

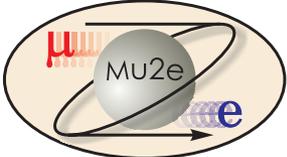
# The Mu2e Experiment

## FNAL Users' meeting 2012



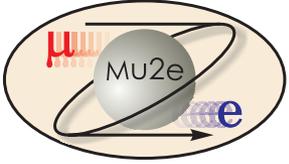
Emma J Barnes  
Boston University  
June 12-13<sup>th</sup> 2012

<http://mu2e.fnal.gov/>



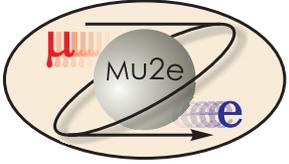
# Summary of Talk

- *Charged Lepton Flavor Violation (CLFV)*
- *Observing CLFV.*
- *The physics of muon to electron conversion*
- *The Mu2e experiment*
- *Backgrounds that can limit sensitivity*
- *Current status*



# Charged Lepton Flavor violation

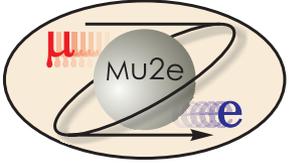
$$\mu^{-} \rightarrow e^{-} \gamma$$



# Charged Lepton Flavor violation

$$\mu^- \rightarrow e^- \gamma$$

$$\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$$



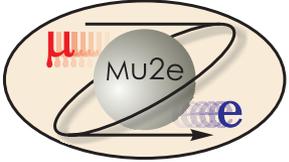
# Charged Lepton Flavor violation

$$\mu^- \rightarrow e^- \gamma$$

$$L_\mu(1) \rightarrow L_e(1)$$

$$\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$$

$$L_\mu(1) \rightarrow L_e(1) + L_{\bar{\nu}_e}(-1) + L_{\nu_\mu}(1)$$



# Charged Lepton Flavor violation

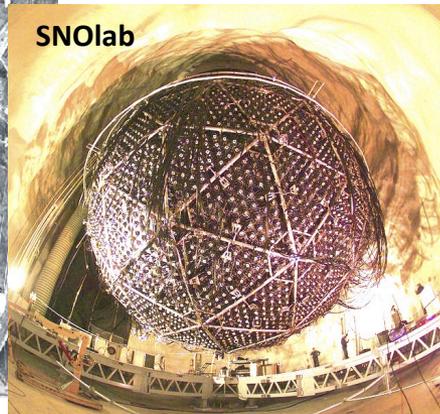
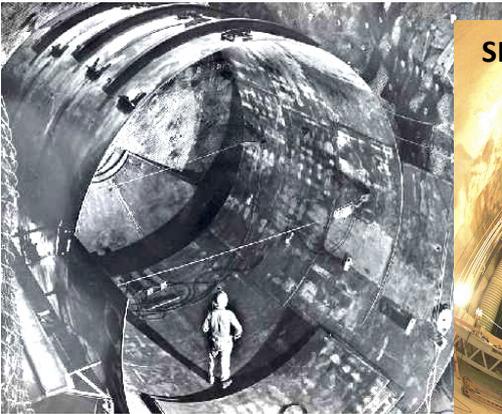
$$\mu^- \rightarrow e^- \gamma$$

$$L_\mu(1) \rightarrow L_e(1)$$

$$\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$$

$$L_\mu(1) \rightarrow L_e(1) + L_{\nu_e}(-1) + L_{\nu_\mu}(1)$$

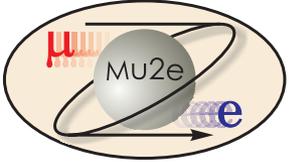
Homestake mine



<http://www.sno lab.ca/content/sno-detector>

- Neutral lepton (neutrino) flavor violation is now an accepted fact.
- Numerous experiments have confirmed many of the parameters of neutrino mixing.

<http://jacksonville.com/entertainment/2012-03-20/story/sky-guy-scientists-hold-fast-speed-limit-light#ixzz1v8b1lsN1>



# Charged Lepton Flavor violation

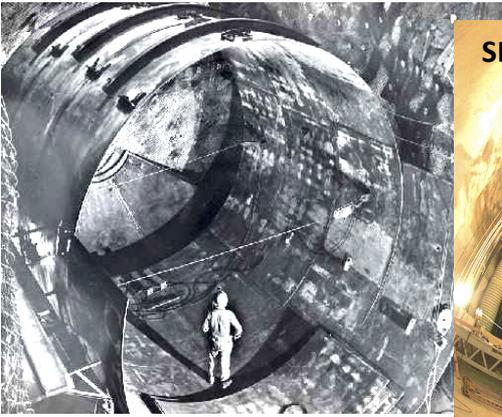
$$\mu^- \rightarrow e^- \gamma$$

$$L_\mu(1) \rightarrow L_e(1)$$

$$\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$$

$$L_\mu(1) \rightarrow L_e(1) + L_{\nu_e}(-1) + L_{\nu_\mu}(1)$$

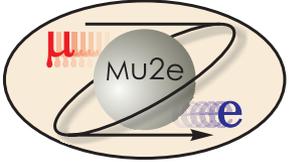
Homestake mine



<http://www.snolab.ca/content/sno-detector>

- Neutral lepton (neutrino) flavor violation is now an accepted fact.
- Numerous experiments have confirmed many of the parameters of neutrino mixing.
- CLFV in the SM can occur through the intermediate mixing of massive neutrinos.
- The rate depends on the neutrino mass splitting and couplings

<http://jacksonville.com/entertainment/2012-03-20/story/sky-guy-scientists-hold-fast-speed-limit-light#ixzz1v8b1lsN1>



# Charged Lepton Flavor violation

$$\mu^- \rightarrow e^- \gamma$$

$$\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$$

$$L_\mu(1) \rightarrow L_e(1)$$

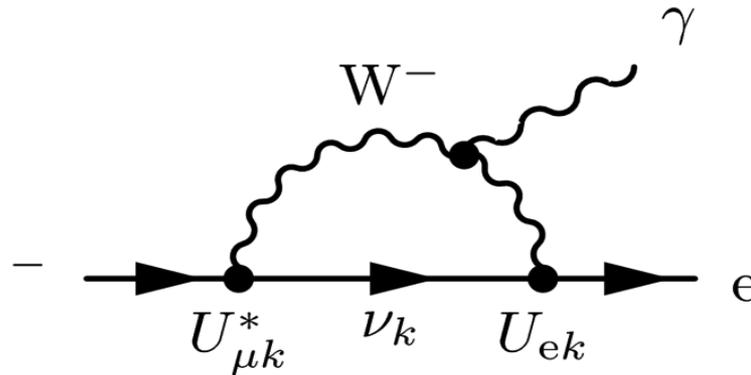
$$L_\mu(1) \rightarrow L_e(1) + L_{\nu_e}(-1) + L_{\nu_\mu}(1)$$

Homestake mine

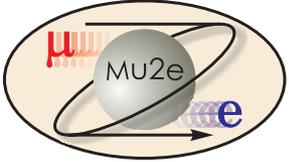


<http://www.snolab.ca/content/sno-detector>

- Neutral lepton (neutrino) flavor violation is now an accepted fact.
- Numerous experiments have confirmed many of the parameters of neutrino mixing.
- CLFV in the SM can occur through the intermediate mixing of massive neutrinos.
- The rate depends on the neutrino mass splitting and couplings



<http://jacksonville.com/entertainment/2012-03-20/story/sky-guy-scientists-hold-fast-speed-limit-light#ixzz1v8b1lsN1>



# Charged Lepton Flavor violation

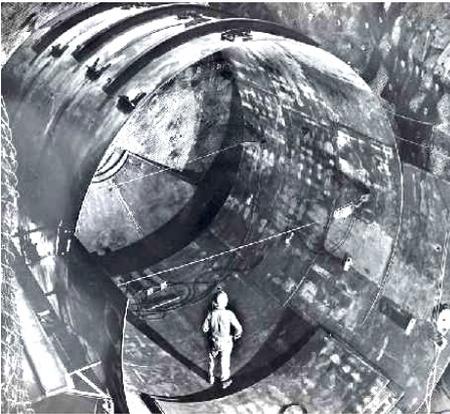
$$\mu^- \rightarrow e^- \gamma$$

$$L_\mu(1) \rightarrow L_e(1)$$

$$\mu^- \rightarrow e^- + \bar{\nu}_e + \nu_\mu$$

$$L_\mu(1) \rightarrow L_e(1) + L_{\bar{\nu}_e}(-1) + L_{\nu_\mu}(1)$$

Homestake mine

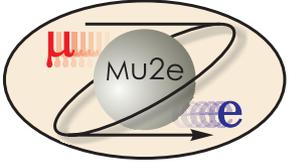


<http://www.snolab.ca/content/sno-detector>

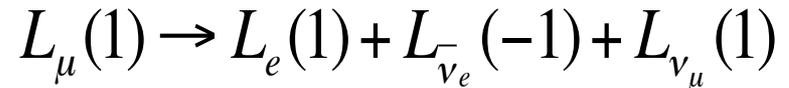
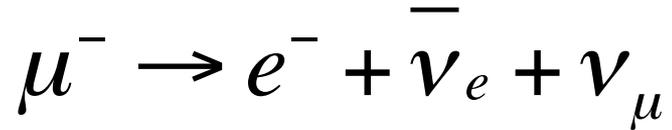
- Neutral lepton (neutrino) flavor violation is now an accepted fact.
- Numerous experiments have confirmed many of the parameters of neutrino mixing.
- CLFV in the SM can occur through the intermediate mixing of massive neutrinos.
- The rate depends on the neutrino mass splitting and couplings

$$Br(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i}^* U_{ei} \frac{\Delta m_{1i}^2}{M_W^2} \right|^2 \sim 10^{-54}$$

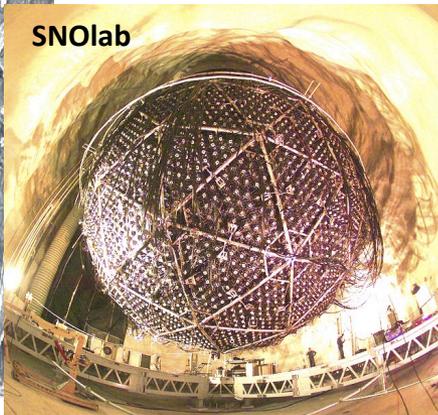
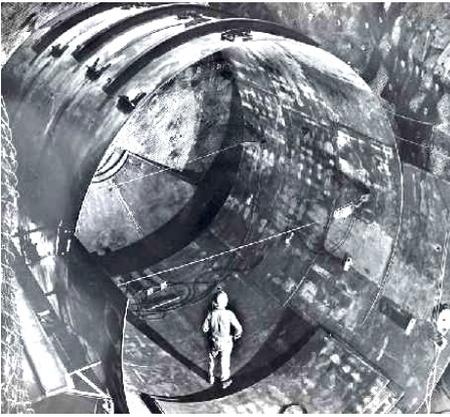
<http://jacksonville.com/entertainment/2012-03-20/story/sky-guy-scientists-hold-fast-speed-limit-light#ixzz1v8b1lsN1>



# Charged Lepton Flavor violation



Homestake mine



<http://www.sno lab.ca/content/sno-detector>

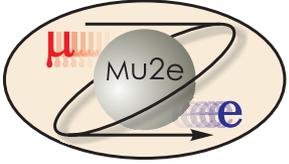
- Neutral lepton (neutrino) flavor violation is now an accepted fact.
- Numerous experiments have confirmed many of the parameters of neutrino mixing.
- CLFV in the SM can occur through the intermediate mixing of massive neutrinos.
- The rate depends on the neutrino mass splitting and couplings

**No Standard model background.**

$$Br(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i}^* U_{ei} \frac{\Delta m_{1i}^2}{M_W^2} \right|^2 \sim 10^{-54}$$

So, if observed it would be unambiguous sign of New Physics

<http://jacksonville.com/entertainment/2012-03-20/story/sky-guy-scientists-hold-fast-speed-limit-light#ixzz1v8b1lsN1>



# Searching for CLFV

3 rare muon processes stand out

$$\mu^{\pm} \rightarrow e^{\pm} \gamma$$

MEGA, MEG, and others...

$$\mu^{\pm} \rightarrow e^{\pm} e^{+} e^{-}$$

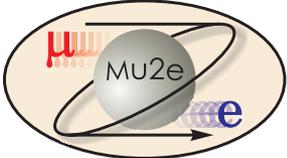
$\text{Br}(\mu^{+} \rightarrow e\gamma) < 2.4 \times 10^{-12}$  [MEG 2011]

$\text{Br}(\mu^{+} \rightarrow e^{+}e^{-}e^{+}) < 1 \times 10^{-12}$  [SINDRUM-I, 1988]

$$\mu^{-} N \rightarrow e^{-} N$$

**Mu2e, comet, SINDRUM**

$R_{\mu e} < 7 \times 10^{-13}$  [SINDRUM-II, 2006]



# Searching for CLFV

## 3 rare muon processes stand out

$$\mu^{\pm} \rightarrow e^{\pm} \gamma$$

**MEGA, MEG, and others...**

$$\mu^{\pm} \rightarrow e^{\pm} e^{+} e^{-}$$

$$\text{Br}(\mu^{+} \rightarrow e\gamma) < 2.4 \times 10^{-12} \text{ [MEG 2011]}$$

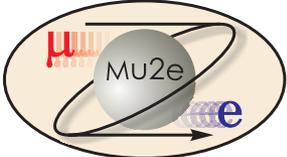
$$\text{Br}(\mu^{+} \rightarrow e^{+} e^{-} e^{+}) < 1 \times 10^{-12} \text{ [SINDRUM-I, 1988]}$$

- *Signal is combination of final state particles.*
- *Accidental coincidences limit usable  $\mu$  rate*

$$\mu^{-} N \rightarrow e^{-} N$$

**Mu2e, comet, SINDRUM**

$$R_{\mu e} < 7 \times 10^{-13} \text{ [SINDRUM-II, 2006]}$$



# Searching for CLFV

## 3 rare muon processes stand out

$$\mu^{\pm} \rightarrow e^{\pm} \gamma$$

MEGA, MEG, and others...

$$\mu^{\pm} \rightarrow e^{\pm} e^{+} e^{-}$$

$$\text{Br}(\mu^{+} \rightarrow e\gamma) < 2.4 \times 10^{-12} \text{ [MEG 2011]}$$

$$\text{Br}(\mu^{+} \rightarrow e^{+}e^{-}e^{+}) < 1 \times 10^{-12} \text{ [SINDRUM-I, 1988]}$$

- Signal is combination of final state particles.
- Accidental coincidences limit usable  $\mu$  rate

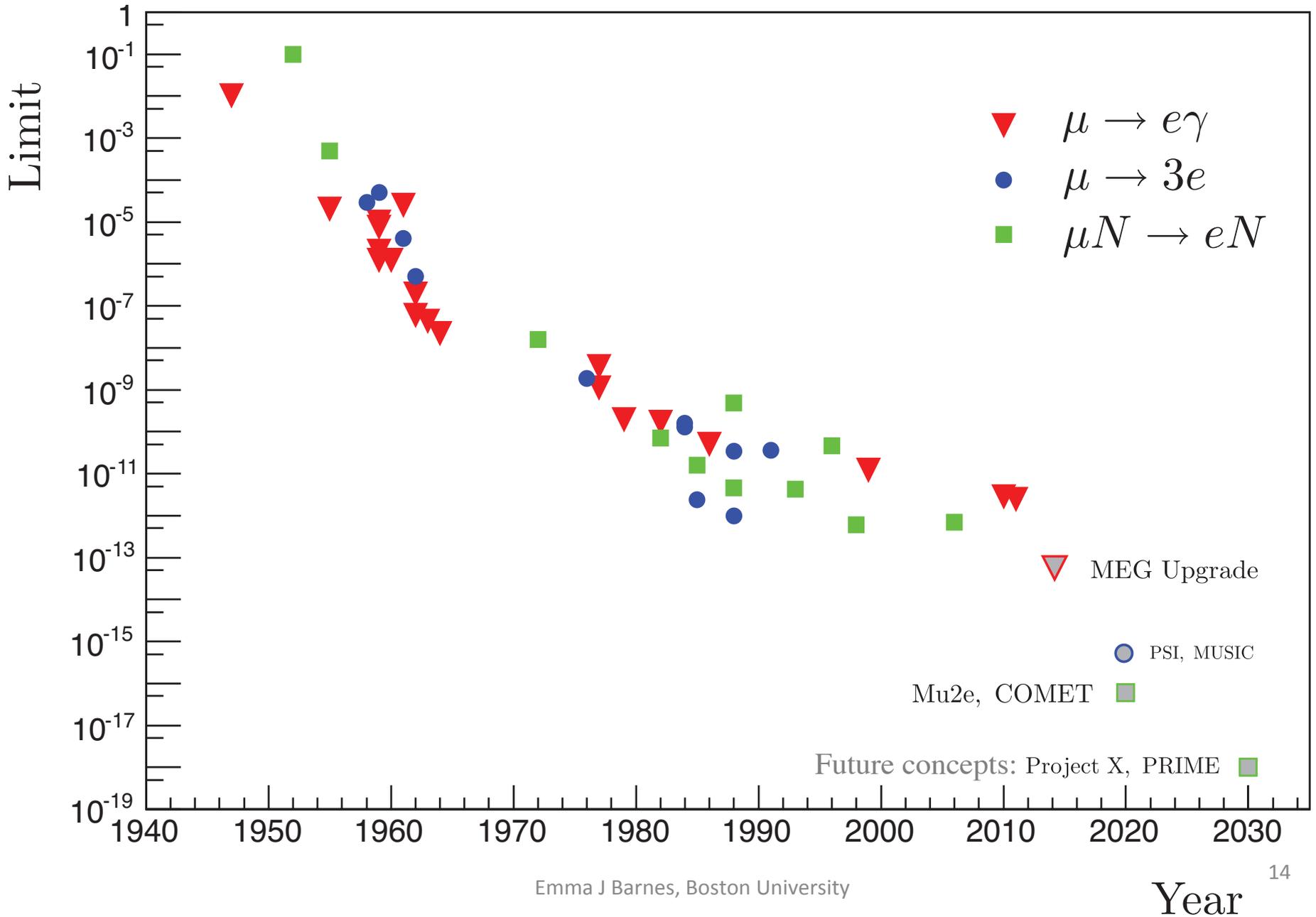
$$\mu^{-} N \rightarrow e^{-} N$$

- Signal is a monoenergetic electron.
- Can take more muons/s.

