE track

Calorimeter radiation damage

- Calorimeter radiation dose driven by beam flash (interaction of proton beam on target)
- Dose from muon capture is x10 smaller
- Dose is mainly in the inner radius
- Highest dose ~ 10 krad/year
- Highest n flux on crystals $\sim 2 \times 10^{11}$ n/cm²/year
- Highest n flux on SiPM $\sim 10^{11}$ n_{1MeVeq}/cm²/year

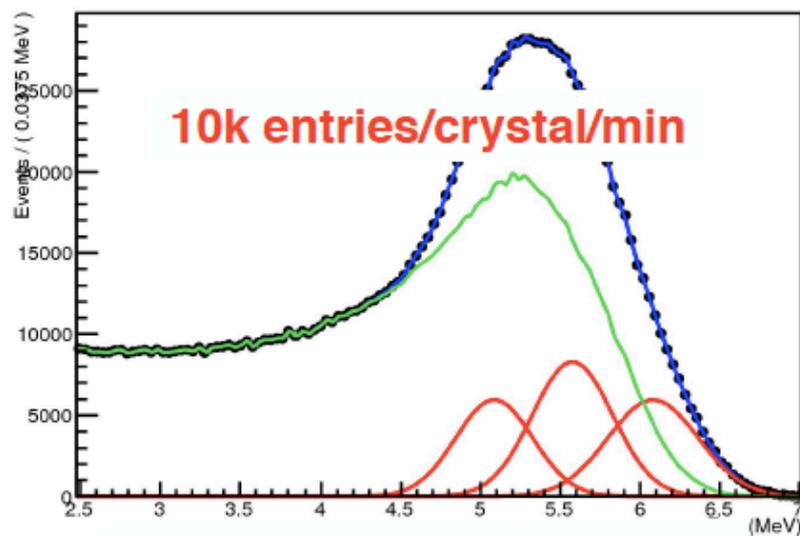


Qualify crystals up to ~ 100 krad, 10^{12} n/cm² This includes a safety factor of 3 for a 3 year run

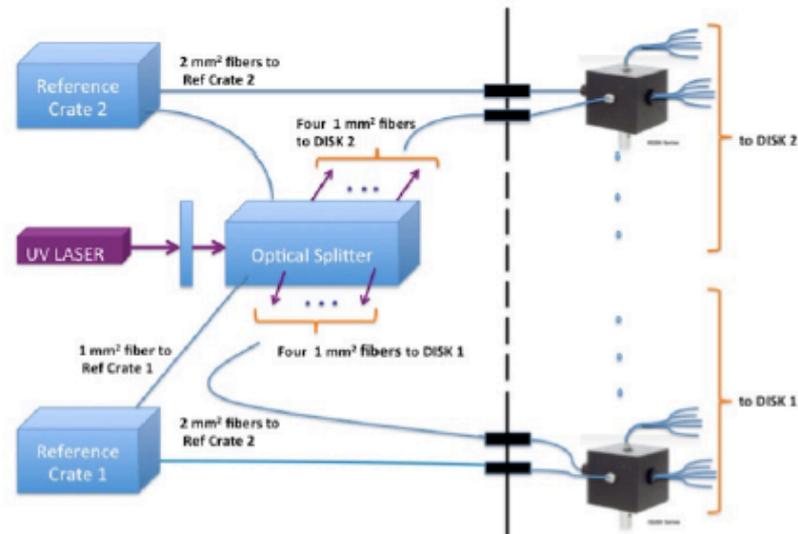
Qualify SiPM up to $\sim 10^{12}$ n_{1MeVeq}/cm²

Calorimeter calibration

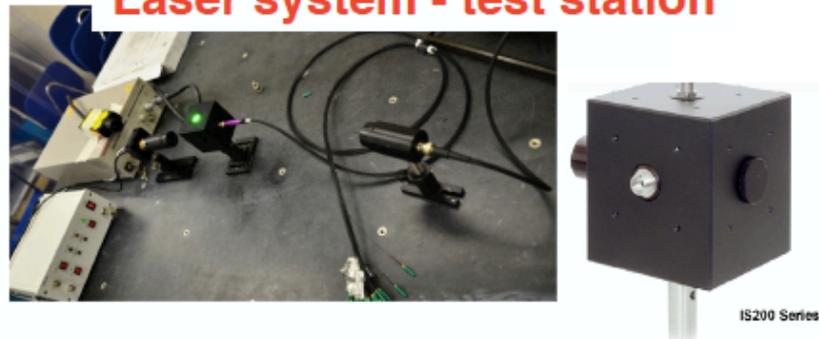
- Liquid source FC 770 + DT generator: 6 MeV + 2 escape peaks
- Laser system to monitor SiPM performance