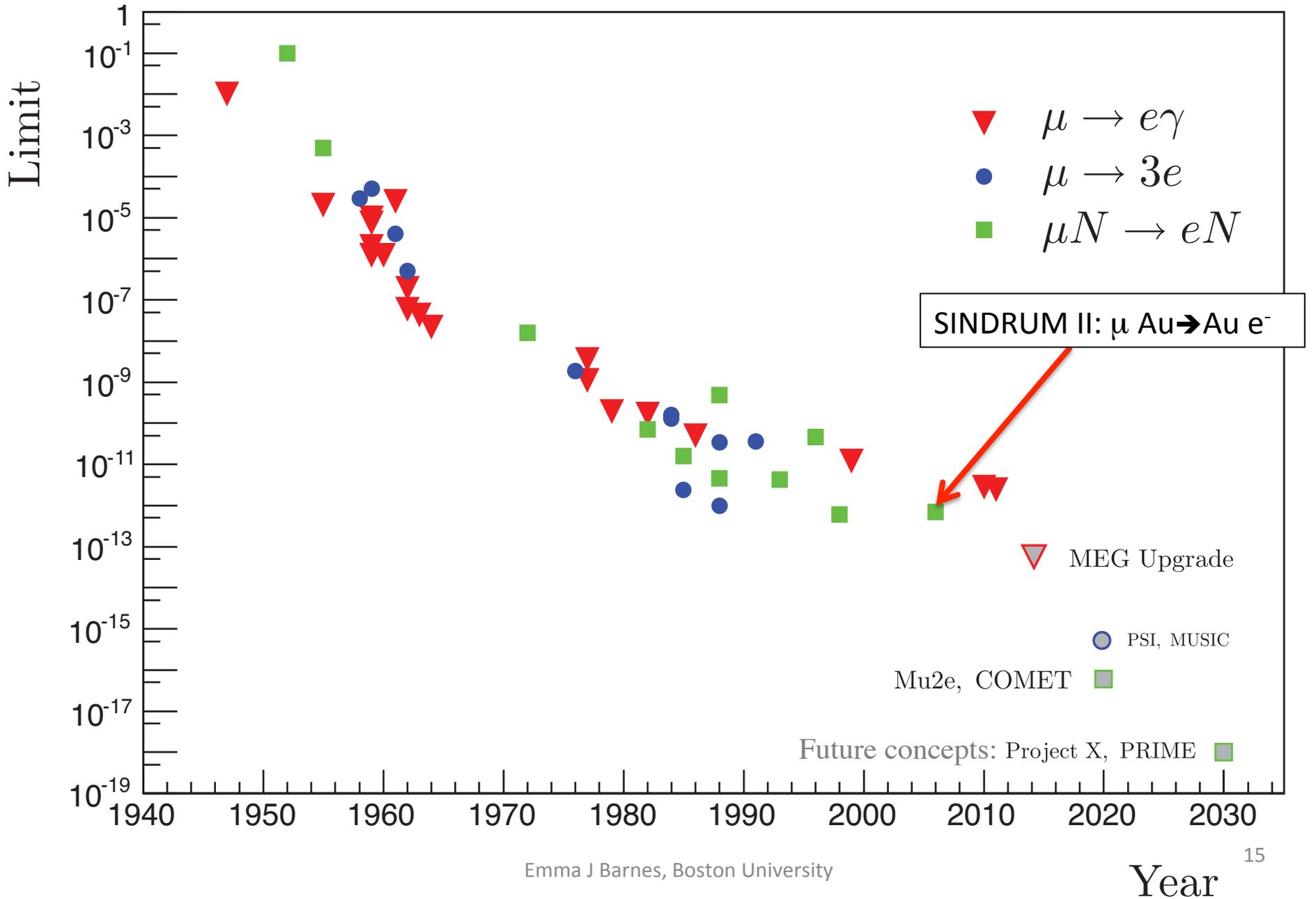
**Mu2e, comet, SINDRUM**

$$R_{\mu e} < 7 \times 10^{-13} \text{ [SINDRUM-II, 2006]}$$

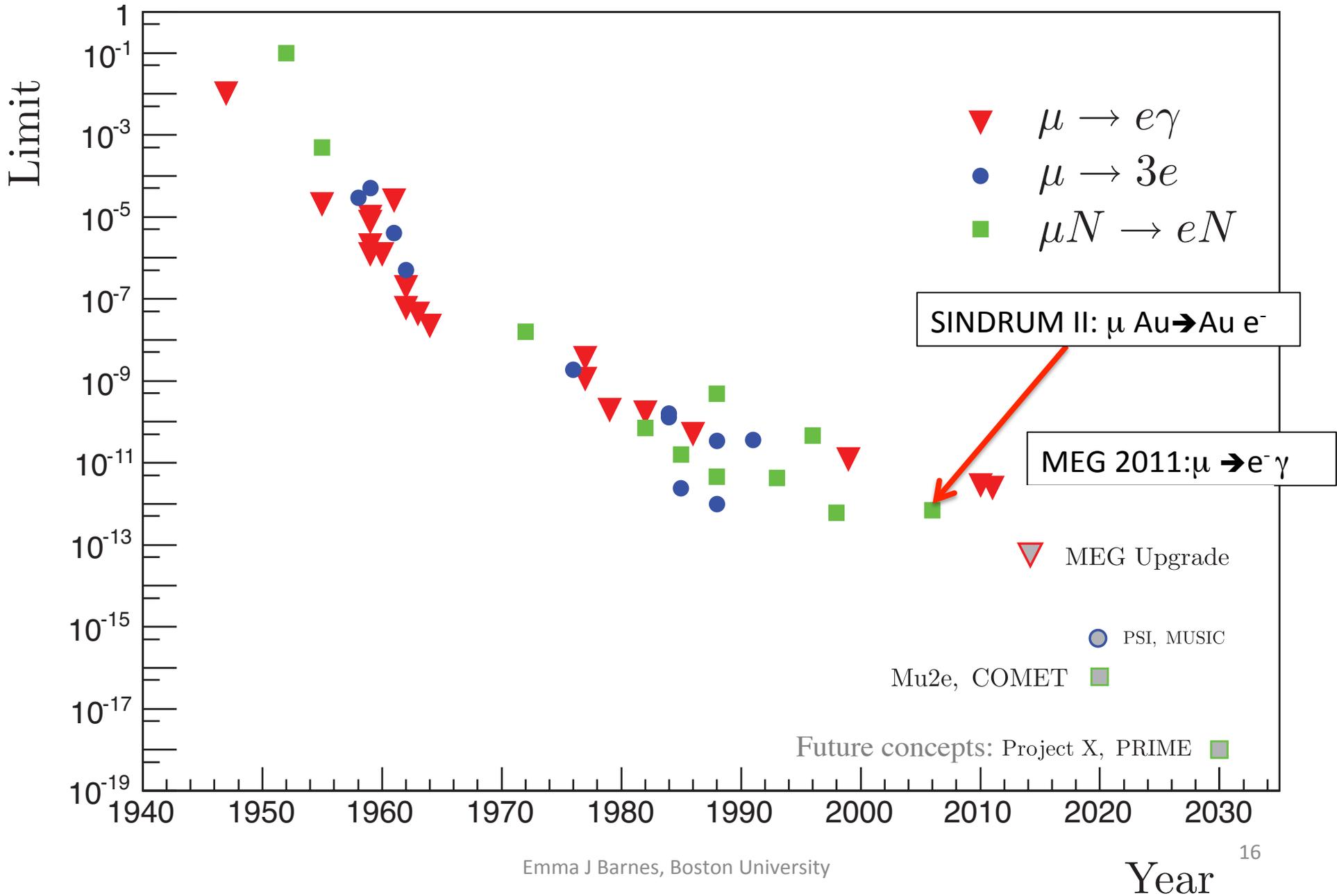
# History of CLFV searches



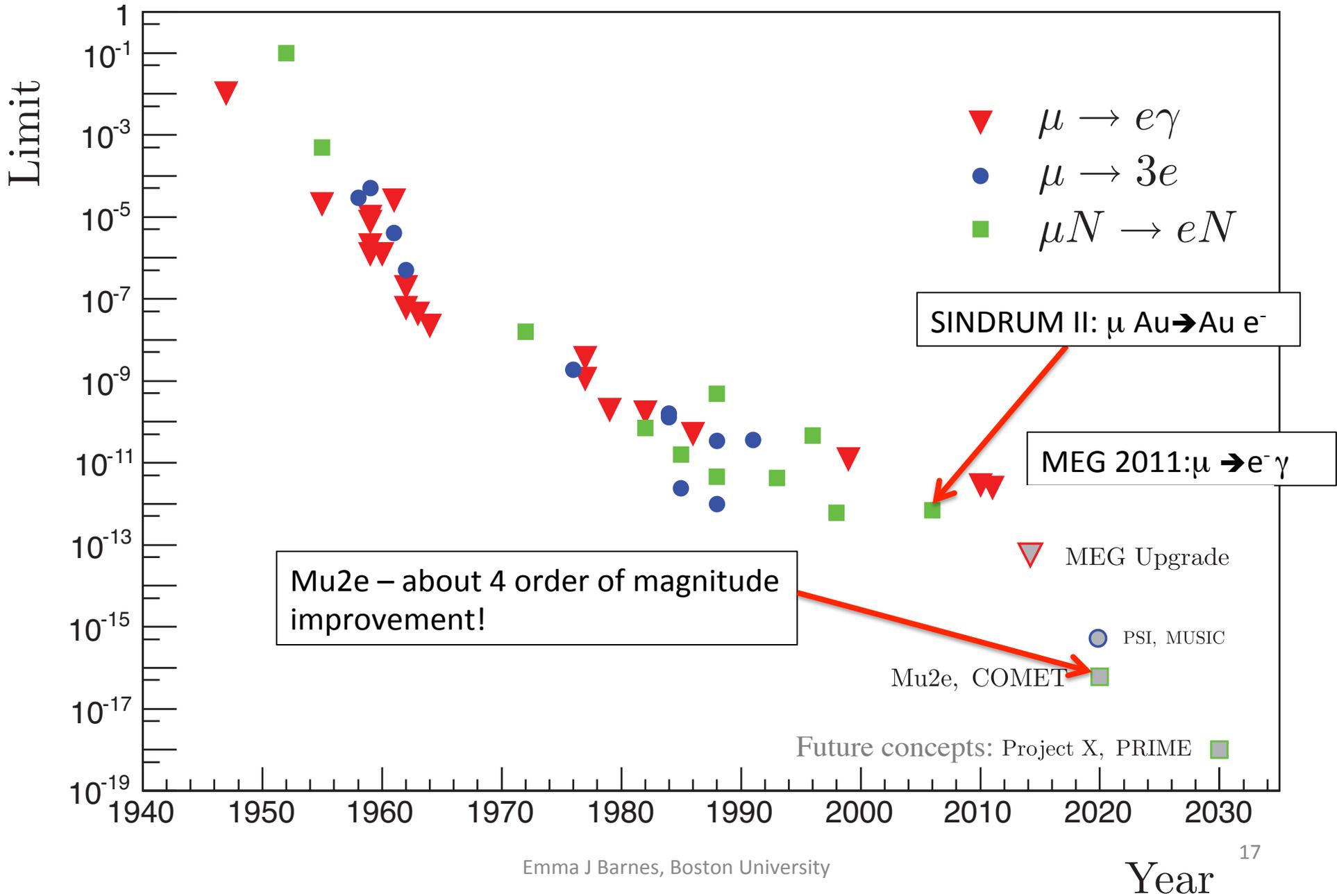
# History of CLFV searches

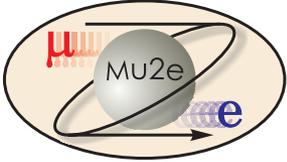


# History of CLFV searches



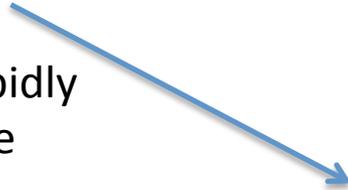
# History of CLFV searches



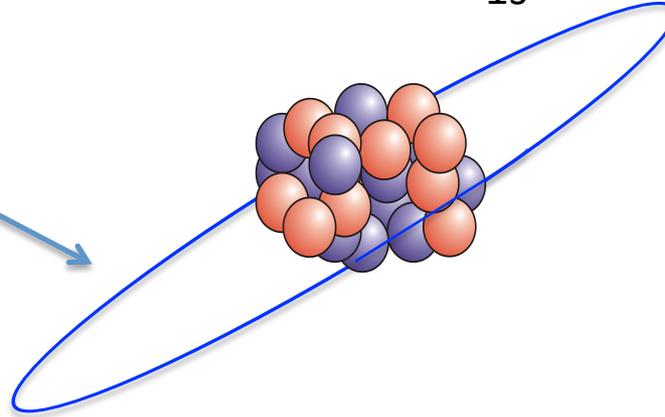


# $\mu \rightarrow e$ conversion

$\mu^-$



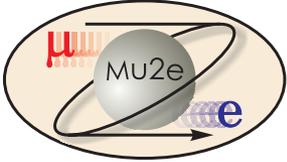
1s



Muons are stopped and rapidly (ps) cascade to the 1s state producing X-Rays.

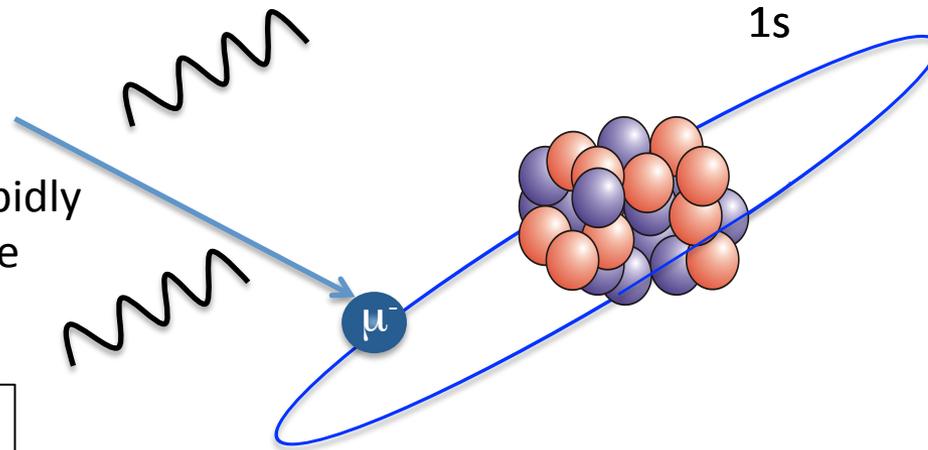
Transition Energy for Al

3d $\rightarrow$ 2p	66 keV
2p $\rightarrow$ 1s	356 keV
3d $\rightarrow$ 1s	423 keV
4p $\rightarrow$ 1s	446 keV



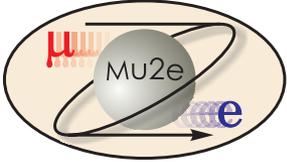
# $\mu \rightarrow e$ conversion

Muons are stopped and rapidly (ps) cascade to the 1s state producing X-Rays.



### Transition Energy for Al

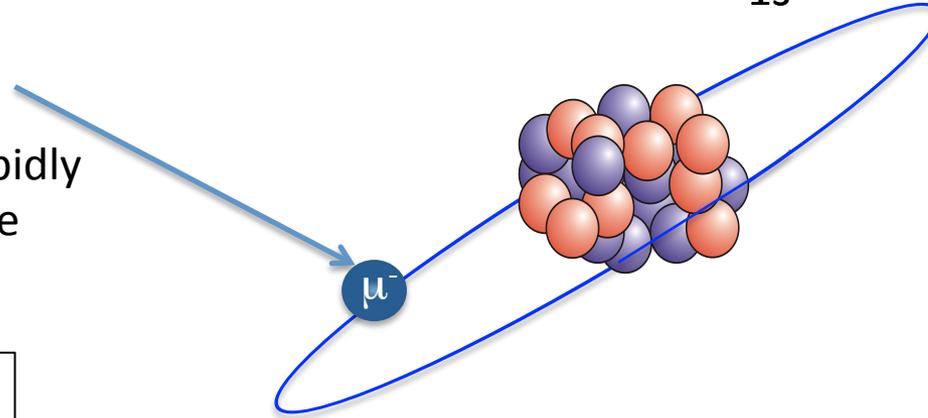
3d → 2p	66 keV
2p → 1s	356 keV
3d → 1s	423 keV
4p → 1s	446 keV



# $\mu \rightarrow e$ conversion

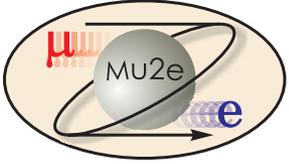
40 % Muon decay in orbit  
60 % Muon capture  $1s$

Muon are stopped and rapidly (ps) cascade to the  $1s$  state producing X-Rays.



Transition Energy for Al

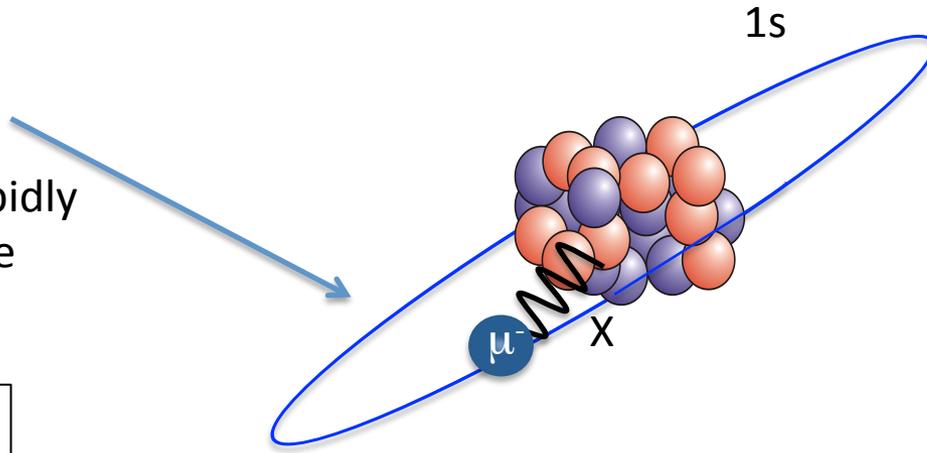
$3d \rightarrow 2p$	66 keV
$2p \rightarrow 1s$	356 keV
$3d \rightarrow 1s$	423 keV
$4p \rightarrow 1s$	446 keV

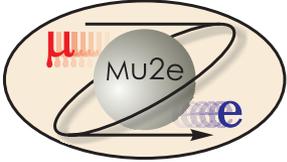


# $\mu \rightarrow e$ conversion

Muons are stopped and rapidly (ps) cascade to the 1s state producing X-Rays.

Transition Energy for Al	
3d $\rightarrow$ 2p	66 keV
2p $\rightarrow$ 1s	356 keV
3d $\rightarrow$ 1s	423 keV
4p $\rightarrow$ 1s	446 keV

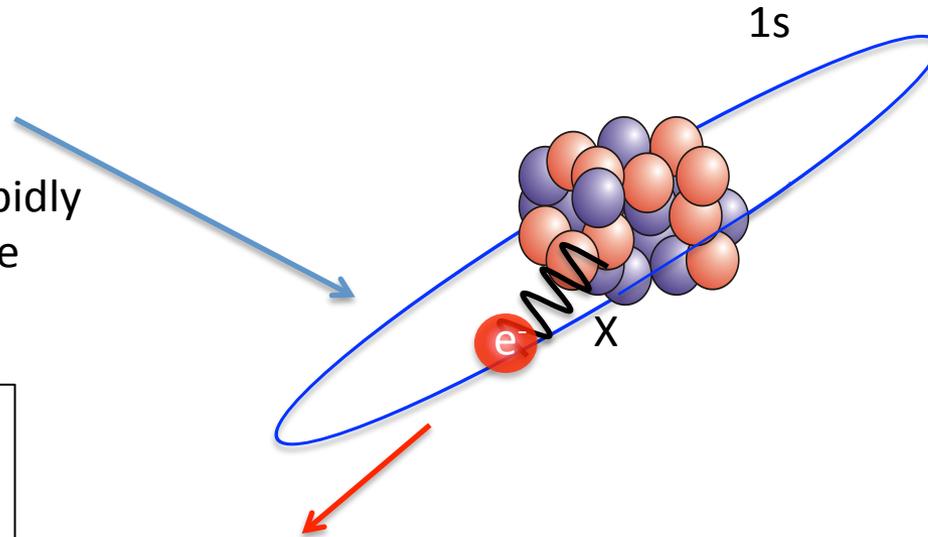


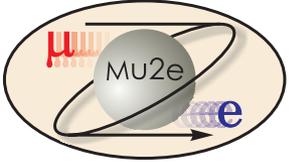


# $\mu \rightarrow e$ conversion

Muons are stopped and rapidly (ps) cascade to the 1s state producing X-Rays.

Transition Energy for Al	
3d $\rightarrow$ 2p	66 keV
2p $\rightarrow$ 1s	356 keV
3d $\rightarrow$ 1s	423 keV
4p $\rightarrow$ 1s	446 keV

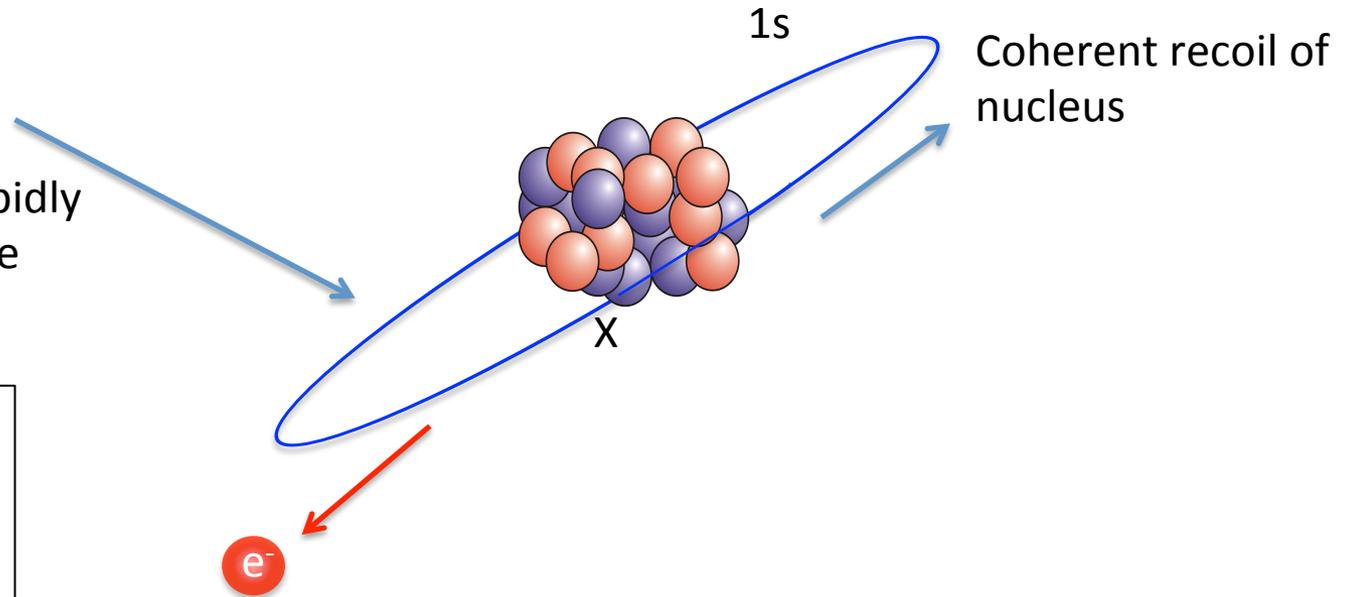


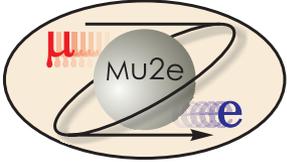


# $\mu \rightarrow e$ conversion

Muons are stopped and rapidly (ps) cascade to the 1s state producing X-Rays.

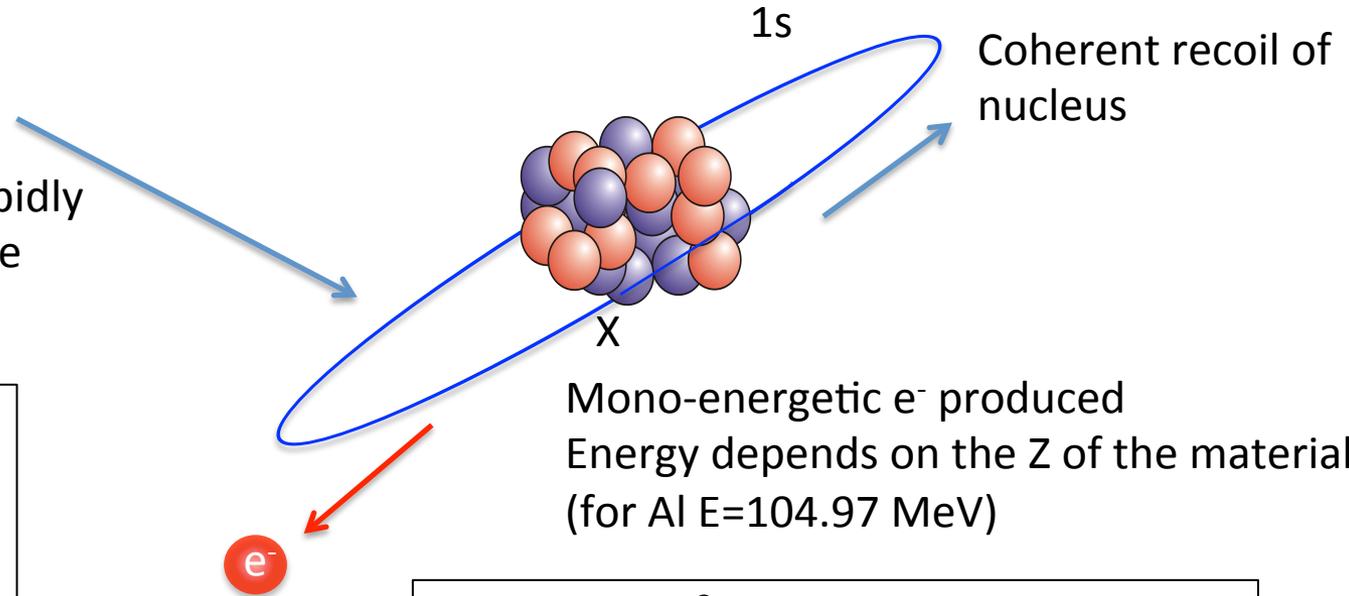
Transition Energy for Al	
3d $\rightarrow$ 2p	66 keV
2p $\rightarrow$ 1s	356 keV
3d $\rightarrow$ 1s	423 keV
4p $\rightarrow$ 1s	446 keV





# $\mu \rightarrow e$ conversion

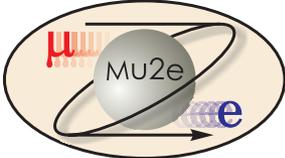
Muons are stopped and rapidly (ps) cascade to the 1s state producing X-Rays.



Transition Energy for Al	
3d → 2p	66 keV
2p → 1s	356 keV
3d → 1s	423 keV
4p → 1s	446 keV

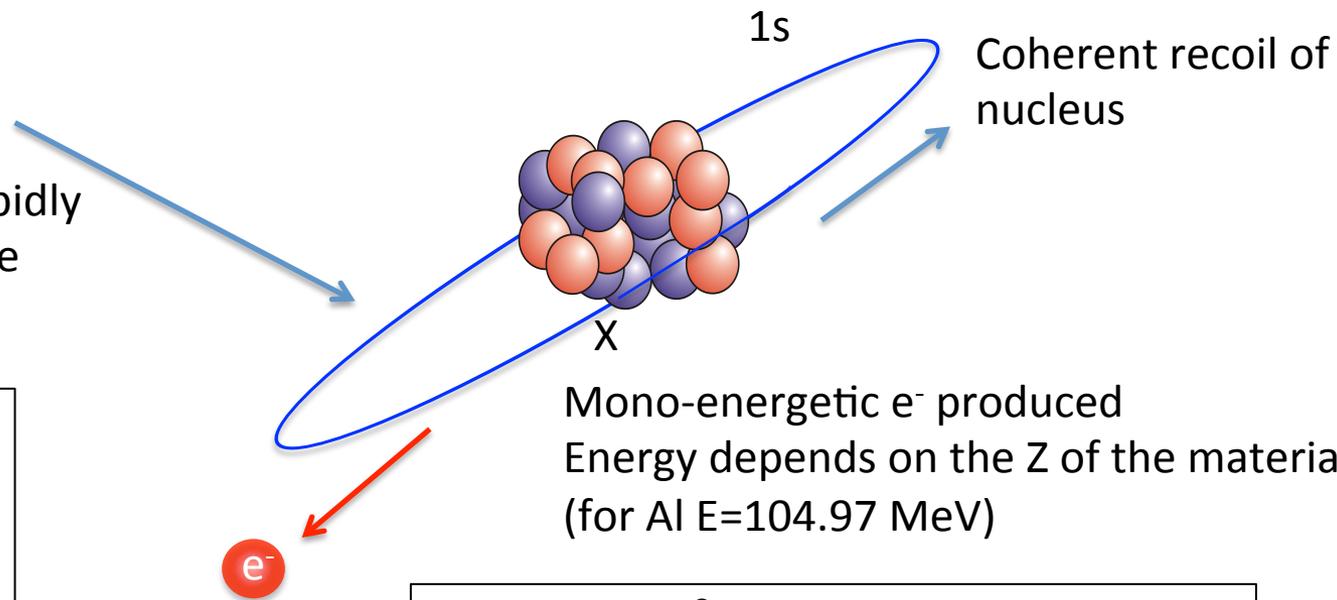
Mono-energetic  $e^-$  produced  
 Energy depends on the Z of the material  
 (for Al  $E=104.97$  MeV)

$$E_{e^-} = m_{\mu}c^2 - BE_{\mu}(Z) - RE(A)$$



# $\mu \rightarrow e$ conversion

Muons are stopped and rapidly (ps) cascade to the 1s state producing X-Rays.



Transition Energy for Al	
3d $\rightarrow$ 2p	66 keV
2p $\rightarrow$ 1s	356 keV
3d $\rightarrow$ 1s	423 keV
4p $\rightarrow$ 1s	446 keV

Mono-energetic  $e^-$  produced  
Energy depends on the Z of the material  
(for Al  $E=104.97$  MeV)

$$E_{e^-} = m_{\mu}c^2 - BE_{\mu}(Z) - RE(A)$$

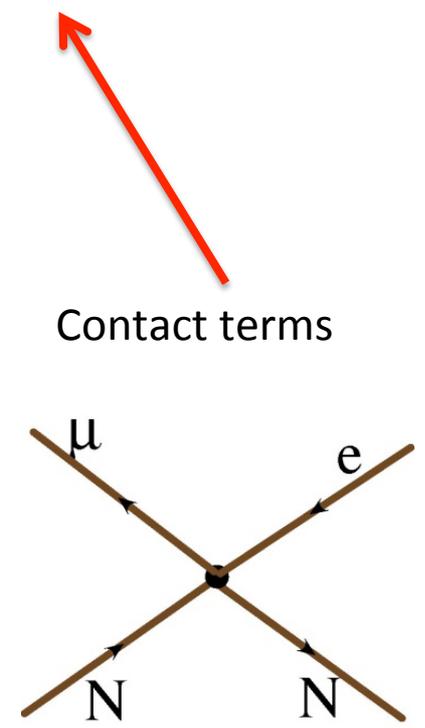
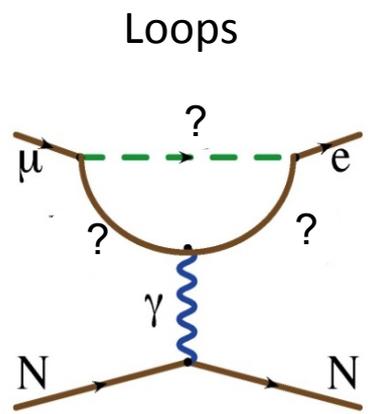
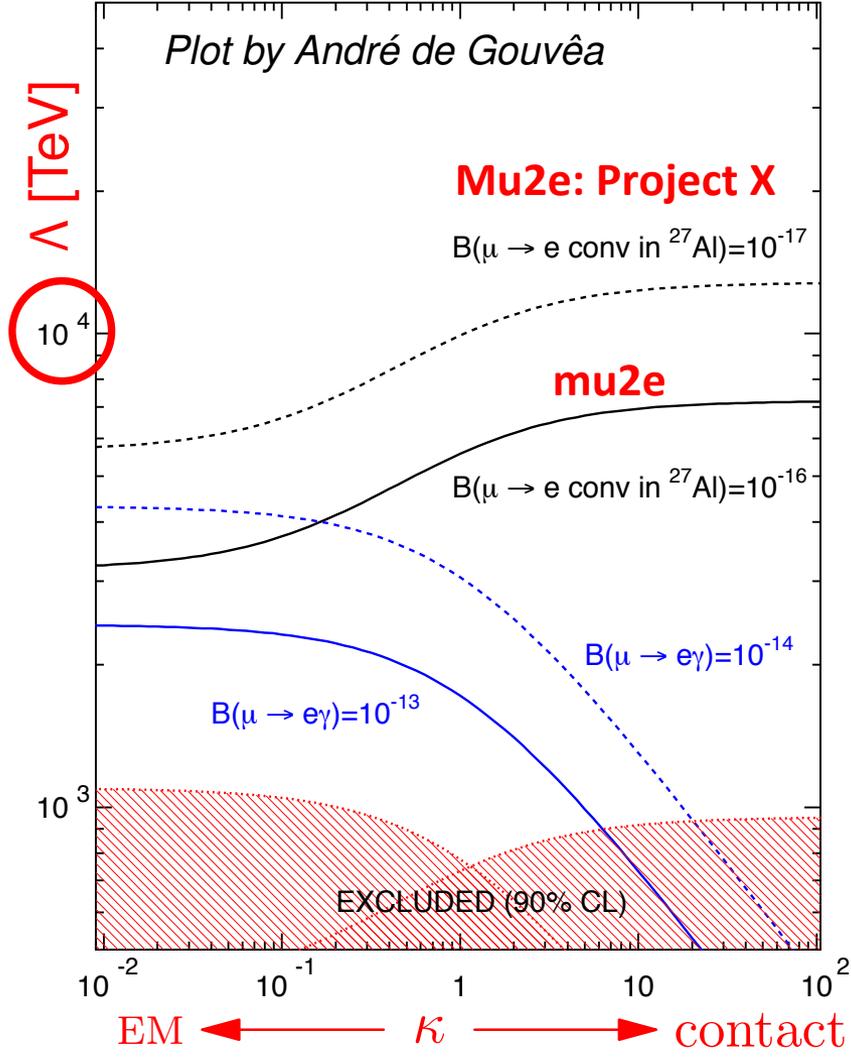
$$R_{\mu e} = \frac{\mu^- + A(Z, N) \rightarrow e^- + A(Z, N)}{\mu^- + A(Z, N) \rightarrow \text{all muon nuclear captures}}$$

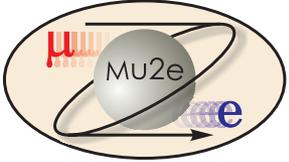


# Mu2e Sensitivity

$$\mathcal{L}_{\text{CLFV}} = \frac{m_\mu}{(\kappa + 1)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1 + \kappa)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L (\bar{u}_L \gamma_\mu u_L + \bar{d}_L \gamma_\mu d_L)$$

$\Lambda$ : Energy scale  
 $\kappa$ : Importance of contact term

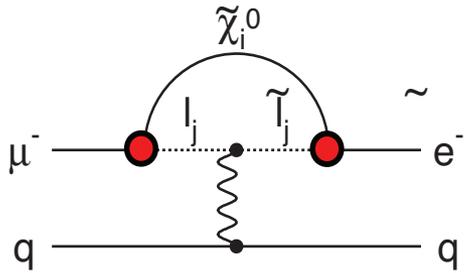




# Contributions to $\mu e$ conversion

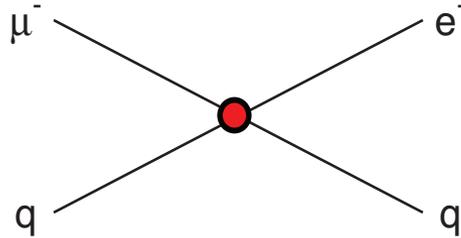
## Supersymmetry

Rate  $\sim 10^{-15}$



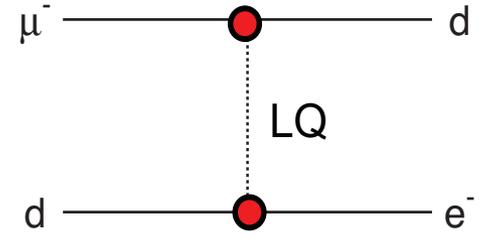
## Compositeness

$\Lambda_c \sim 3000$  TeV



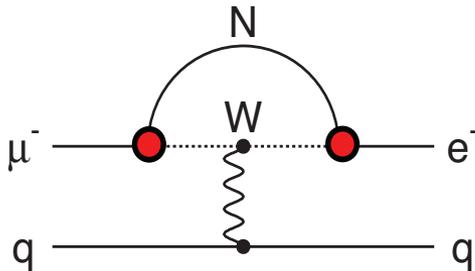
## Leptoquark

$M_{LQ} = 3000 (\lambda_{\mu d} \lambda_{e d})^{1/2} \text{ TeV}/c^2$



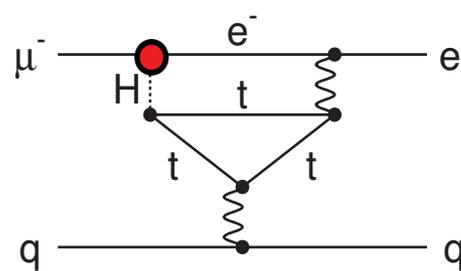
## Heavy Neutrinos

$|U_{\mu N} U_{e N}|^2 \sim 8 \times 10^{-13}$



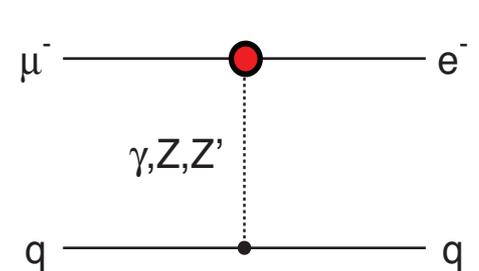
## Second Higgs Doublet

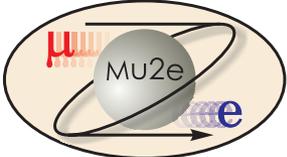
$g(H_{\mu e}) \sim 10^{-4} g(H_{\mu\mu})$