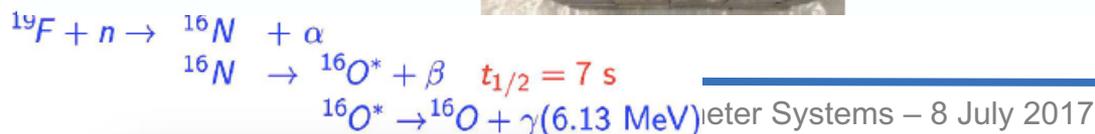
Liquid source prototype



Laser system - test station

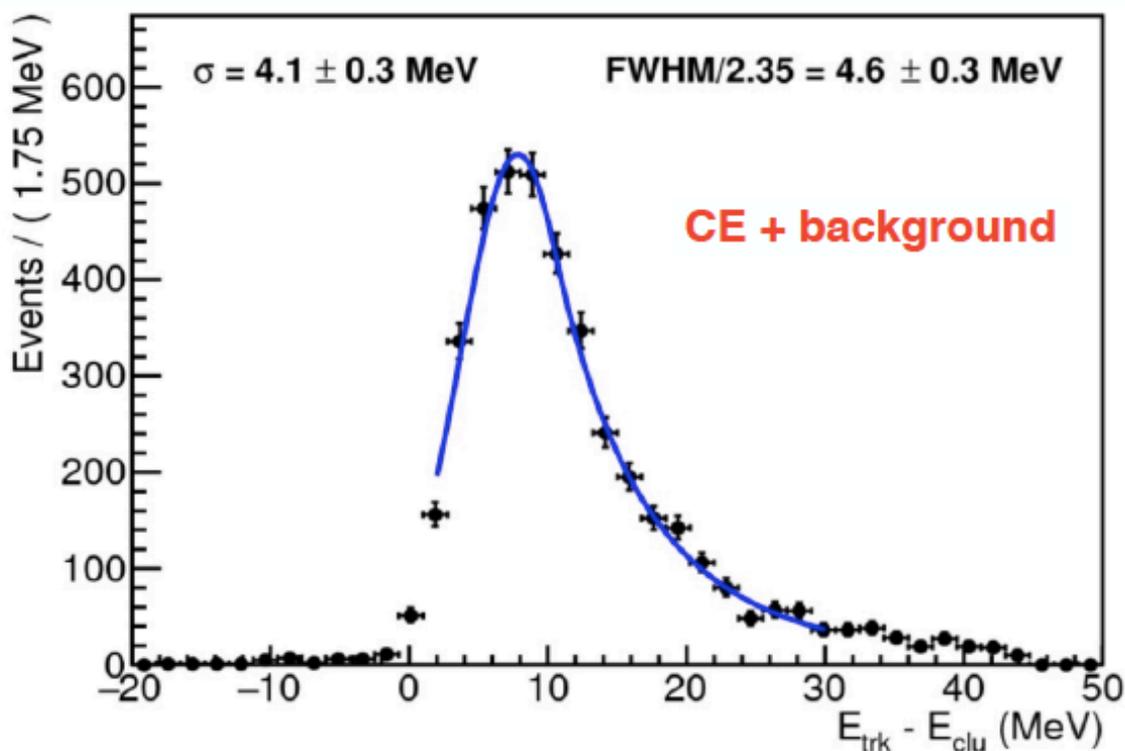


IS200 Series

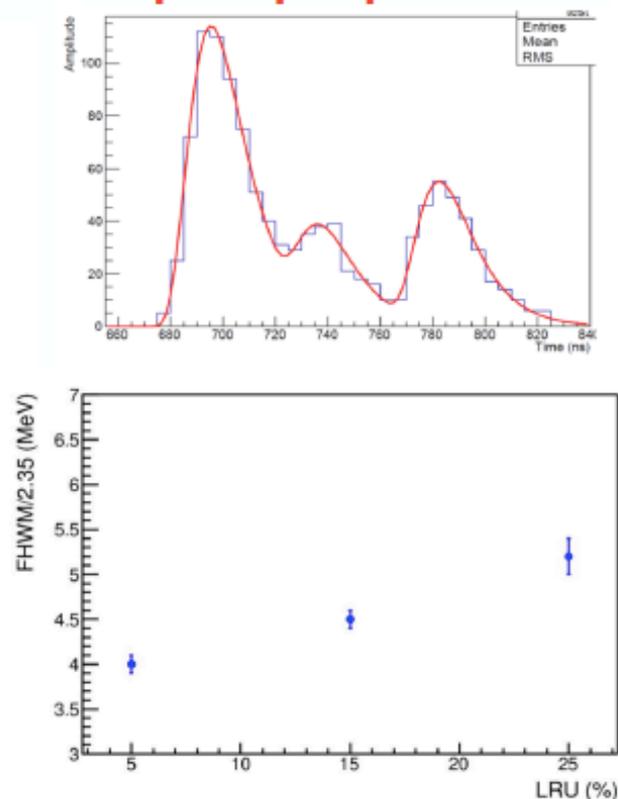


Calorimeter simulation

- Offline simulation including background hits
- Experimental effects included: longitudinal response uniformity (LRU), electronic noise, digitization, etc
- Waveform-based analysis to improve pileup separation