## Heavy $Z'$ Anomal. Z Coupling

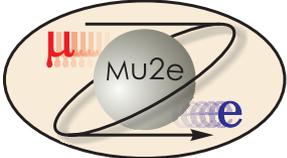
$M_{Z'} = 3000 \text{ TeV}/c^2$



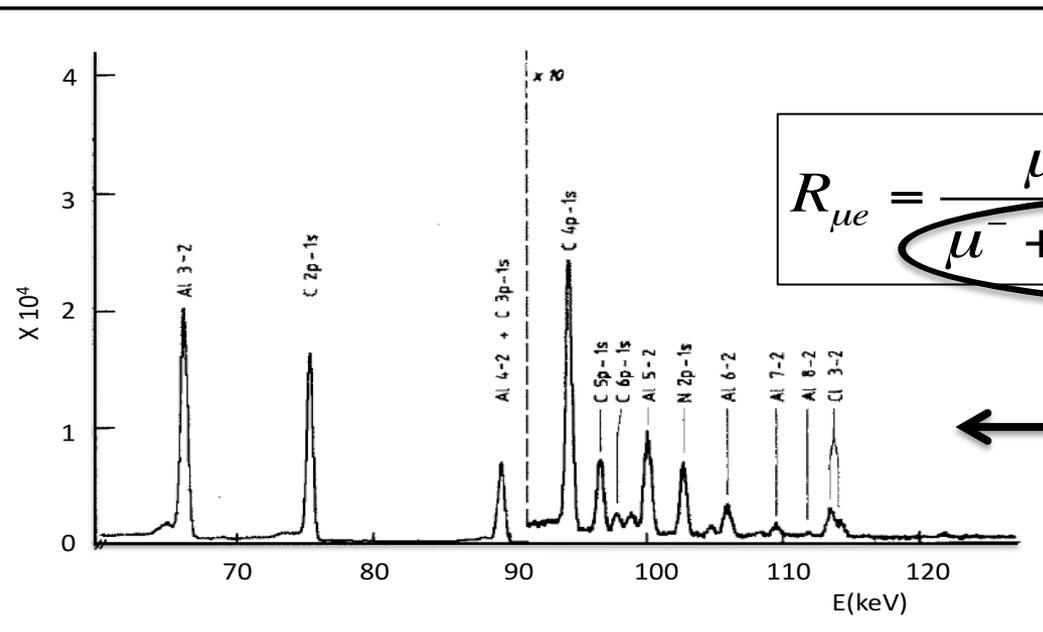


# Measuring $\mu e$ conversion

1. Generate pulsed beam of low momentum **negative muons**.
2. Stop muons in thin foils and form **muonic atoms**.
  - $\mu^-$  in aluminium:  $\tau_{1/2} = 864$  ns.
3. Wait for **prompt background** to **decay**.
  - Need suitable beam repetition rate.
4. Then measure the electron spectrum.
  - The signal: mono-energetic electrons at 105 MeV



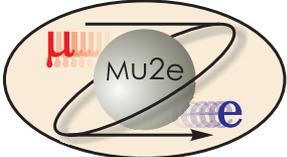
# Measuring $\mu e$ conversion



$$R_{\mu e} = \frac{\mu^- + A(Z, N) \rightarrow e^- + A(Z, N)}{\mu^- + A(Z, N) \rightarrow \text{all muon captures}}$$

- With sensitivity to R at 90% C.L.  $< 5 \times 10^{-17}$
- 4 orders of magnitude better than current limits
  - Need more than  $10^{17}$  stopped muons!
  - --  $3.6 \times 10^{20}$  protons on target (3 year run)

**Need to reduce any background that could affect our sensitivity.**

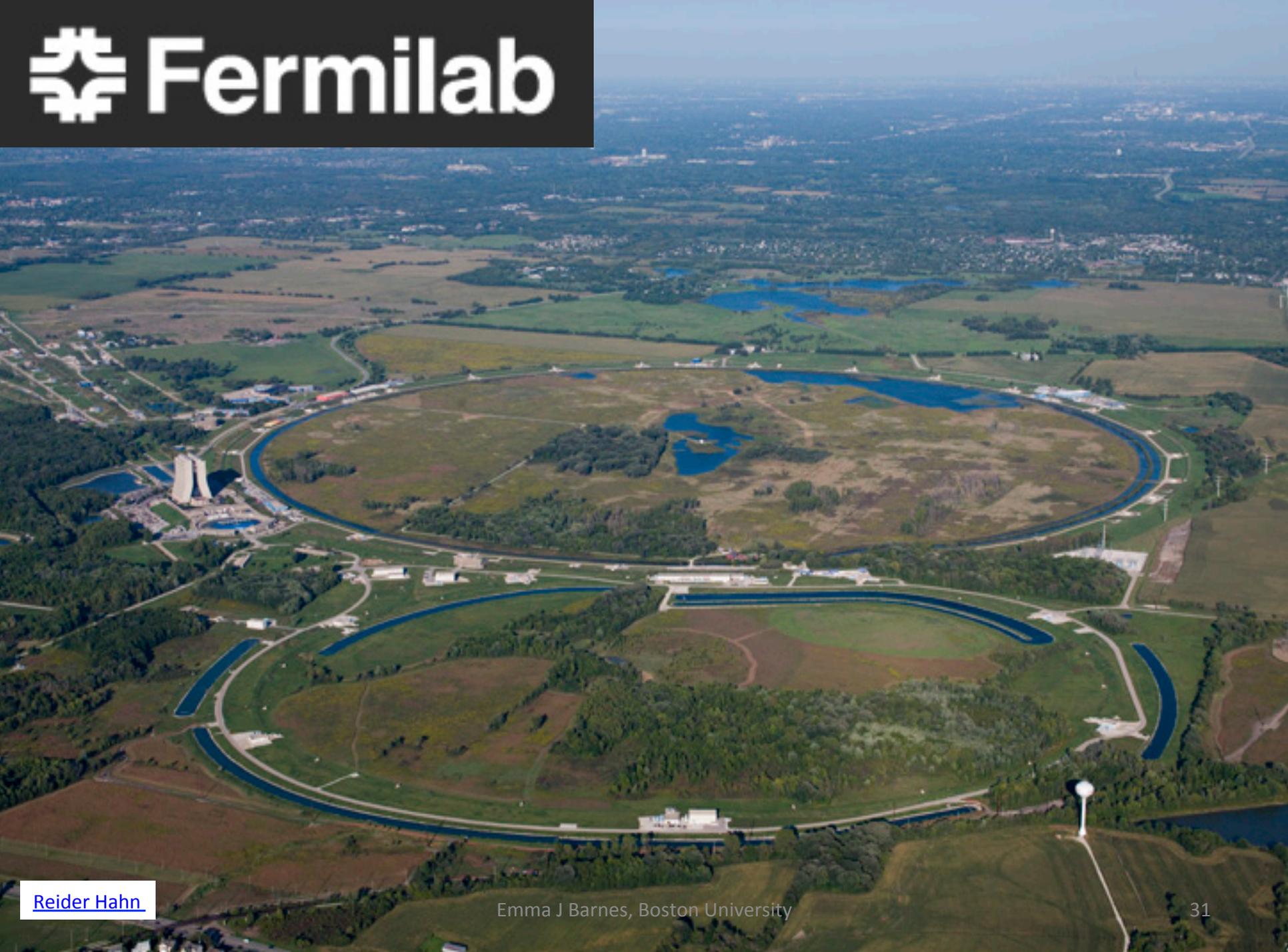


# Needle in a haystack?

- Muon decay in orbit (DIOs)
- Radiative muon/pion capture
$$\pi^- N \rightarrow \gamma N^* \quad \pi^- N \rightarrow e^+ e^- N^*$$
- Long transit time backgrounds (pbars, pions)
- Cosmic ray induced electrons
- Muon/pion decay in flight
- Beam electrons that can scatter into the target

**Need to reduce any background that could affect our sensitivity.**

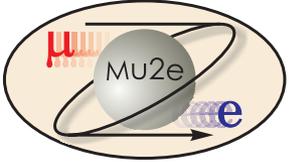
**Mu2e has been designed with this in mind.**







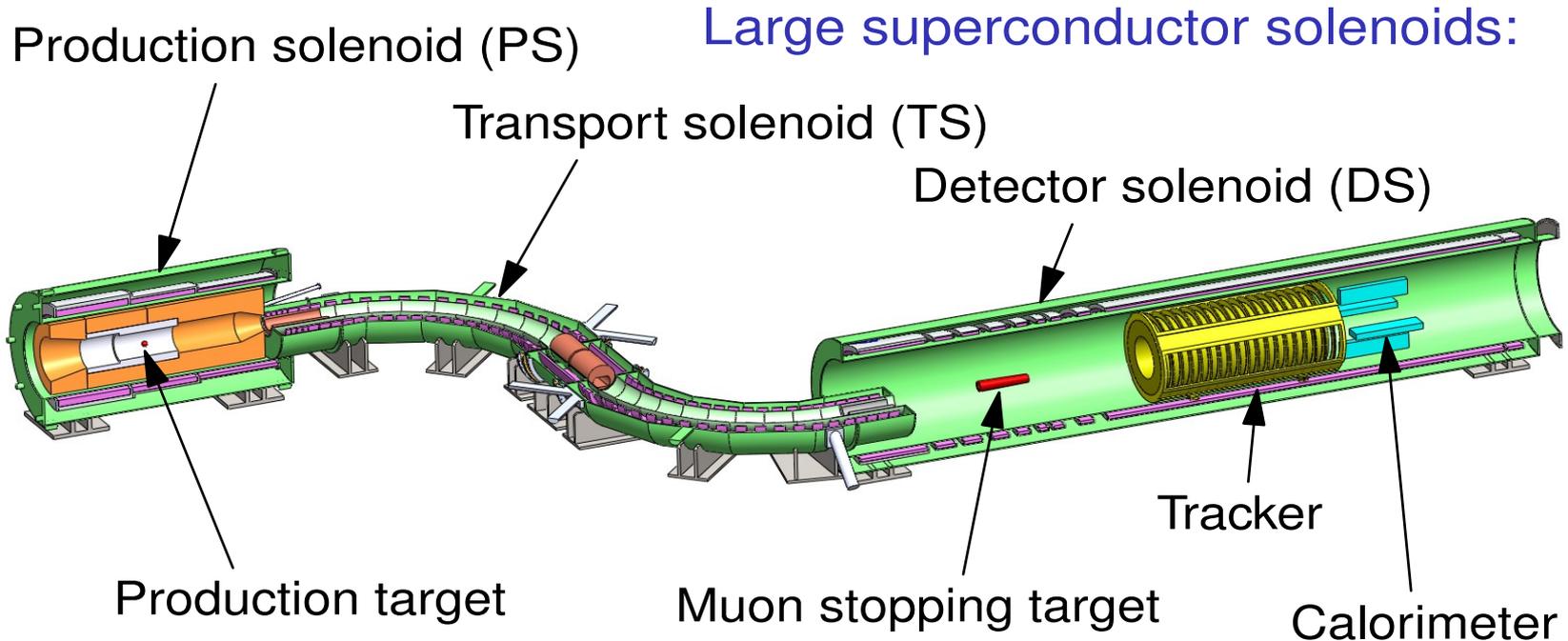




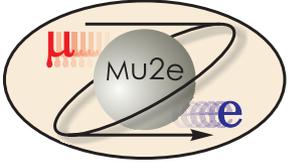
# The mu2e Detector System

**Transport:** S-curve removes line of sight background and also sign-selects.

**Detector:** Stopping Target, Tracking and Calorimeter

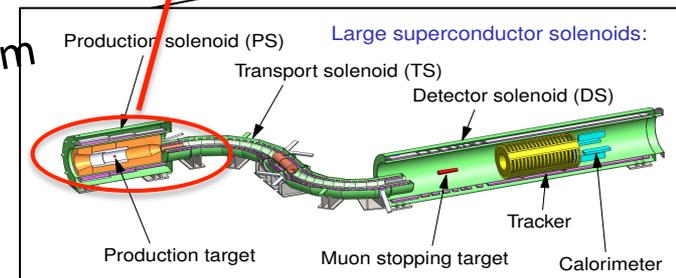
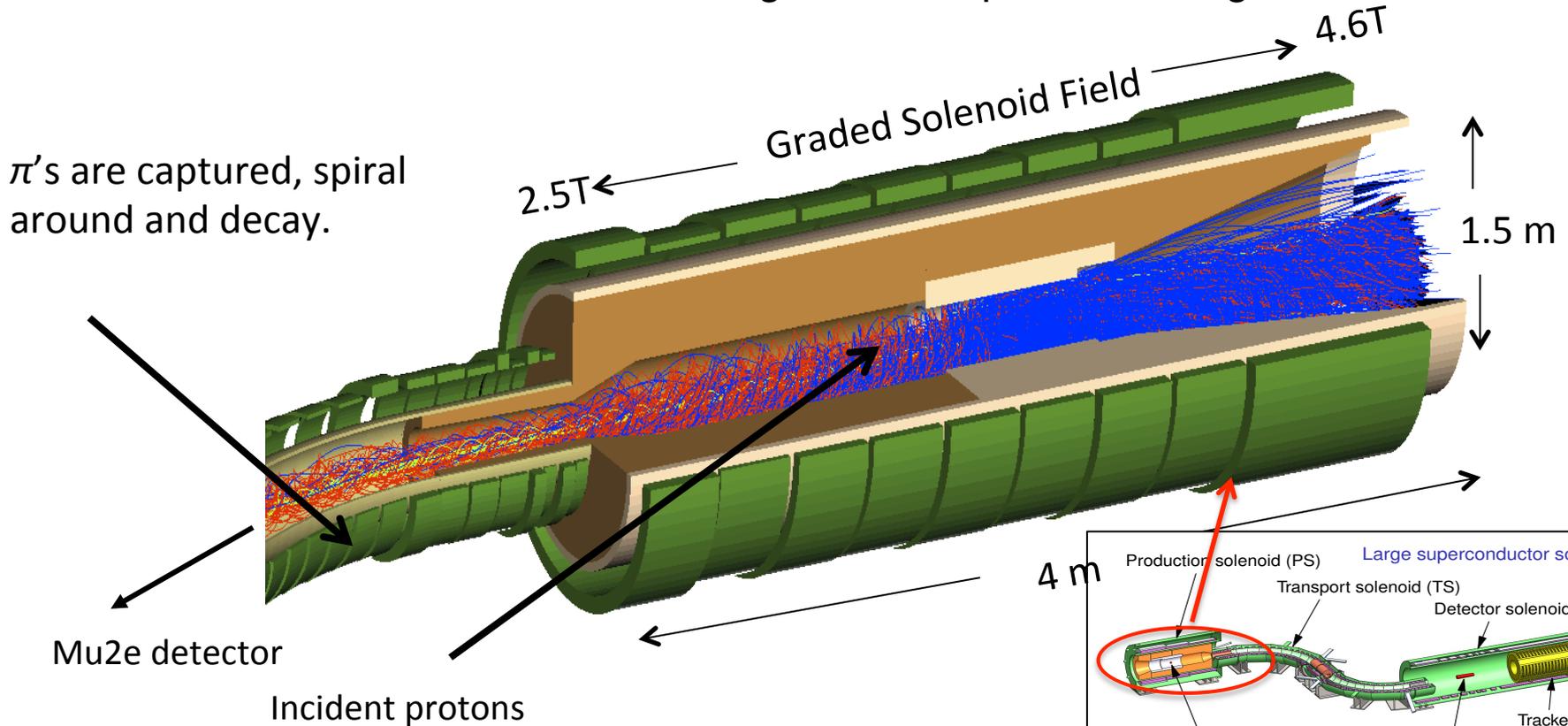


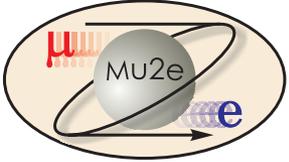
**Production:** 8 GeV protons enter the PS and encounter a tungsten target.  
Up to 8 kW beam power ( $6 \times 10^{12}$  Protons/s)  
 $\pi$ 's are then produced which decay into accepted  $\mu$ 's



# The production Solenoid (PS)

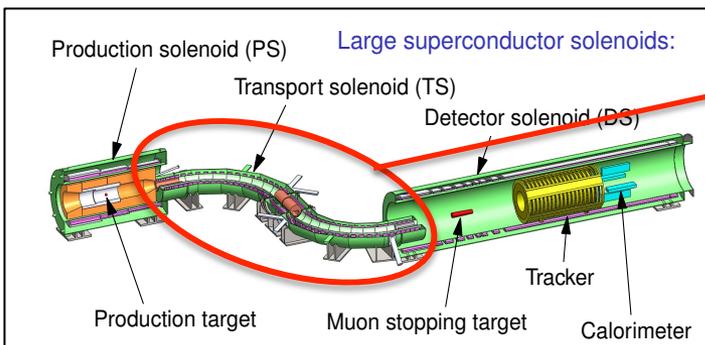
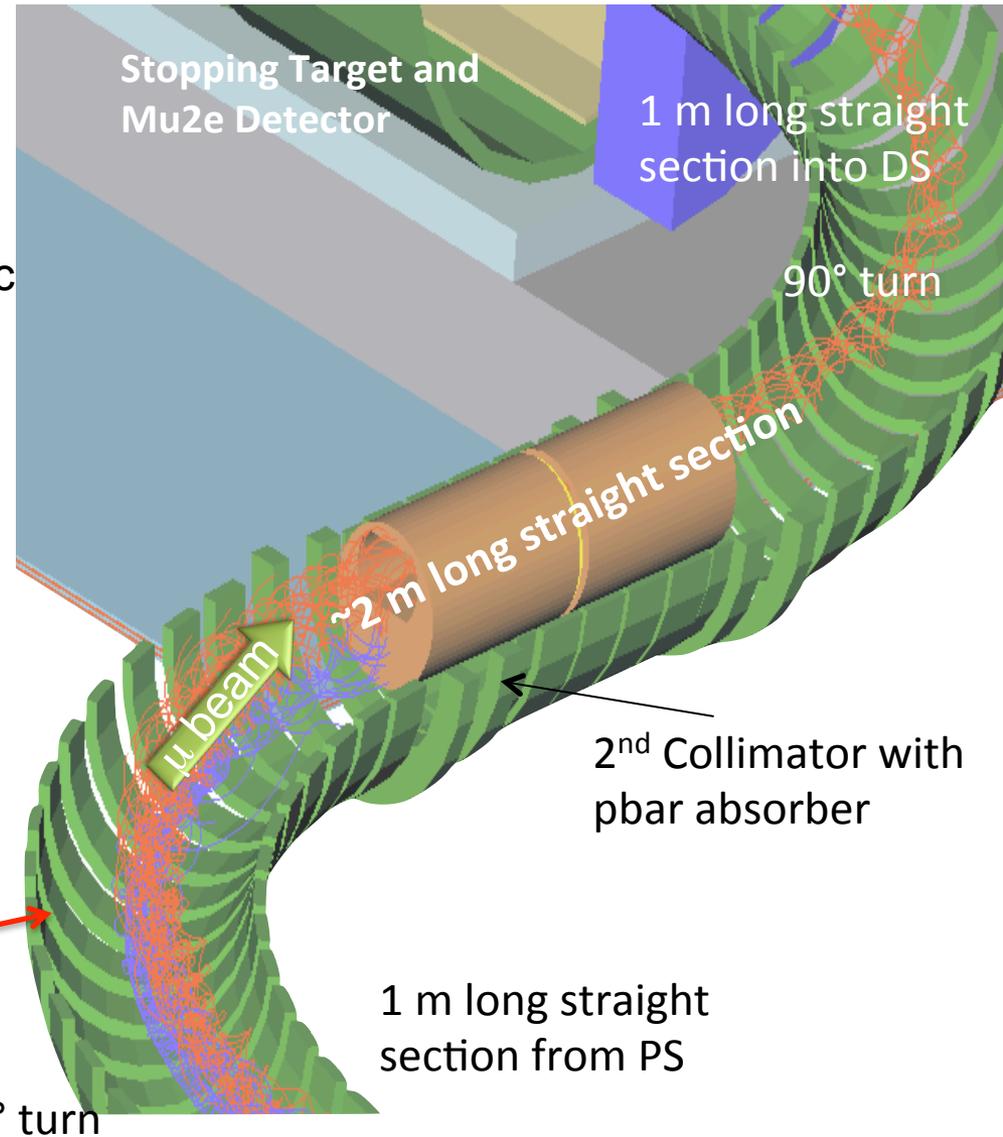
- 8 GeV protons strike a tungsten production target near the center of the Production Solenoid.
- The PS is a high field magnet with a graded (4.6 T  $\rightarrow$  2.5 T) solenoidal field.
- The PS will be used to capture pions which can then decay into muons.
- A heat and radiation shield, constructed from bronze, will line the inside of the PS to limit the heat load and radiation damage to the superconducting cable.

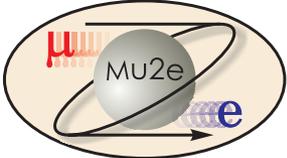




# The Transport Solenoid (TS)

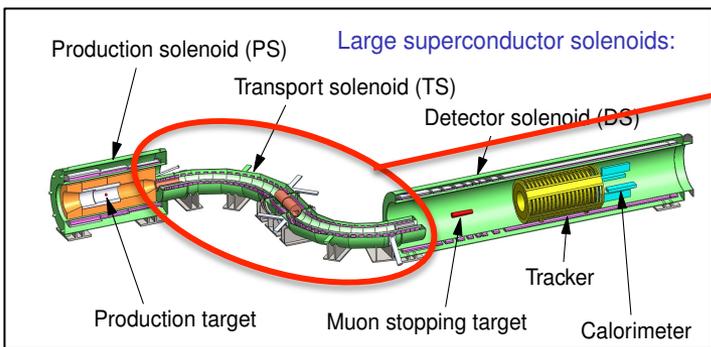
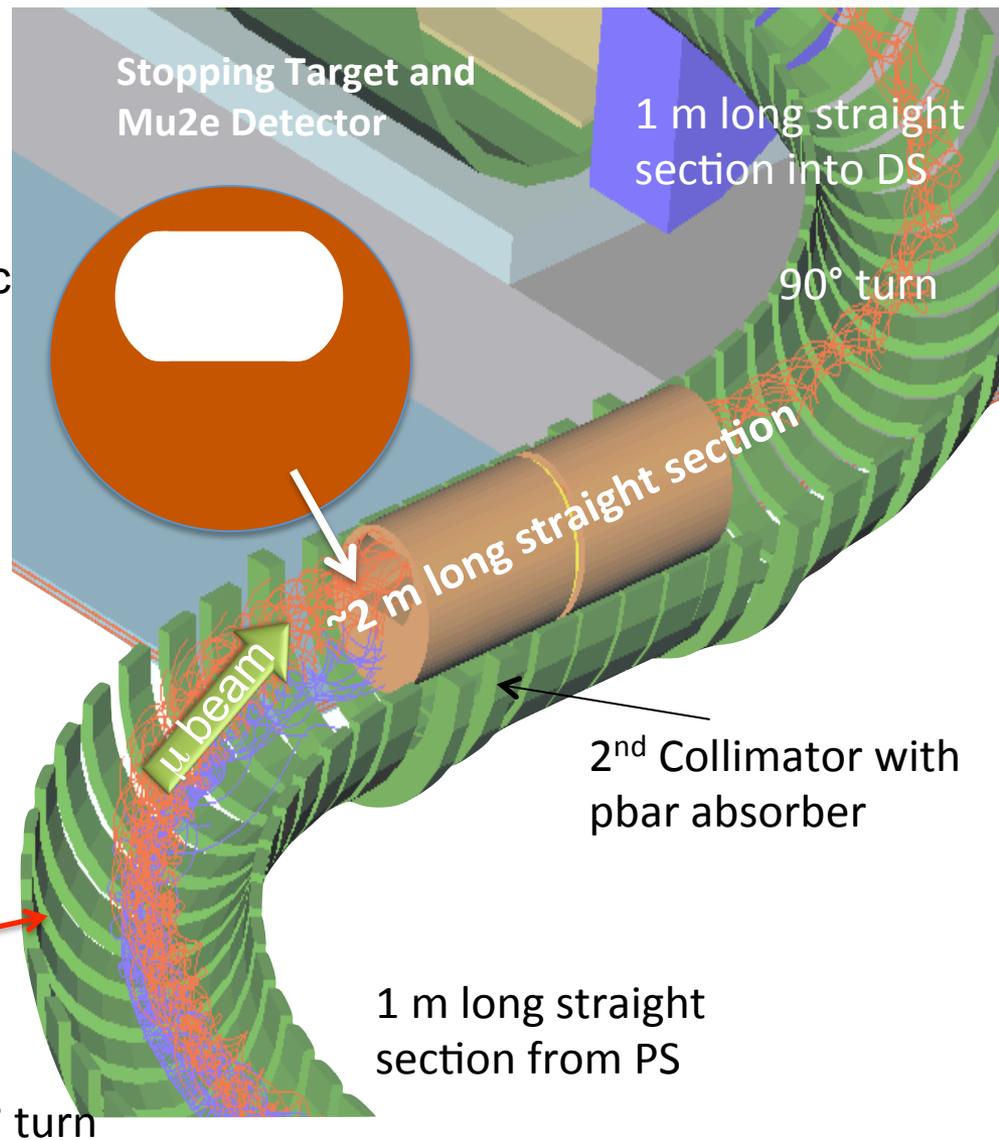
- S-shaped to reduced any line-of-sight photons and neutrons.
- It consists of a set of superconducting solenoids and toroids that form a magnetic channel that transmits low energy negatively charged muons from the PS.
- High energy negatively charged particles, positively charged particles and line-of-sight neutral particles will nearly all be eliminated by the two 90° bends combined with a series of absorbers and collimators.





# The Transport Solenoid (TS)

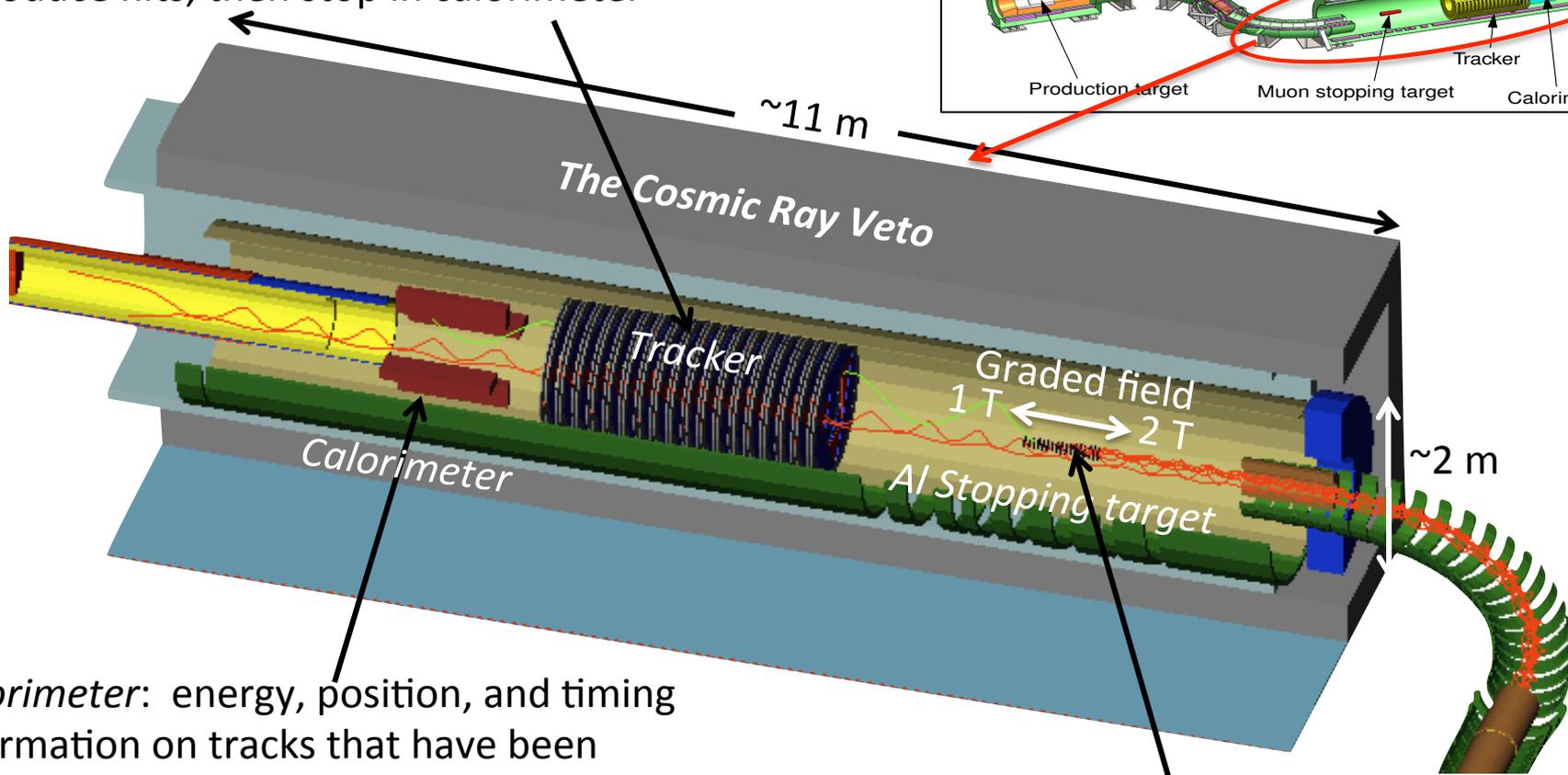
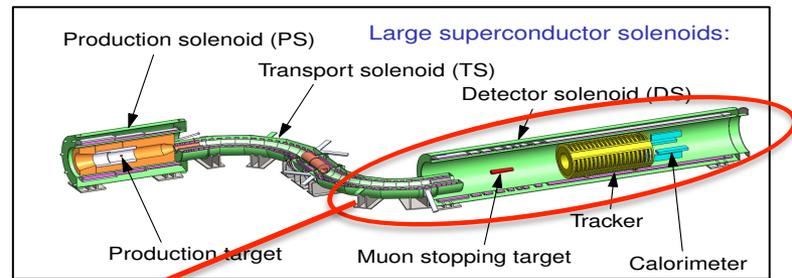
- S-shaped to reduced any line-of-sight photons and neutrons.
- It consists of a set of superconducting solenoids and toroids that form a magnetic channel that transmits low energy negatively charged muons from the PS.
- High energy negatively charged particles, positively charged particles and line-of-sight neutral particles will nearly all be eliminated by the two 90° bends combined with a series of absorbers and collimators.





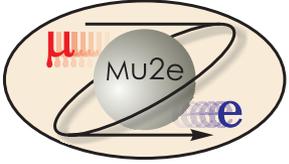
# The Detector Solenoid (DS)

Signal events pass *through* tracker and produce hits, then stop in calorimeter



**Calorimeter:** energy, position, and timing information on tracks that have been reconstructed by the tracker.  
1936 LYSO crystals arranged in four vanes

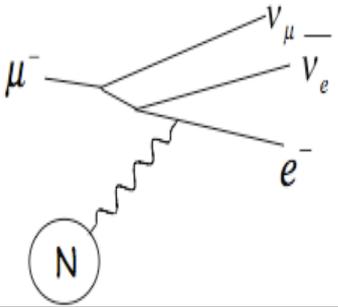
17 thin Al disks. Radii decreases from 8.3 cm upstream to 6.53 cm downstream.



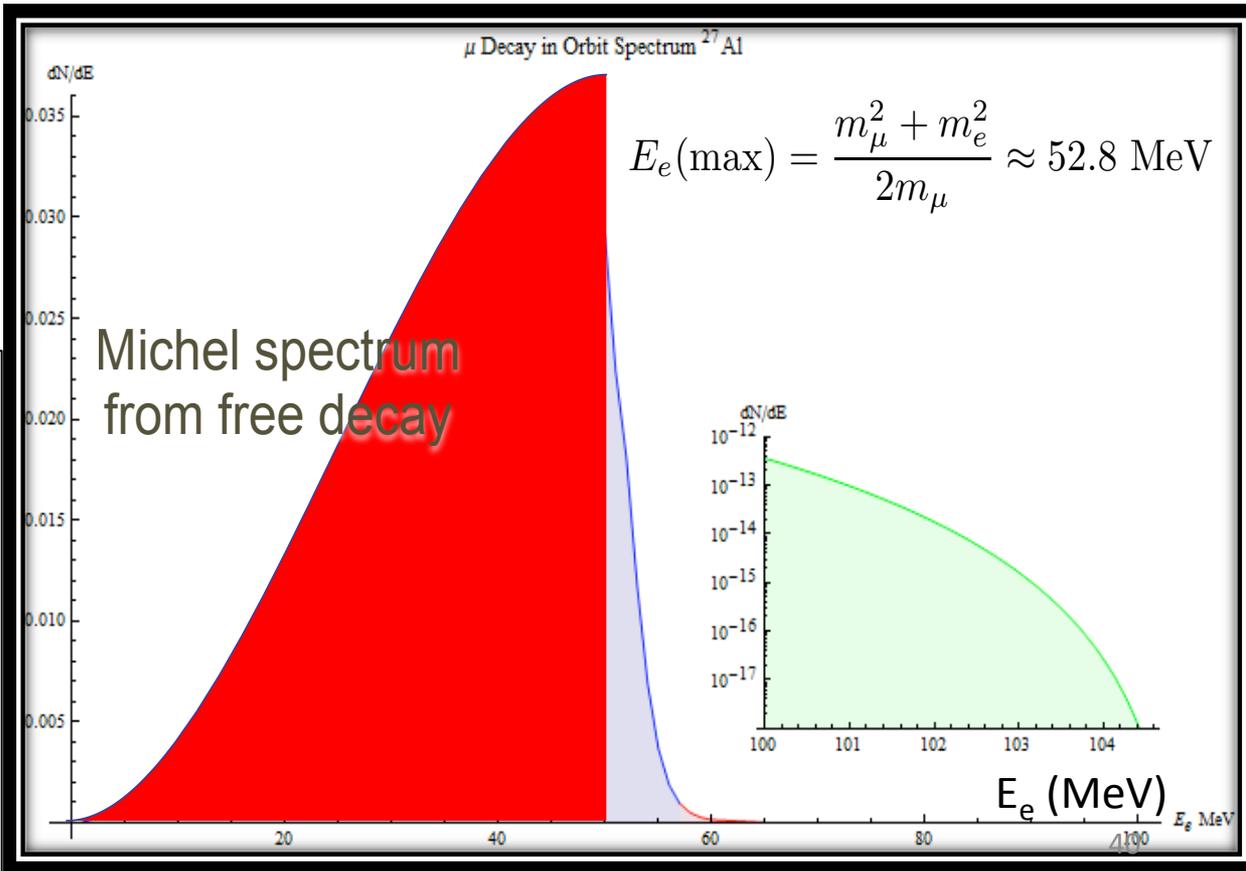
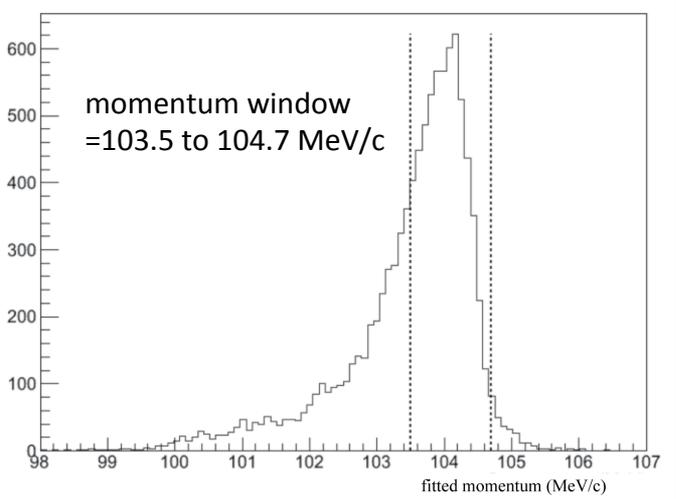
# Decay In Orbit (DIO)

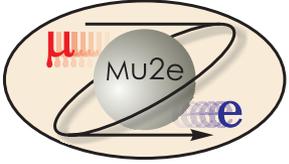


- DIOs contribute ~50% background in the signal energy region.
- Most DIO e<sup>-</sup>s have energies below 52.8 MeV, the kinematic limit for decay of a free muon.
- For decay off a bound muon the outgoing e<sup>-</sup> can recoil off the nucleus.
- If the neutrinos are at rest the e<sup>-</sup> can have exactly the conversion energy E<sub>CE</sub>=104.97 MeV.
- Small rate  $\propto (E_{ce} - E_{electron})^5$
- $\sim 3 \times 10^{-13}$  e<sup>-</sup>s >100 MeV