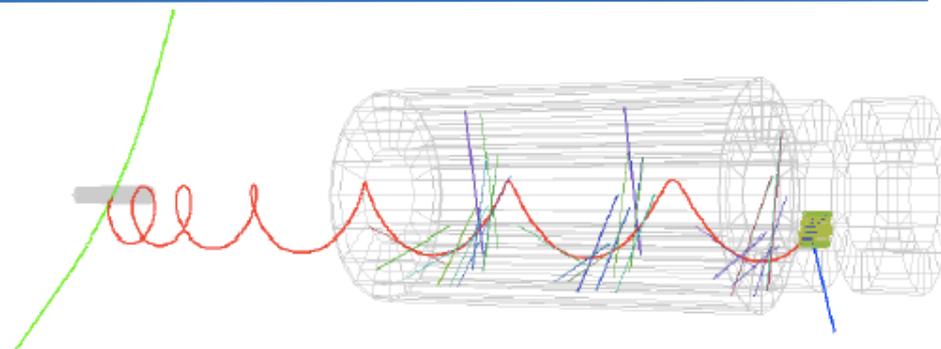
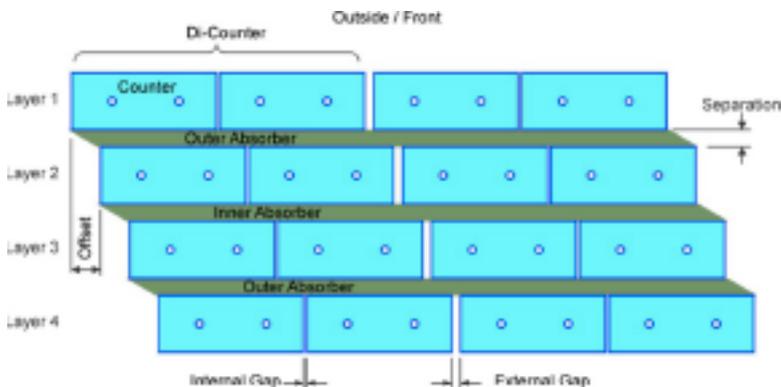


pile-up separation

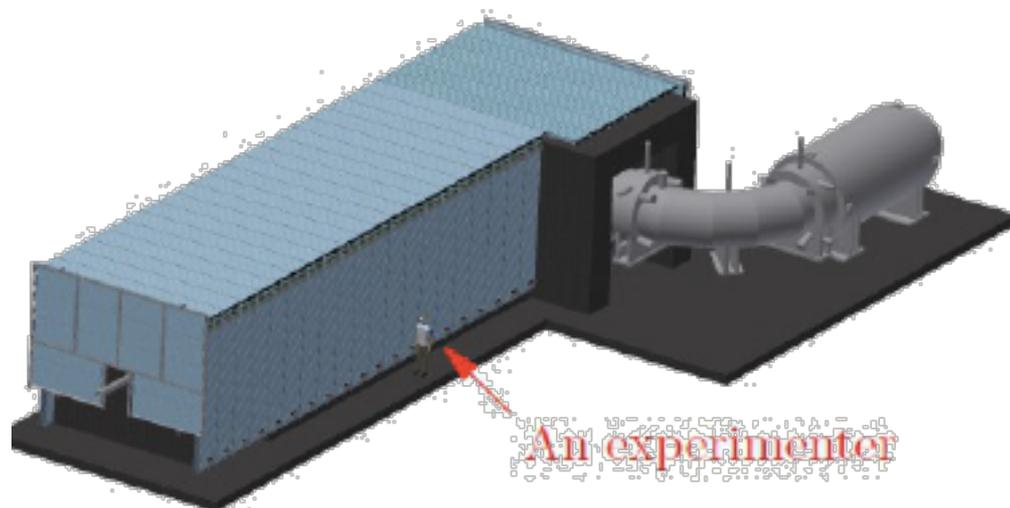


Cosmic Ray Veto

Cosmic ray muons will produce one fake signal event per day without a CRV. The muon itself can fake a 105 MeV e^- or it can knock out an e^- from material in the detector (delta ray), which can fake the signal
 → Passive shielding + PID trk/EMC + CRV



- Four layers of extruded plastic scintillator, (5×2) cm²
- WLS fiber + SiPM readout



Nbkg (cosmics): 0.05 (3 years run)

- CR veto inefficiency = 10^{-4}
- $\epsilon_{\text{CRV}} = 99.99\%$
- $\epsilon_{\text{CRVplane}} = 99.6\%$ (14 pe/MeV for longest module)
- Thr @ 0.5-1 MeV to reduce n/γ , avoiding dead time (m.i.p. 2 MeV/cm)
- $\frac{3}{4}$ layers hit: 125 ns veto

The COMET Experiment

phase I

phase II