Track fit momentum

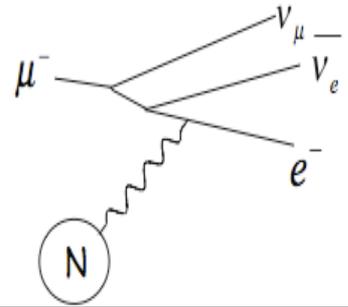




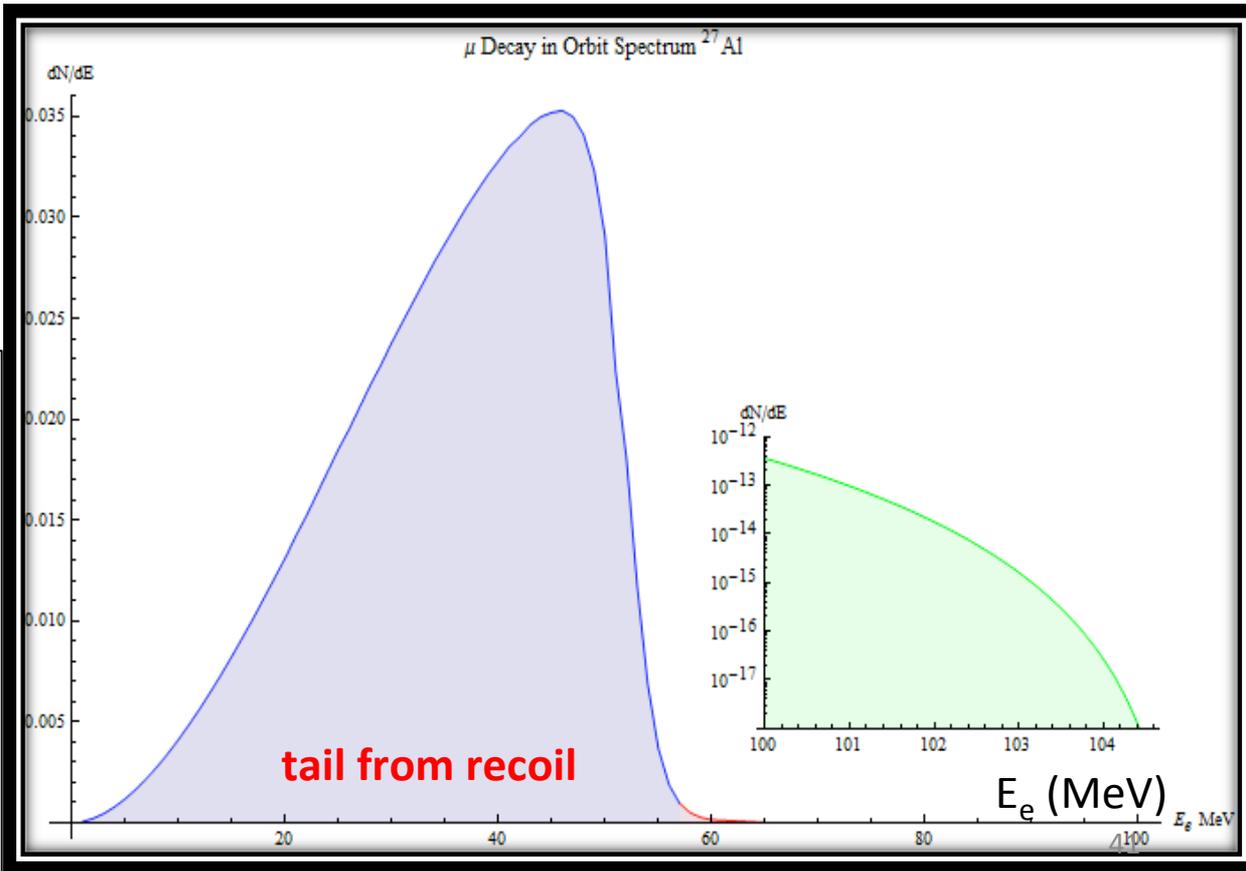
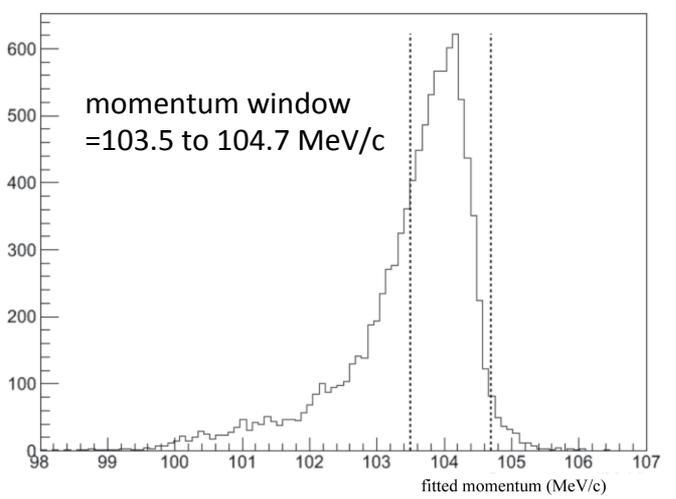
# Decay In Orbit (DIO)

$$\mu^- + A(Z, N) \rightarrow A(Z, N) + e^- + \bar{\nu}_e + \nu_\mu$$

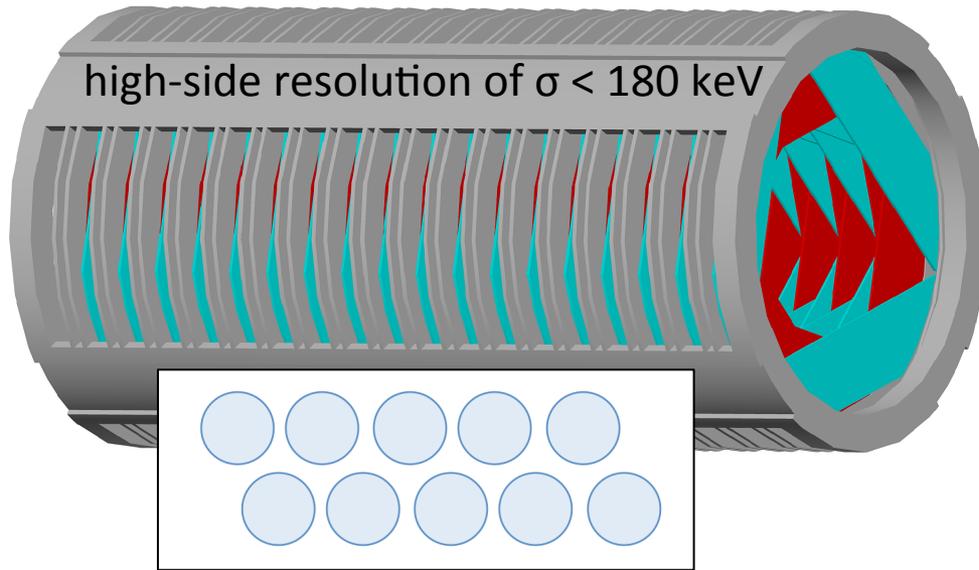
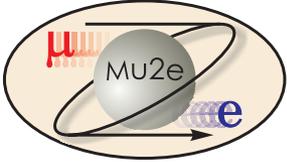
- DIOs contribute ~50% background in the signal energy region.
- Most DIO e<sup>-</sup>s have energies below 52.8 MeV, the kinematic limit for for decay of a free muon.
- For decay off a bound muon the outgoing e<sup>-</sup> can recoil off the nucleus.
- If the neutrinos are at rest the e<sup>-</sup> can have exactly the conversion energy E<sub>CE</sub>=104.97 MeV.
- Small rate  $\propto (E_{ce} - E_{electron})^5$
- $\sim 3 \times 10^{-13}$  e<sup>-</sup>s >100 MeV



Track fit momentum

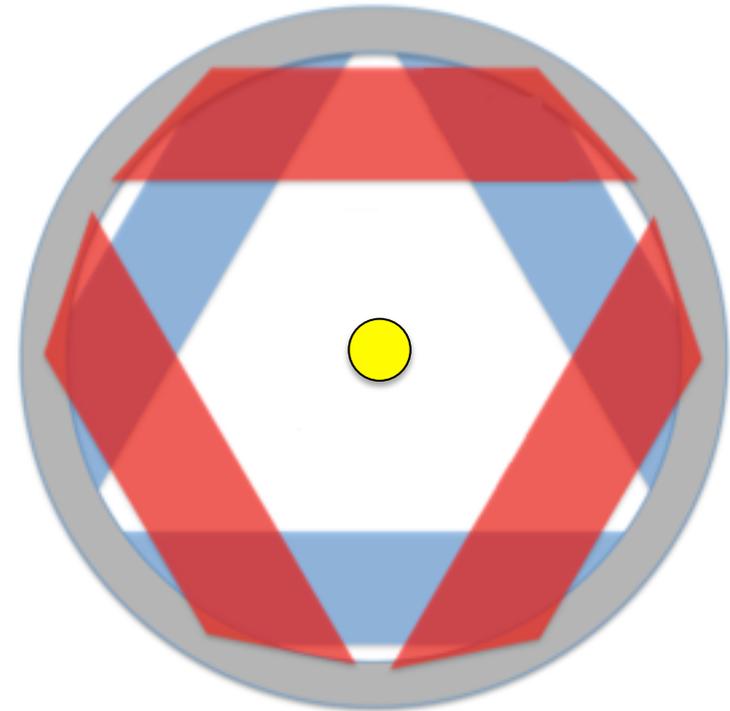


# The Tracker

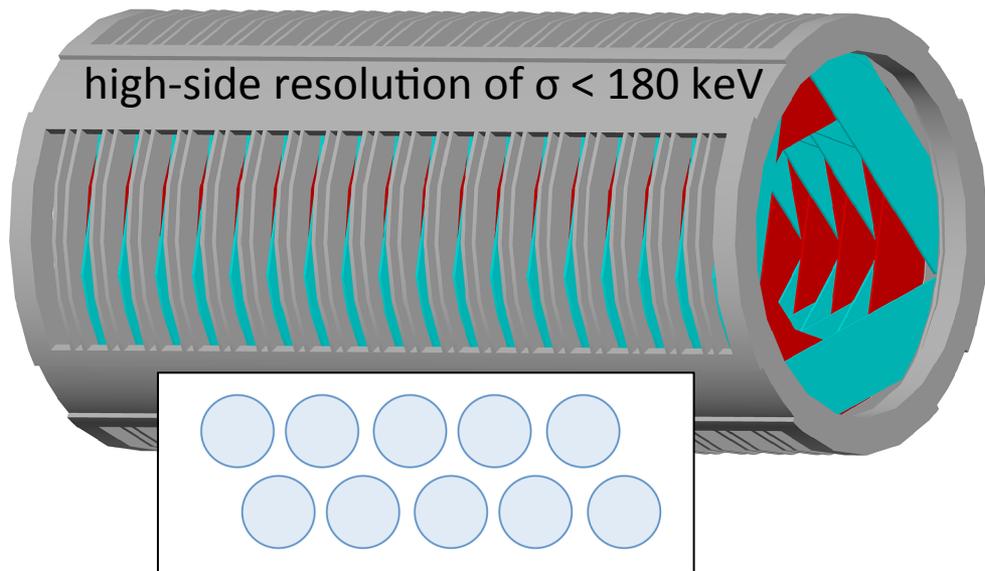
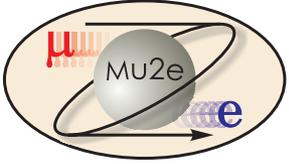


20  $\mu\text{m}$  sense wire inside a 5 mm diameter tube made of 15  $\mu\text{m}$  thick metalized Mylar<sup>®</sup>.  
 ~22,000 straws distributed into 18 measurement stations across a ~3 m length.  
 Planes are constructed from two layers of straws and are operated in a vacuum.

- The tracker is designed to intercept only a small fraction of the significant flux of electrons from muon decays-in-orbit (DIO) (energies  $< 60\text{MeV}$ ).

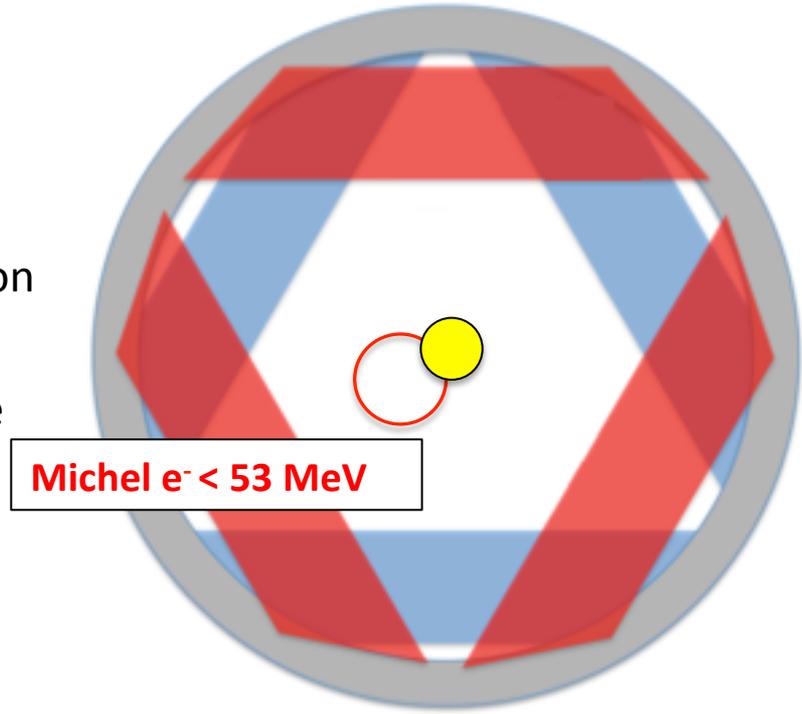


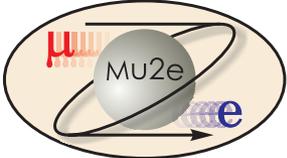
# The Tracker



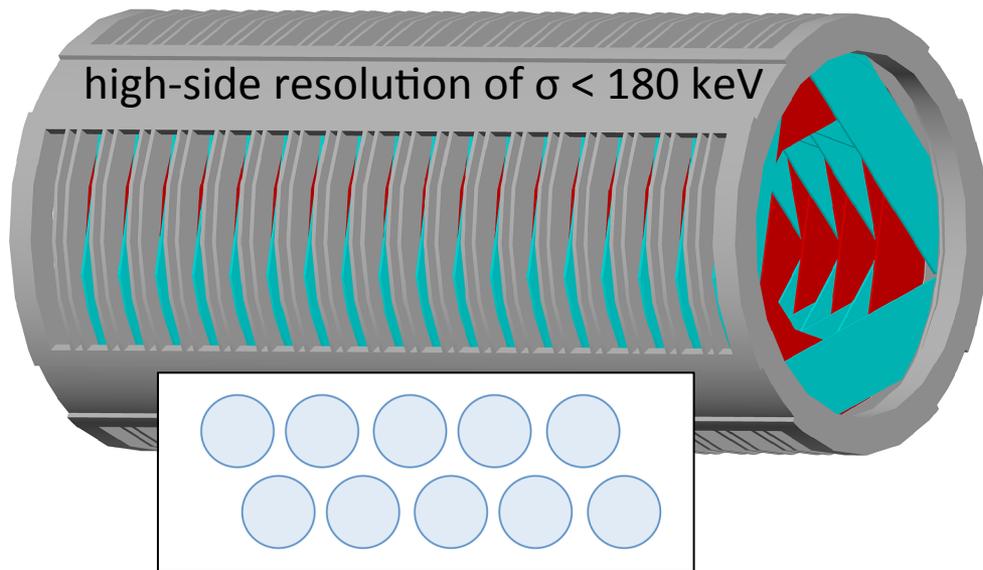
20  $\mu\text{m}$  sense wire inside a 5 mm diameter tube made of 15  $\mu\text{m}$  thick metalized Mylar<sup>®</sup>.  
 ~22,000 straws distributed into 18 measurement stations across a ~3 m length.  
 Planes are constructed from two layers of straws and are operated in a vacuum.

- The tracker is designed to intercept only a small fraction of the significant flux of electrons from muon decays-in-orbit (DIO) (energies < 60MeV).
- $e^-$  with energies < 53 MeV will curl in the field of the DS and pass unobstructed through the hole in the center of the tracker.



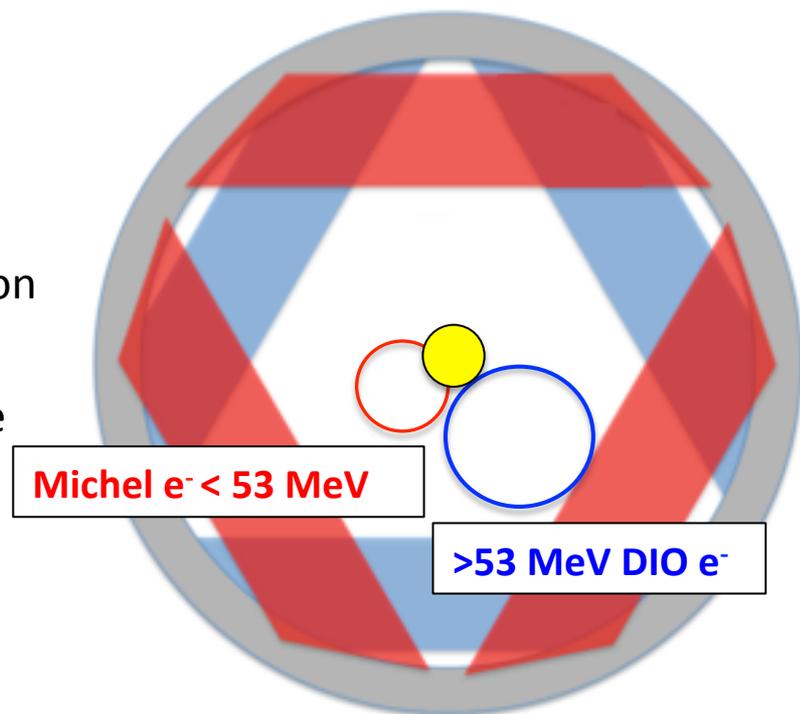


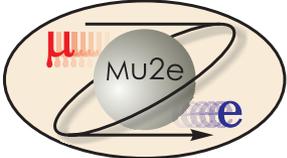
# The Tracker



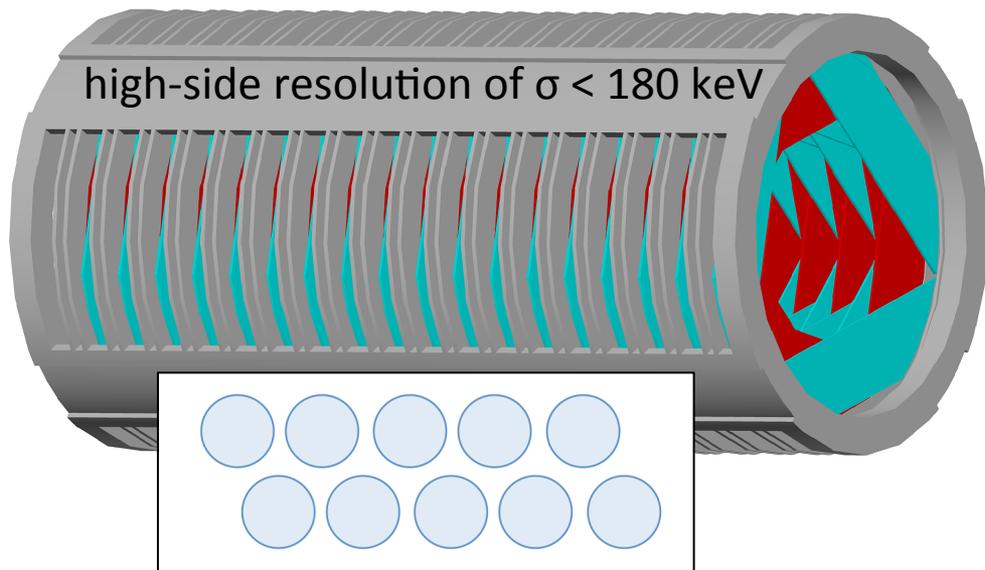
20  $\mu\text{m}$  sense wire inside a 5 mm diameter tube made of 15  $\mu\text{m}$  thick metalized Mylar<sup>®</sup>.  
~22,000 straws distributed into 18 measurement stations across a ~3 m length. Planes are constructed from two layers of straws and are operated in a vacuum.

- The tracker is designed to intercept only a small fraction of the significant flux of electrons from muon decays-in-orbit (DIO) (energies  $< 60\text{MeV}$ ).
- $e^-$  with energies  $< 53\text{ MeV}$  will curl in the field of the DS and pass unobstructed through the hole in the center of the tracker.
- $e^-$  with energies  $> 53\text{ MeV}$  (~3% total rate) will be observed in the tracker.



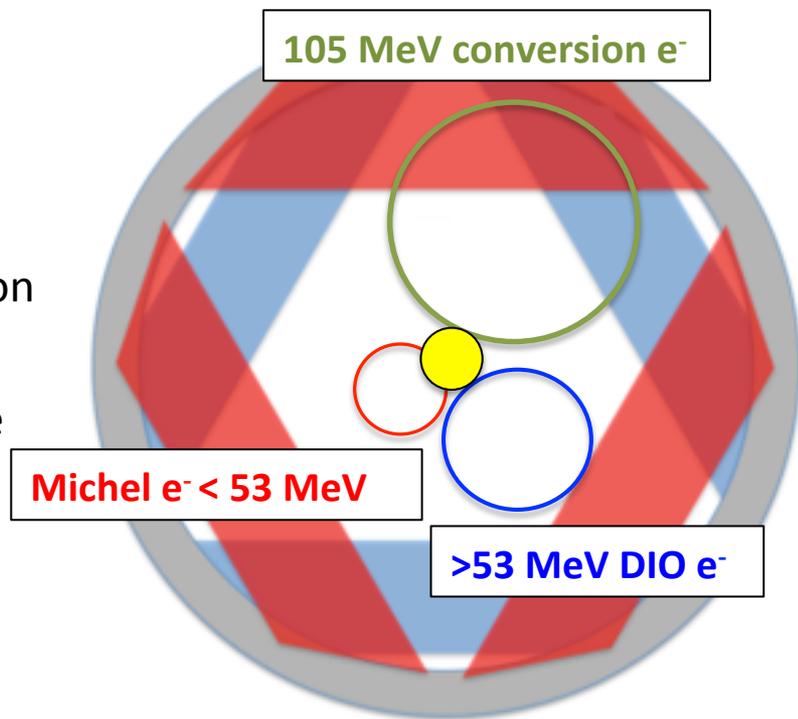


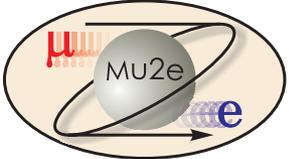
# The Tracker



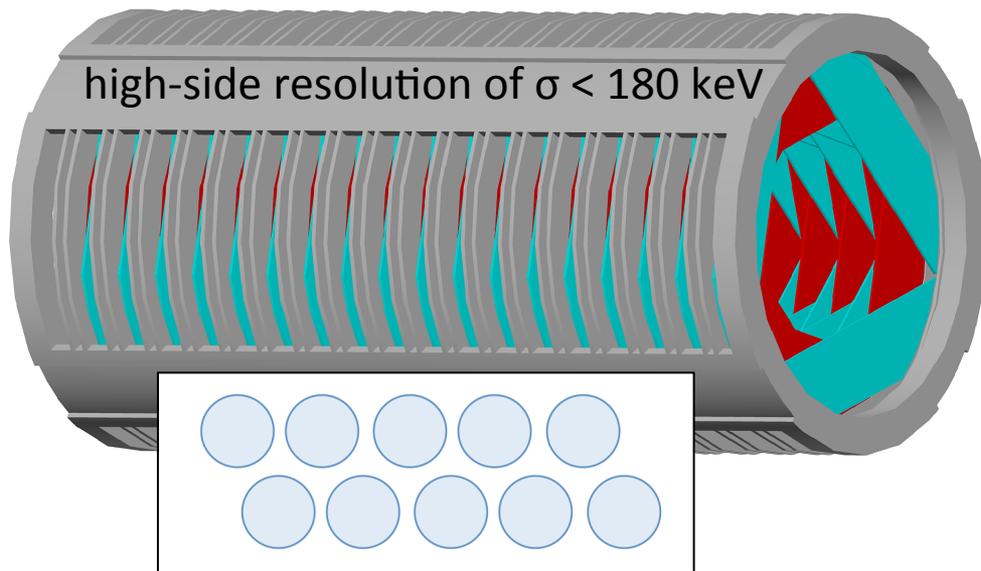
20  $\mu\text{m}$  sense wire inside a 5 mm diameter tube made of 15  $\mu\text{m}$  thick metalized Mylar<sup>®</sup>.  
~22,000 straws distributed into 18 measurement stations across a ~3 m length.  
Planes are constructed from two layers of straws and are operated in a vacuum.

- The tracker is designed to intercept only a small fraction of the significant flux of electrons from muon decays-in-orbit (DIO) (energies  $< 60\text{MeV}$ ).
- $e^-$  with energies  $< 53 \text{ MeV}$  will curl in the field of the DS and pass unobstructed through the hole in the center of the tracker.
- $e^-$  with energies  $> 53 \text{ MeV}$  (~3% total rate) will be observed in the tracker.

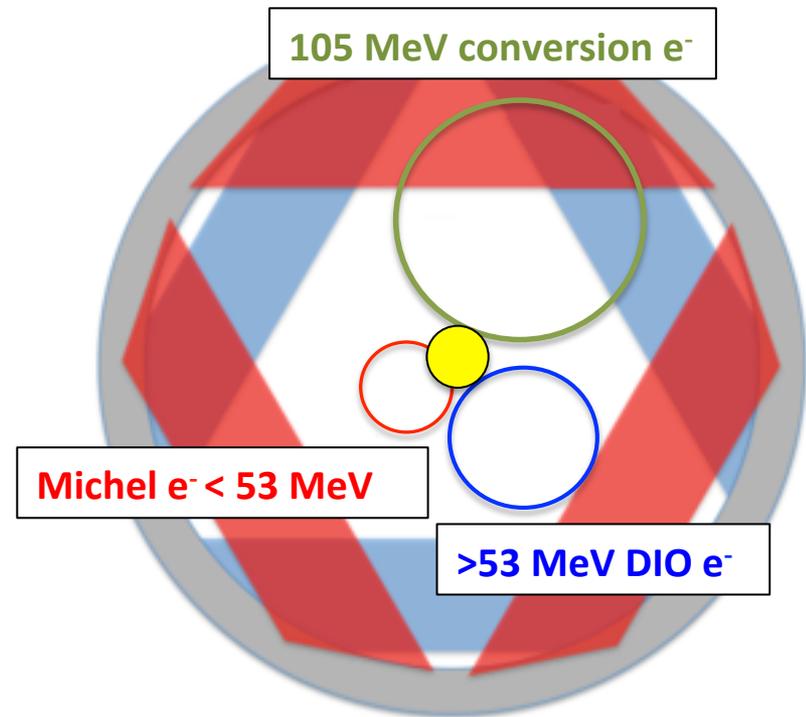
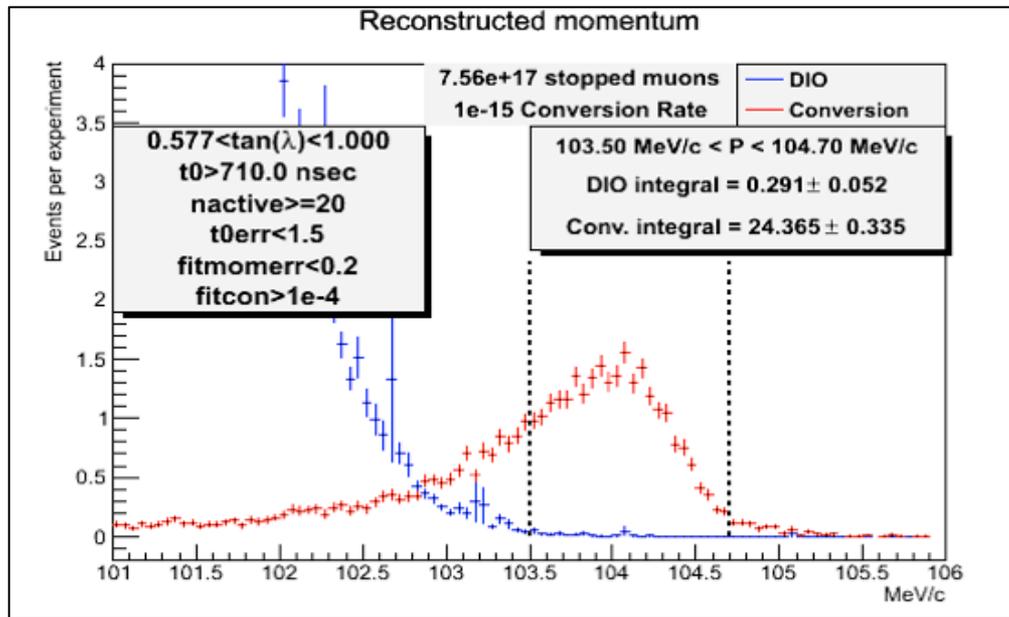


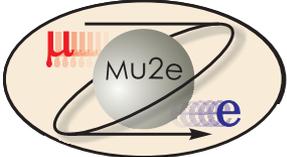


# The Tracker

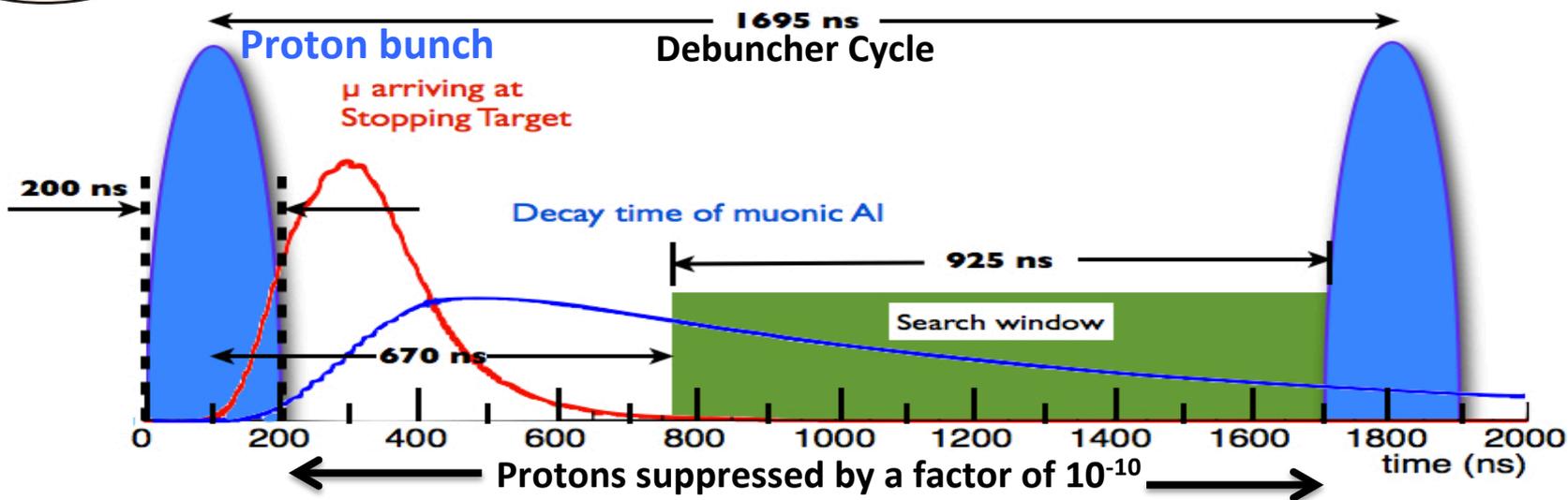


20  $\mu\text{m}$  sense wire inside a 5 mm diameter tube made of 15  $\mu\text{m}$  thick metalized Mylar<sup>®</sup>.  
 ~22,000 straws distributed into 18 measurement stations across a ~3 m length.  
 Planes are constructed from two layers of straws and are operated in a vacuum.





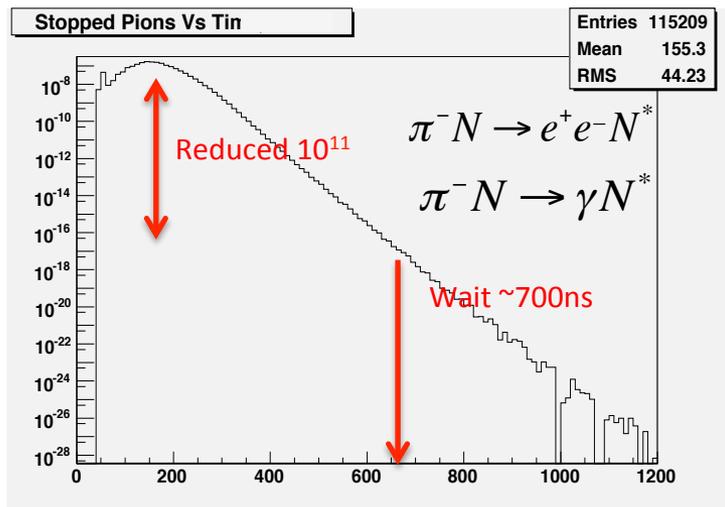
# Pulsed beam structure

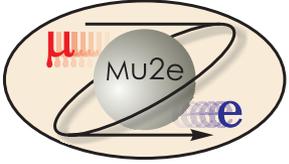


- Delivery ring supplies a single circulating bunch where protons are resonantly extracted to the beamline.
- Pulsed beam to Mu2e every cyclotron period of 1695 ns (2x muons lifetime) for 8 GeV protons.
- Muon lifetime in Al is long (864 ns) the loss of muons is acceptable if the time between pulses is not much longer than the muon lifetime and one simply waits for the pions to decay (21 ns).

**Extinction level of  $10^{-10}$  between bunches is crucial!**

$$\text{Extinction} = \frac{\text{number of protons striking the PT between spills}}{\text{number striking it during the spills.}}$$





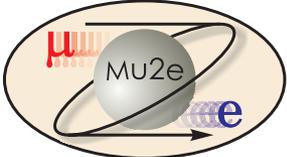
# Backgrounds and sensitivities

3 years of  $1.2 \times 10^{20}$  protons/year (8 kW beam power)

Background	Background Estimate
Muon decay-in-orbit	$0.22 \pm 0.06$
Cosmic Rays	$0.05 \pm 0.013$
Radiative Pion Capture	$0.03 \pm 0.007$
Pion decay In-Flight	$0.003 \pm 0.0015$
Muon decay In-Flight	$0.01 \pm 0.003$
Antiproton Induced	$0.10 \pm 0.05$
Beam electrons	$0.0006 \pm 0.0003$
Radiative muon capture	$< 2 \times 10^{-6}$
Total (Add in quadrature)	$0.41 \pm 0.08$

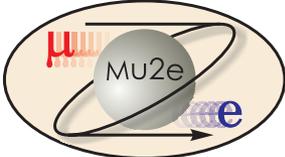
**GOAL  $< 2.4 \times 10^{-17}$  single event sensitivity**

**$R_{\mu e} \sim 10^{-15}$  40 signal events**



# Road Map

- Aim for a 4 orders of magnitude improvement compared to SINDRUM II which achieved  $R_{\mu e} < 7 \times 10^{-13}$ .
- CD-0 approval.
- Lehman review for CD-1 last week!
- Start taking data in 2019
- 3 year running time



# The Mu2e Collaboration



- Fermi National Accelerator Laboratory
- Lewis University
- University of Illinois, Urbana-Champaign
- Los Alamos National Laboratory
- University of Massachusetts, Amherst Muons, Inc.
- Northwestern University
- Northern Illinois University
- Rice University
- University of Houston
- University of Virginia
- University of Washington



- Istituto Nazionale di Fisica Nucleare Pisa Istituto Nazionale di Fisica Nucleare, Lecce
- Laboratori Nazionali di Frascati



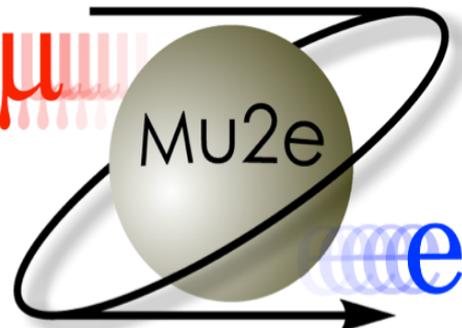
- Institute for Nuclear Research, Moscow
- Joint Institute for Nuclear Research, Dubna

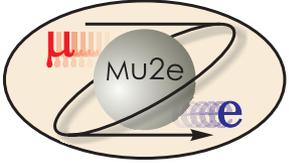
- Boston University
- Brookhaven National Laboratory
- University of California, Berkeley
- Lawrence Berkeley National Laboratory
- University of California, Irvine
- California Institute of Technology
- City University of New York
- Duke University

**~130 collaborators**

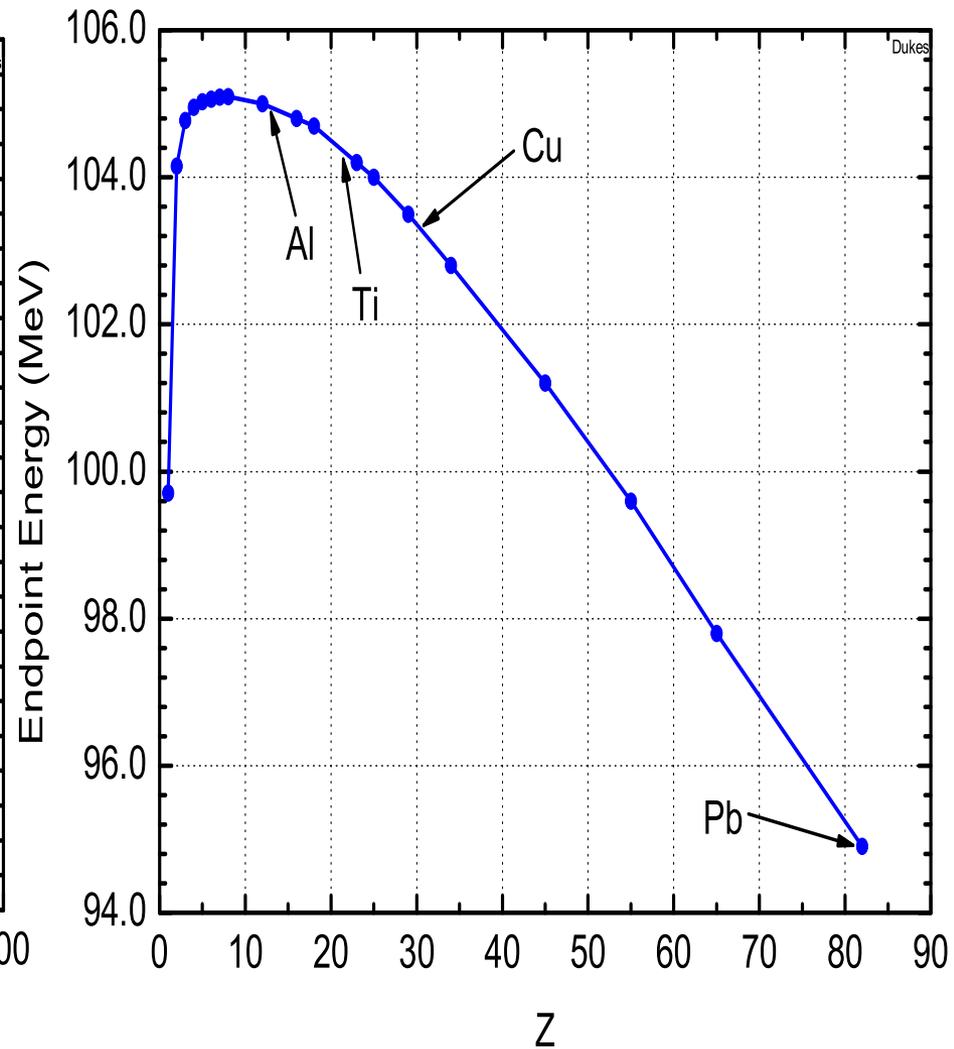
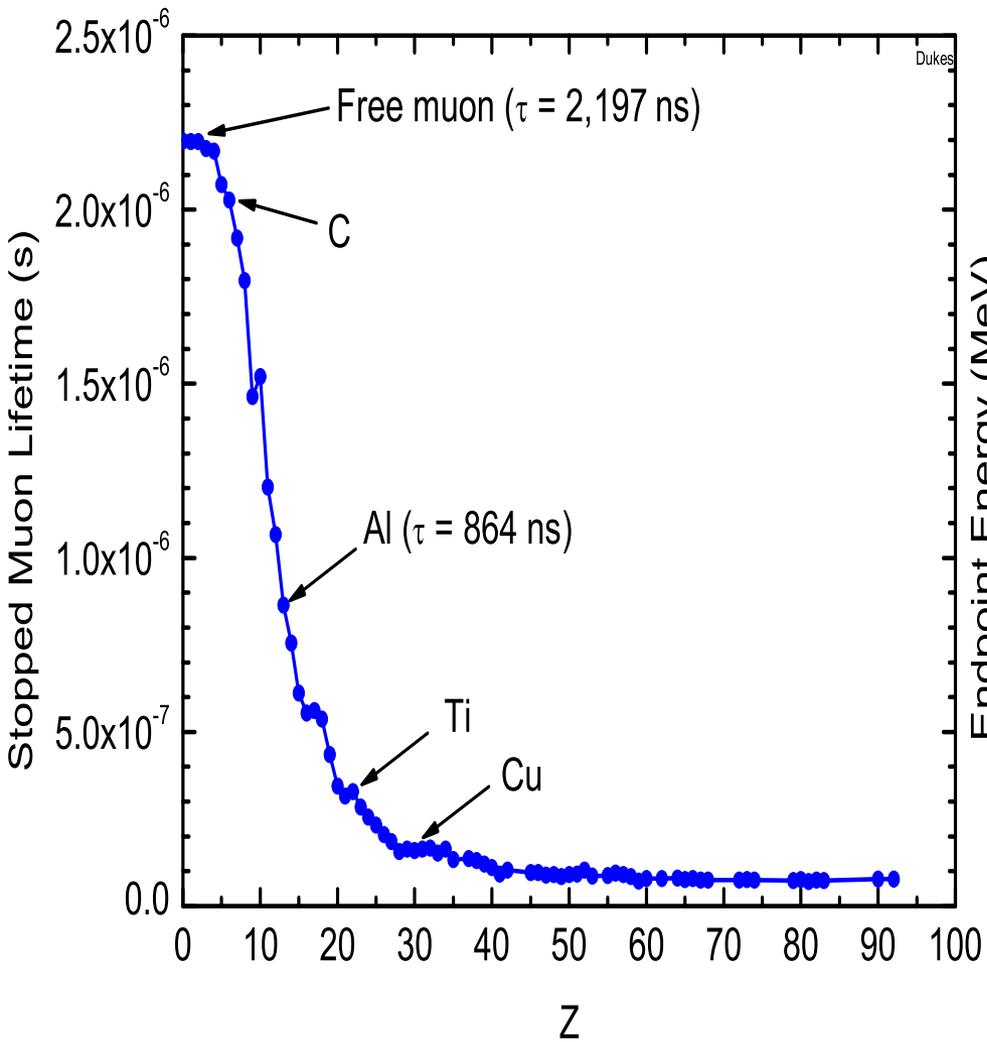
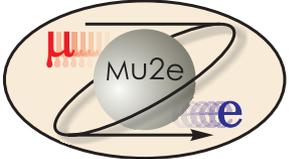
<http://mu2e.fnal.gov/>

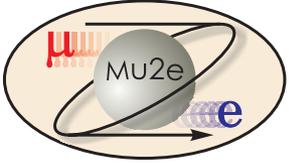
# Thanks for listening!





# Backup Slides



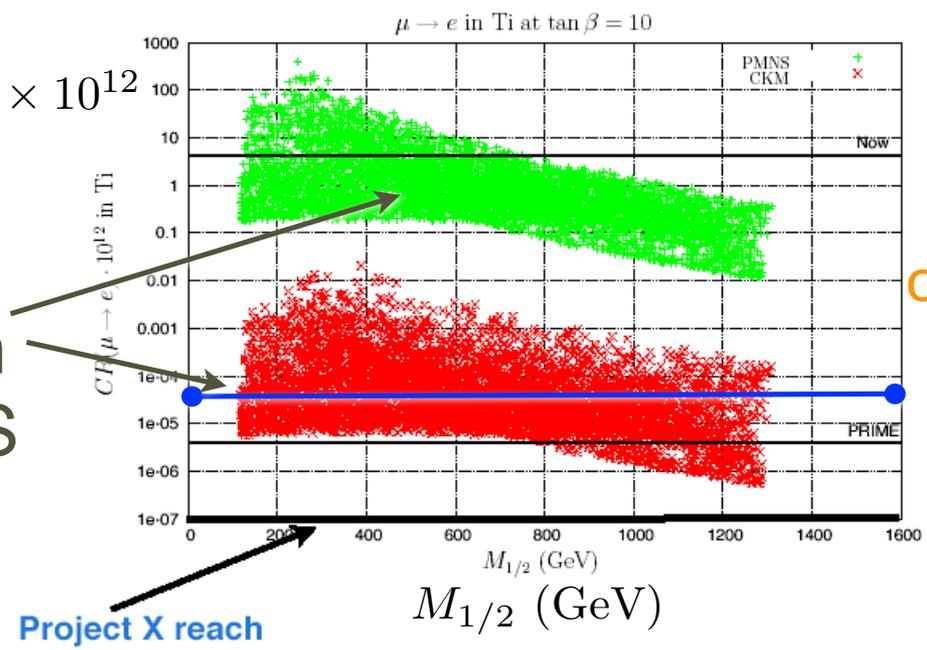


# Mu2e and the LHC

Neutrino-Matrix Like (PMNS) Minimal Flavor Violation(CKM)

$$BR(\mu \rightarrow e) \times 10^{12}$$

measurement can distinguish between PMNS and MFV

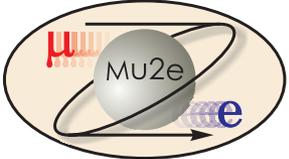


Current  $\mu e$  conversion

Mu2e

L. Calibbi, A. Faccia, A. Masiero, S. Vempati, hep-ph/0605139

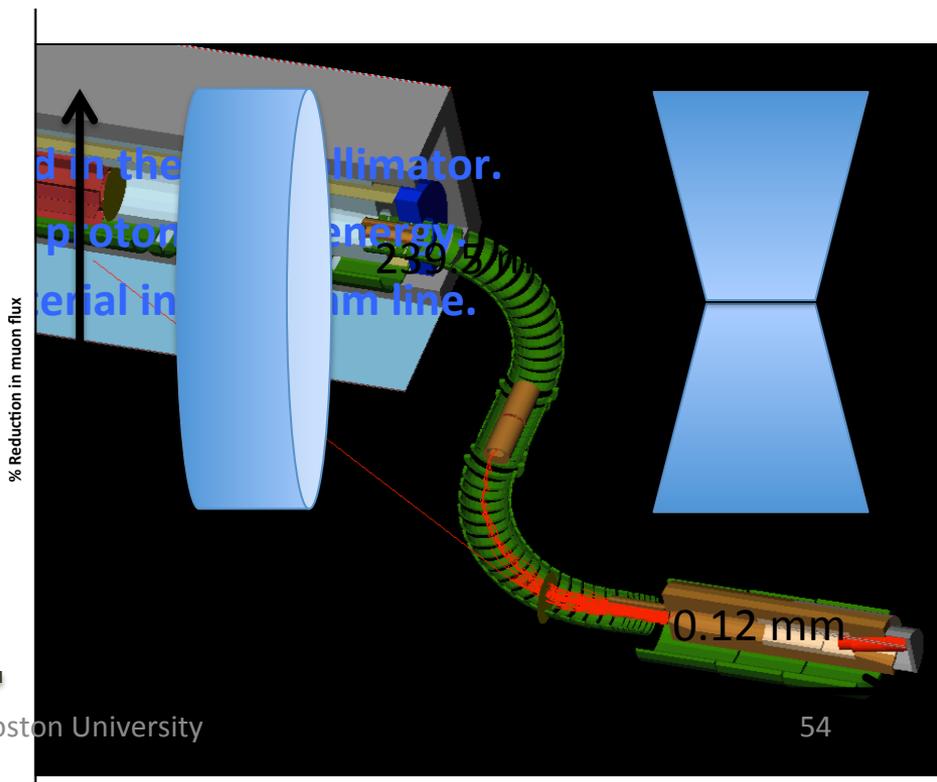
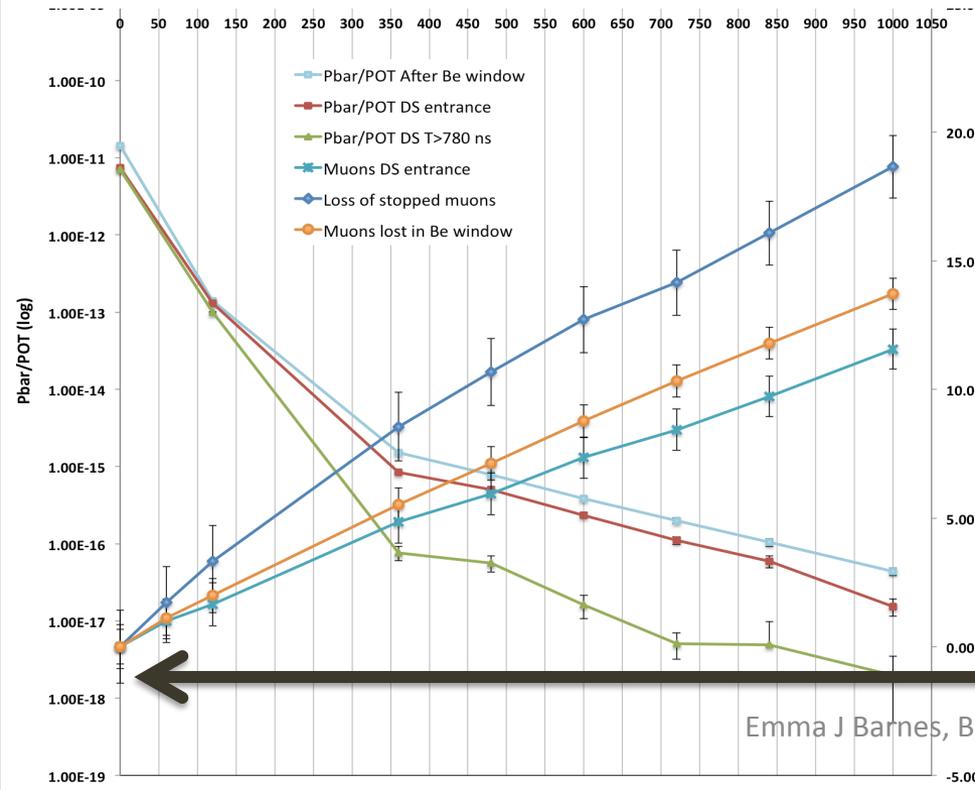
*complementarity between Lepton Flavor Violation (LFV) and LHC experiments*



# Antiprotons

- The 8 GeV K.E proton beam can produce antiprotons in sufficient numbers to be a serious background.
- Spiral slowly through the TS entering the DS.
- They do not decay (well not that we know of!)
- Annihilate on nuclei (such as in the stopping target) releasing a large number of secondaries (energetic  $\pi$ ,  $\gamma$ , n and p.)

**Can reduce the Pbar flux to  $533 \pm 124$  Pbars throughout entire mu2e run**



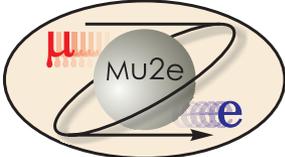
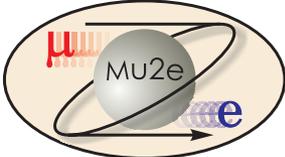


Table 3.1: The table gives the total  $\bar{p}$  induced backgrounds for different incident proton momenta.

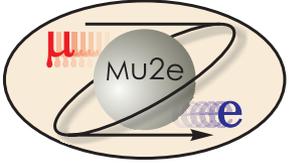
$p$ momentum (GeV/c)	Number of $\bar{p}/p$ produced	Number of $\bar{p}/p$ entering transport	Background events
5	$3.9 \times 10^{-10}$	$7.4 \times 10^{-15}$	$7 \times 10^{-7}$
6	$5.3 \times 10^{-8}$	$8.0 \times 10^{-13}$	$8 \times 10^{-5}$
7	$1.4 \times 10^{-6}$	$1.2 \times 10^{-11}$	$1.2 \times 10^{-3}$
8	$8.5 \times 10^{-6}$	$6.8 \times 10^{-11}$	$7 \times 10^{-3}$



# Beam requirements

<b>Parameter</b>	<b>Design Value</b>	<b>Requirement</b>	<b>Unit</b>
Booster synchrotron repetition rate	15	$> 10.5^2$	Hz
Booster synchrotron beam intensity	$4.0 \times 10^{12}$	$4.0 \times 10^{12}$	Protons/batch
Time between beam pulses	1685	$> 864$	nsec
Out of time extinction factor	$10^{-10}$	$\leq 10^{-10}$	
Pulse full width	$\pm 100$	$\leq \pm 130$	nsec
Pulse rms width	40	$\leq 50$	nsec
Duration of spill	54	$> 20$	msec
Beam Power on Target	8	-----	kW
Average proton intensity per pulse	31	$< 50$	Mp/pulse
Pulse to Pulse intensity variation	50	$< 50$	%
Minimum Target rms spot size <sup>3</sup>	1	0.5	mm
Maximum Target rms spot size <sup>3</sup>	1	2.0	mm
Target rms beam divergence	0.5	$< 20$	mrad

Table 5.1. Summary of the Mu2e Proton Beam Requirements



# Accelerator timeline

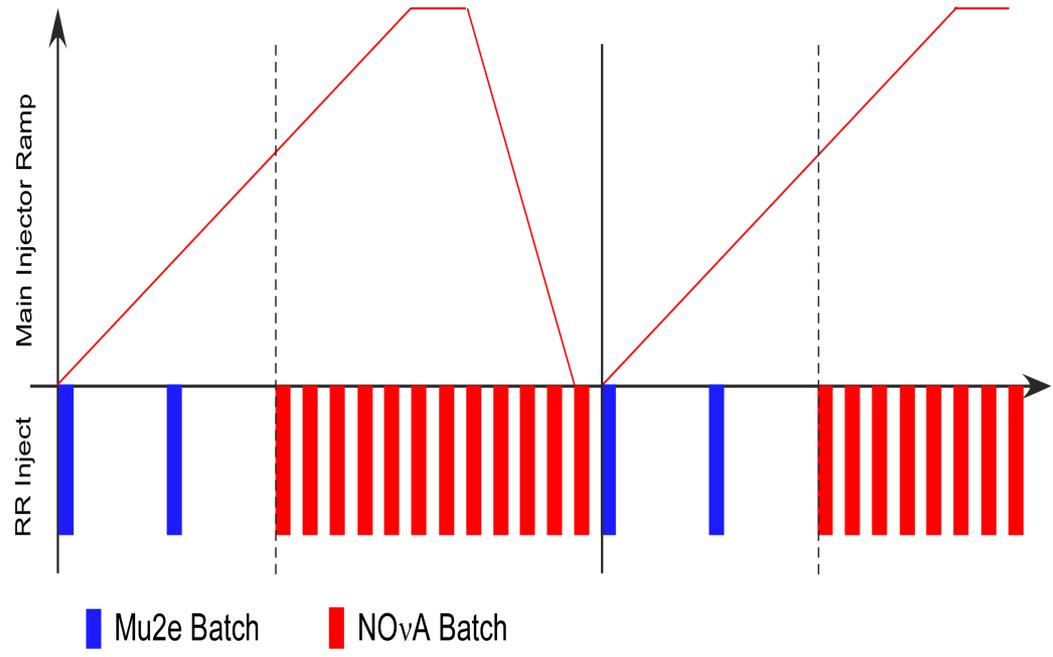


Figure 5.2. The accelerator timeline is shared between Mu2e and NOvA. The blue and red bars represent Mu2e and NOvA proton batch injections respectively. Mu2e Recycler Ring beam manipulations occur in the first eight 15 Hz ticks. NOvA proton batches are slip-stacked during the remaining twelve 15 Hz ticks. The total length of a cycle is 20 ticks = 1.333 sec.

